

Latest Super-K Atmospheric Neutrino Oscillation Analysis

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Now2012, Conca Specchiulla (Otranto Italy)

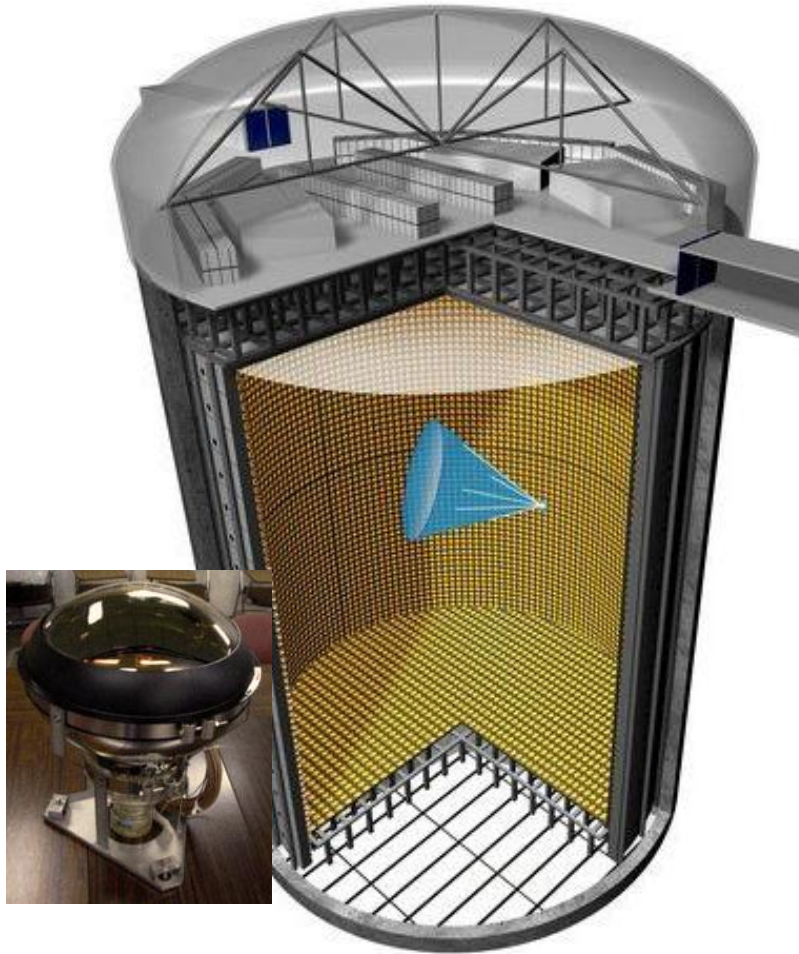
Introduction

- ❑ Atmospheric neutrinos are a good tool for studying L/E-style oscillations in a broad sense
 - ❑ They can also provide information on sub-leading effects

- ❑ Now that θ_{13} is known to be large the goals of atmospheric neutrino measurements become more focused:
 - ❑ Can we determine **Mass Hierarchy** ?
 - ❑ Can we determine the θ_{23} **octant** ?
 - ❑ Can we constrain the allowed values of δ_{cp} ?

- ❑ Several atmospheric neutrino analyses :
 - ❑ **Three-flavor oscillation analysis** → **Today**
 - ❑ L/E Analysis
 - ❑ Two-flavor Analysis
 - ❑ Tau Appearance analysis

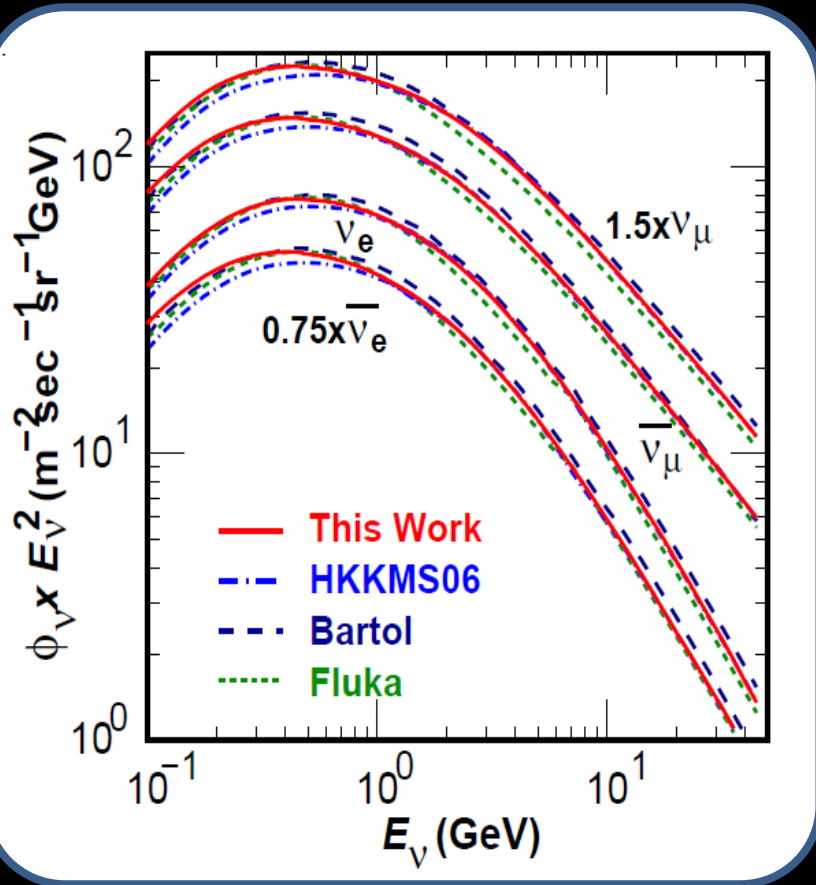
Super-Kamiokande : Introduction



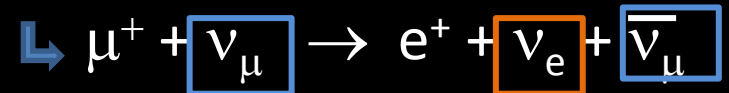
- ❑ SK discovered atmospheric neutrino oscillations exploiting the disappearance of upward-going muon events
- ❑ Event samples are classified by
 - ❑ Number of rings (**Single-** or **Multiple-**)
 - ❑ PID of out-going lepton (**e-** or **μ -like**)
 - ❑ Event energy is **fully** or **partially** contained
 - ❑ Upward-going muons
- ❑ Today: **3903 days of data**
 - ❑ ~34,000 Events
- ❑ Currently **statistics limited**

- ❑ 22.5 kton FV Volume
- ❑ Ring Imaging Water Cherenkov detector
- ❑ 11,146 20" Phototubes
- ❑ Data taken over four periods since 1996

Atmospheric Neutrino Generation

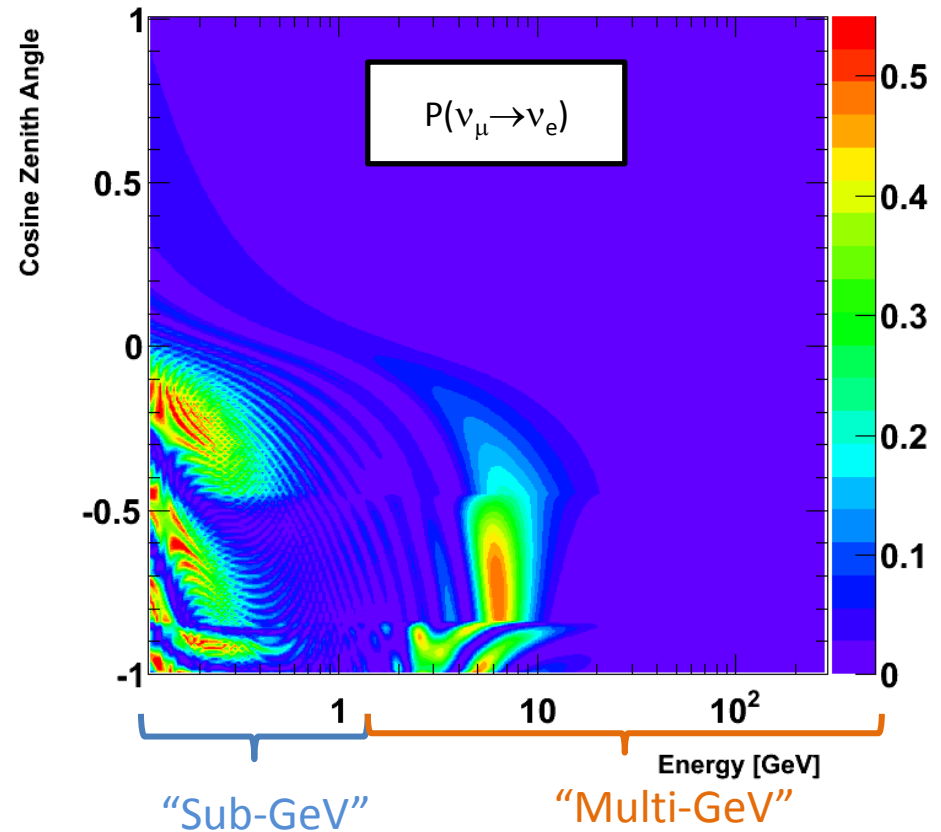
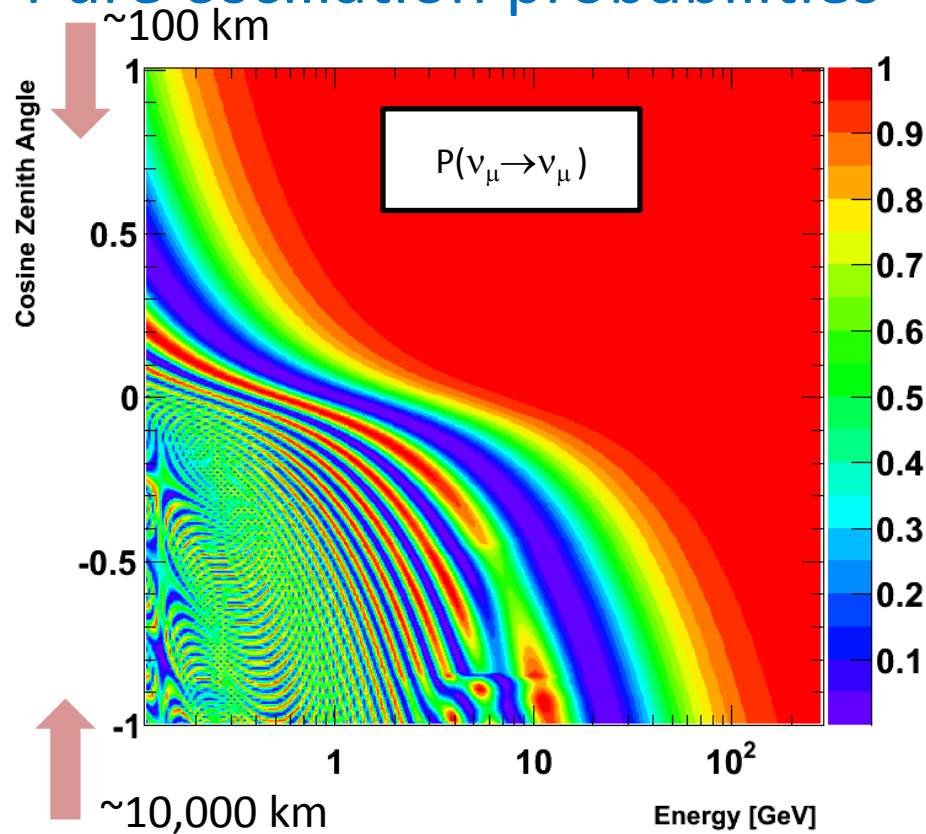


- Cosmic rays strike air nuclei and the decay of the outgoing hadrons gives neutrinos



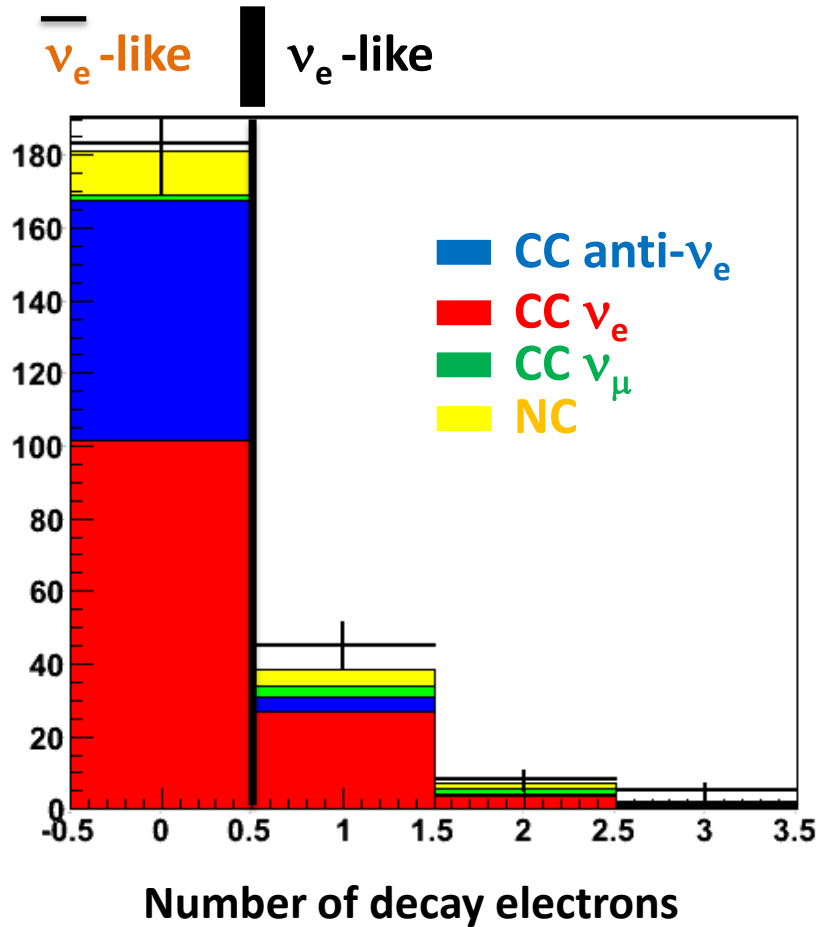
- Isotropic about the Earth
 - Path length to the detector spans 10 – 10,000 km
- Both neutrinos and antineutrinos
 - about 30% of the final samples are antineutrinos
- Spans many decades in energy $\sim 100 \text{ MeV} - 100 \text{ TeV}^+$
- Excellent tool for broad studies of neutrino oscillations
 - Access to sub-leading effects with high statistics

Pure oscillation probabilities



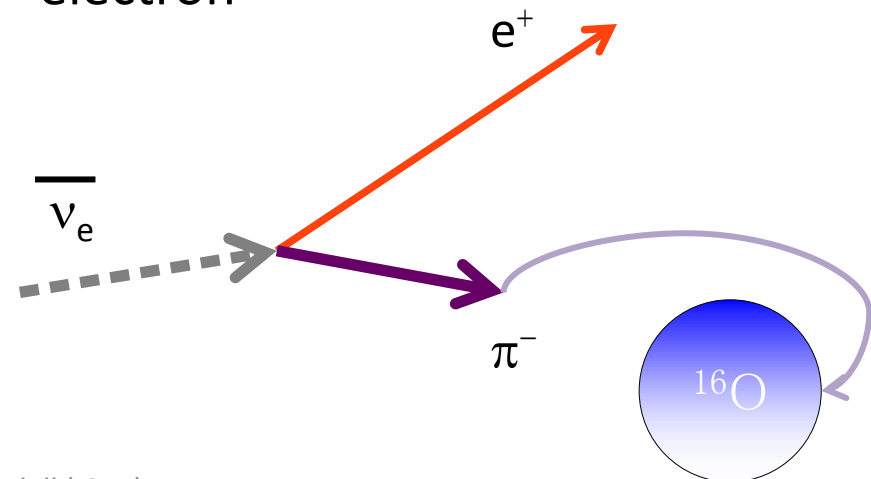
- ❑ With large θ_{13} resonant enhancement of the $P(\nu_\mu \rightarrow \nu_e)$ probability occurs in matter
- ❑ Resonance occurs only for (anti-)neutrinos under the Normal (Inverted) Hierarchy
 - ❑ Size of resonance is a function of θ_{23} octant, θ_{13} , and δ_{cp}
 - ❑ "Signal" present in both μ -like and e -like samples
 - ❑ Ideally we would like **separate neutrino and antineutrino** samples

Sample Selection : Multi-GeV Single-Ring anti- ν_e and ν_e -like



(Multi-ring events are in general more complicated separation is done using a likelihood)

- Separate neutrinos from anti-neutrinos in the single-ring sample using the number of observed decay electrons
- The outgoing π^- from an anti-neutrino CC-1 π event can be absorbed on a ^{16}O nuclei before it decays. The lack of an outgoing muon means there is no possibility of a subsequent Michel (decay) electron



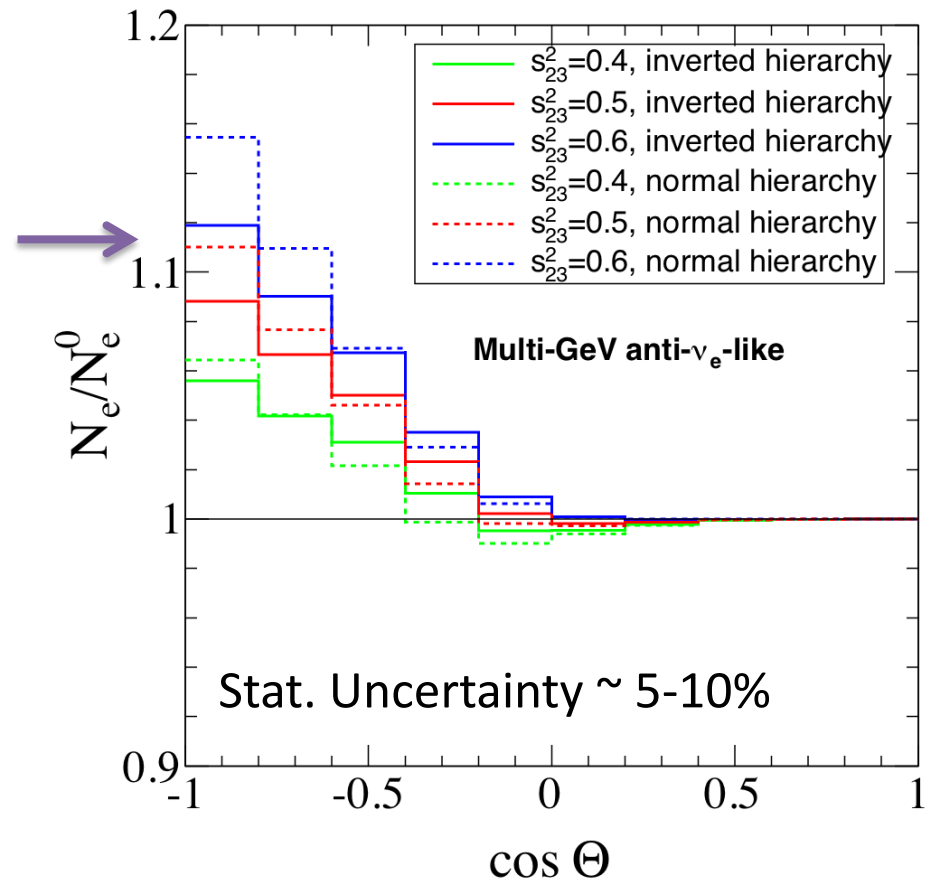
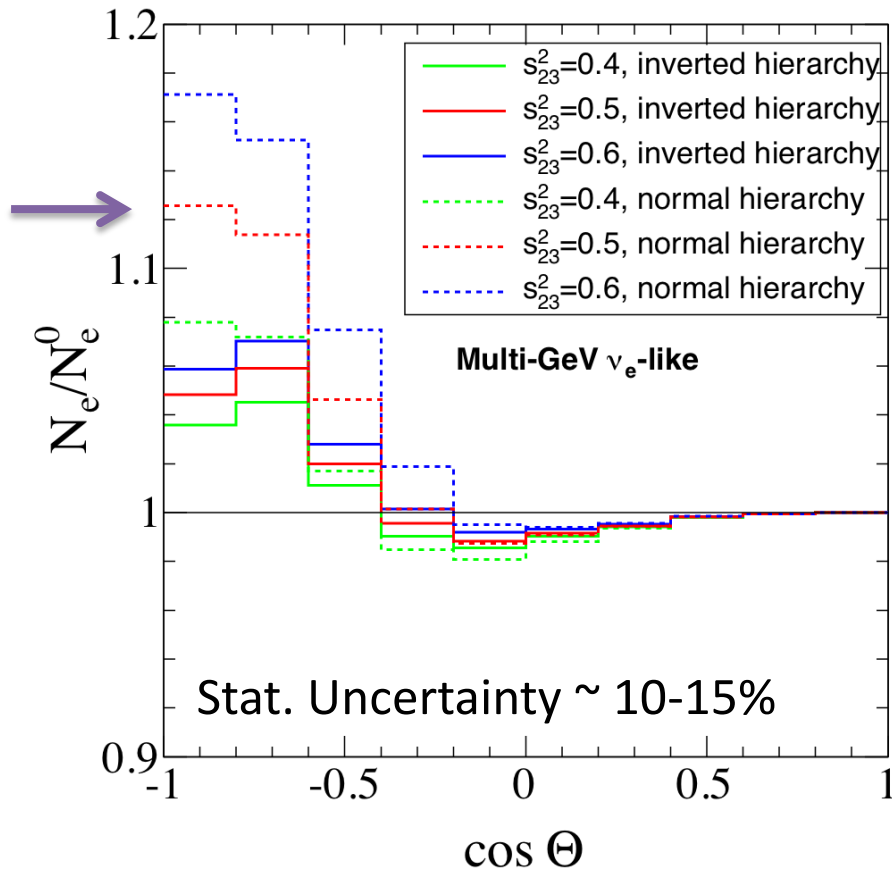
Sample Composition

Composition (%)		CC ν_e	CC anti- ν_e	CC $\nu_\mu + \text{anti-}\nu_\mu$	NC
ν_e like	1R	60.2	10.6	13.5	14.8
	MR	57.5	17.4	10.7	13.7
Anti- ν_e like	1R	55.7	36.6	1.1	6.4
	MR	51.9	20.7	8.2	19.7

Composition (%)		CC ν_e	CC anti- ν_e	CC $\nu_\mu + \text{anti-}\nu_\mu$	NC
ν_μ like	1R	0.2	0.08	98.8	0.2
	MR	2.5	0.3	91.7	4.4

- Generally the background component of the e-like signal samples is **~20-30%**
- Muon-like samples on the other hand tend have high-purity and reasonable sensitivity to small effects

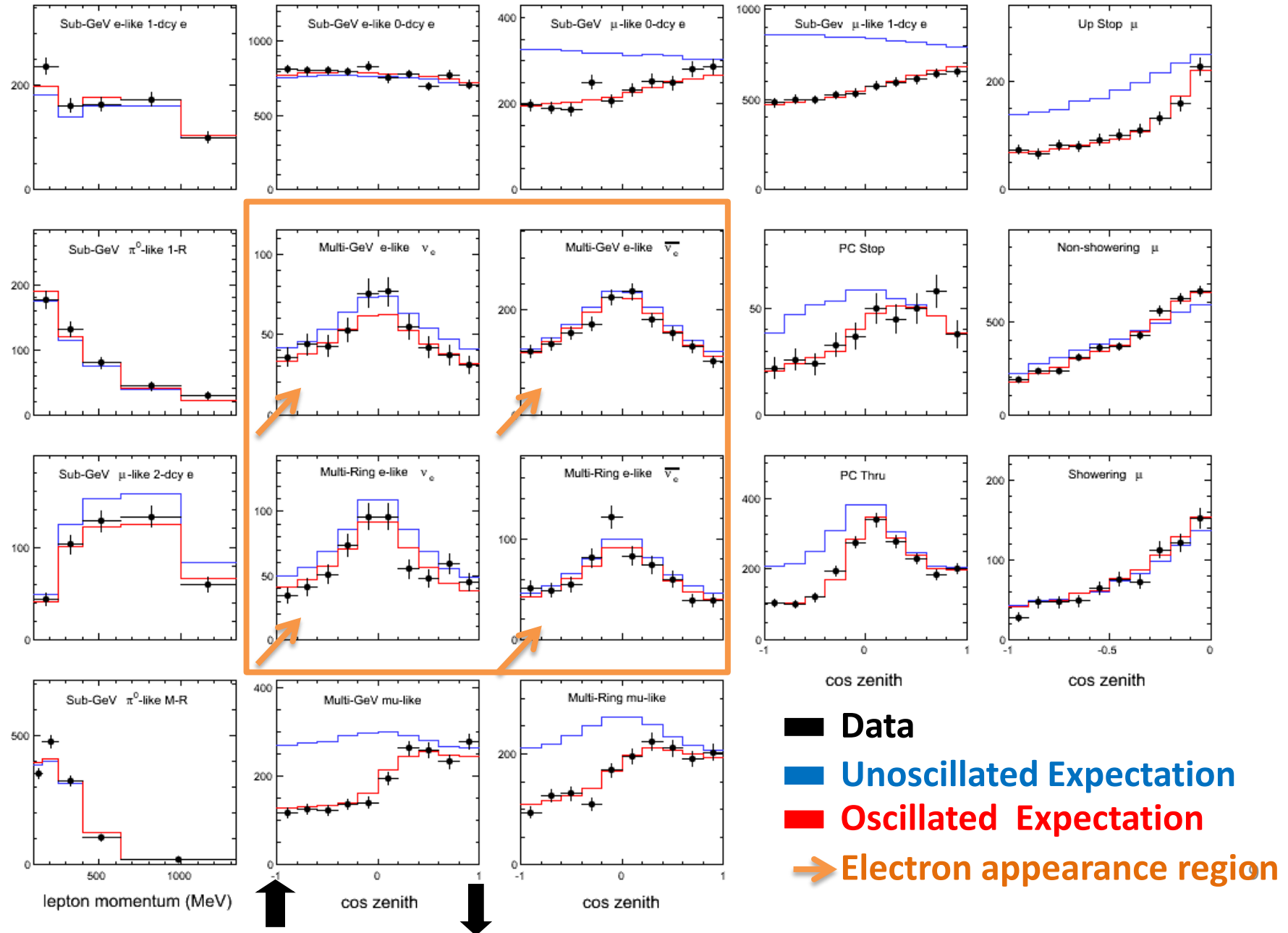
Expected Effects : Single-Ring e-like and anti-e-like samples



→ Equivalent MC

- ❑ Effect of the θ_{23} octant can be large.
- ❑ That from δ_{cp} is an additional $\pm 2\%$ on each line
- ❑ Statistical uncertainty remains cumbersome

Super-K I+II+III+IV Data – Make Use of All Potential Signals



About Systematics

- ❑ Super-K analysis considers 151 sources of systematic uncertainty from the usual cadre of errors
 - ❑ Flux, cross-section, detector performance
- ❑ This is a partial listing of things relevant to three-flavor issues

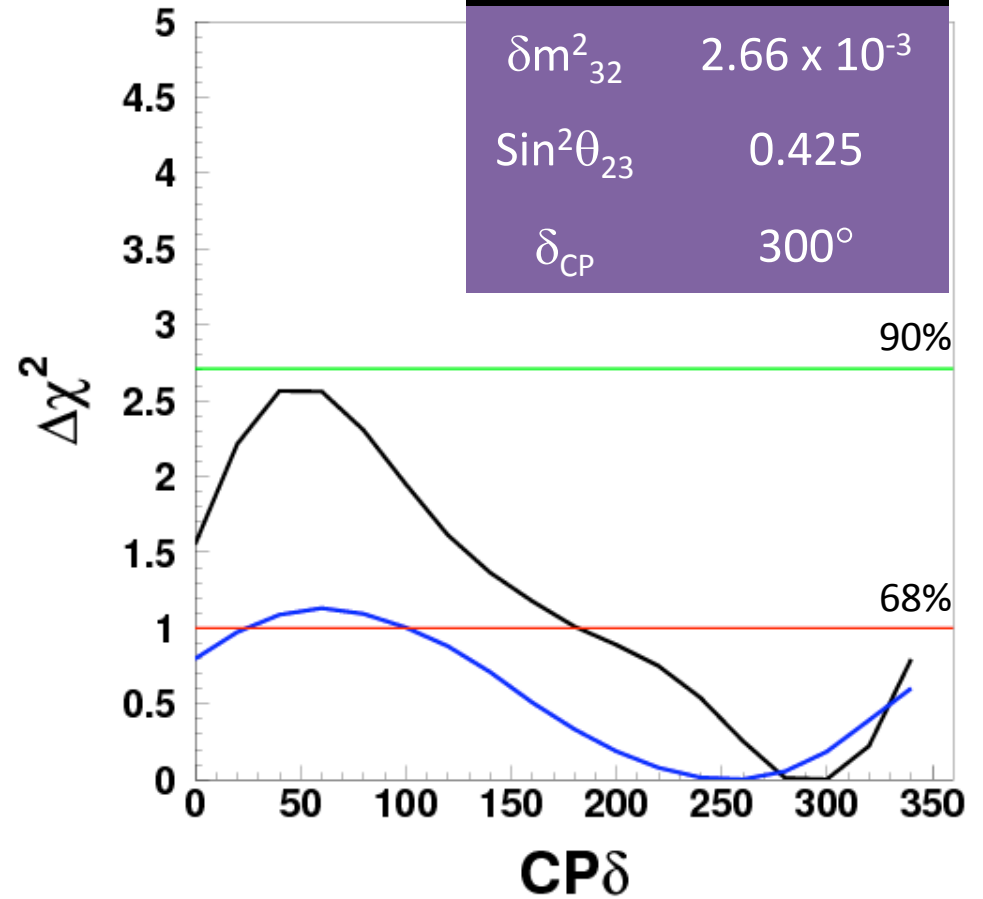
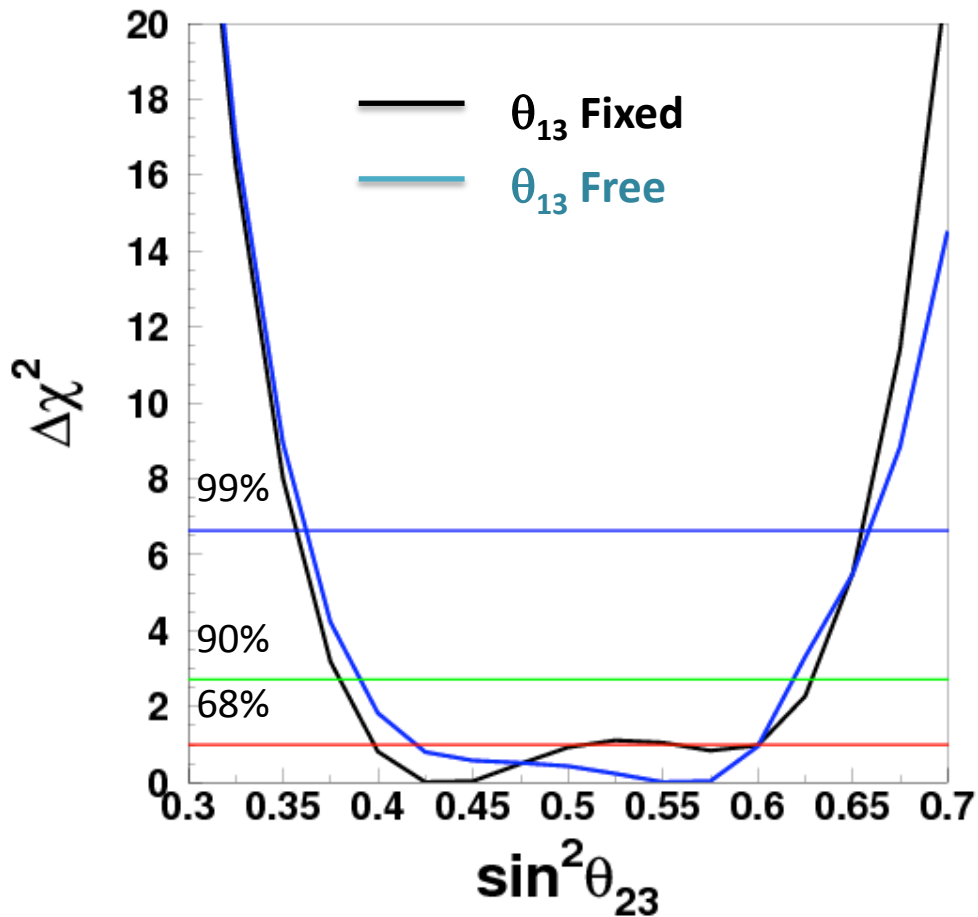
Error Source	Uncertainty
ν_e vs. anti- ν_e sample selection	7%
Charged-Neutral Pion Production	40%
Tau Production Cross section	25%
DIS Cross Section	5-10%
NC / CC Ratio	20%
Single-Pion Production	20%
Flux Normalization above 1 GeV	7%
Flux Ratio ν to $\bar{\nu}$ above 1 GeV	5-8%

← Tau events look similar to e-like events and they are upward-going

Results

- + Simultaneous fitting performed on Δm^2_{32} , θ_{23} , δ_{cp} , θ_{13}
- + Both hierarchies considered
- + Two Analyses : θ_{13} Free / Fixed to reactor value

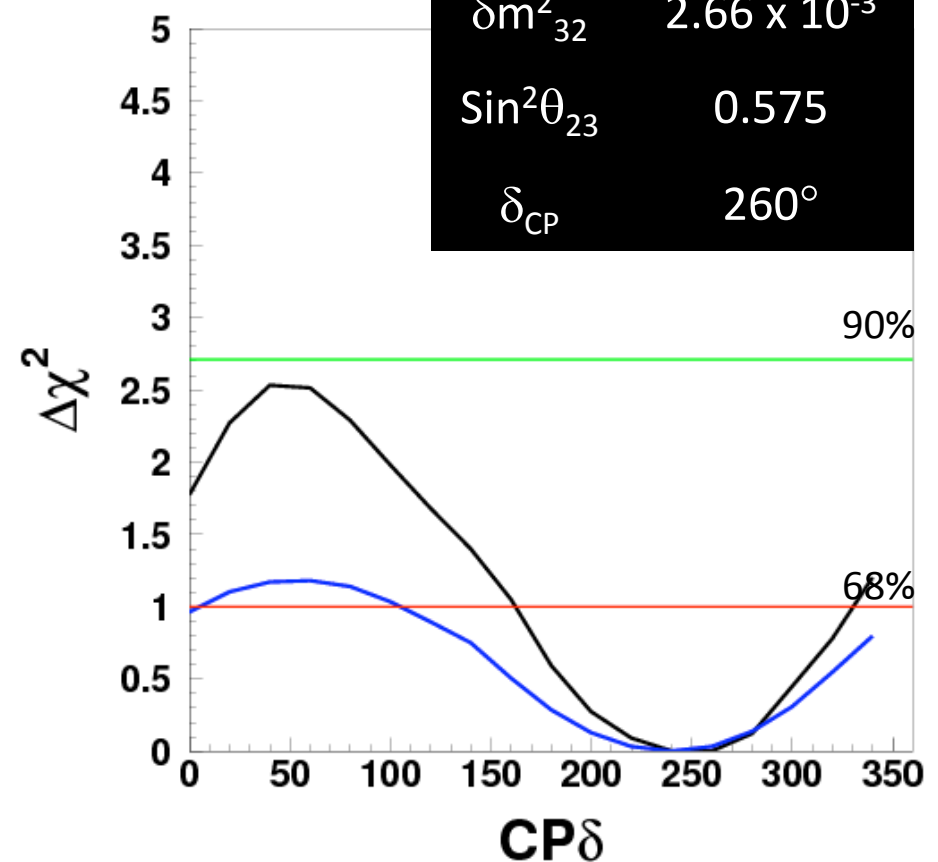
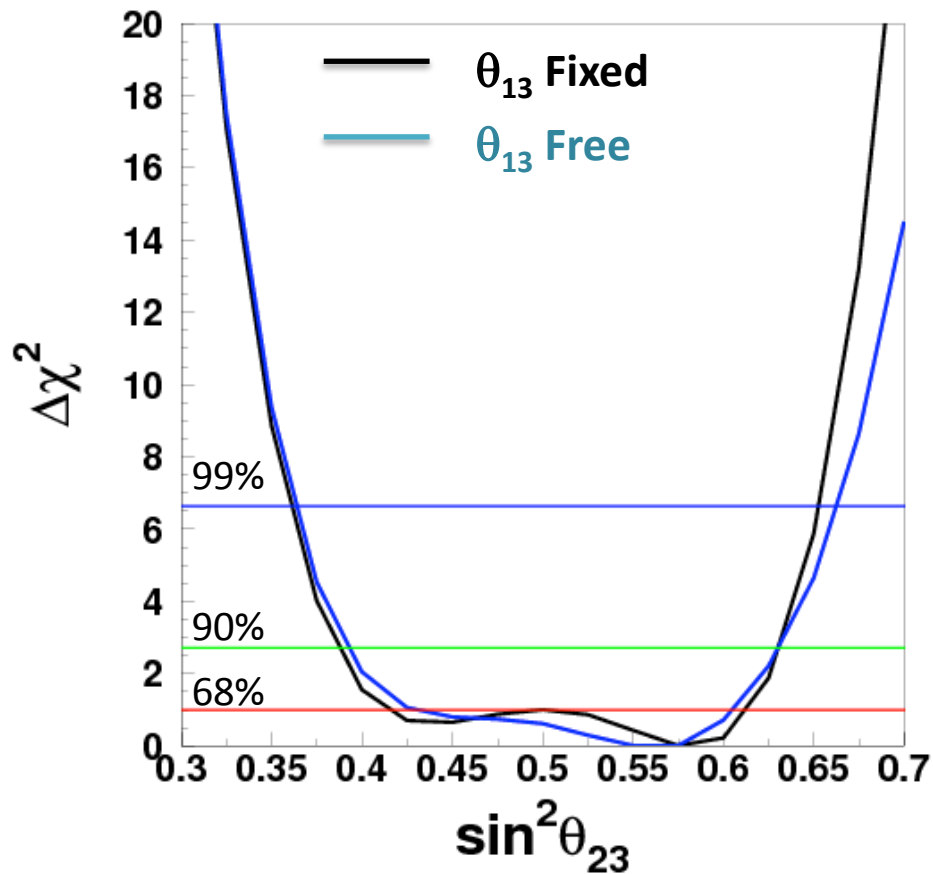
Normal Hierarchy Results



NH	Best -Fixed
δm^2_{32}	2.66×10^{-3}
$\sin^2\theta_{23}$	0.425
δ_{CP}	300°

- ❑ Fixing the value of θ_{13} to the reactor experiments value moves the preferred octant from the second to the first (only a very weak preference)
- ❑ Weak constraint on δ_{cp} , all values allowed at 90% C.L. in both fits

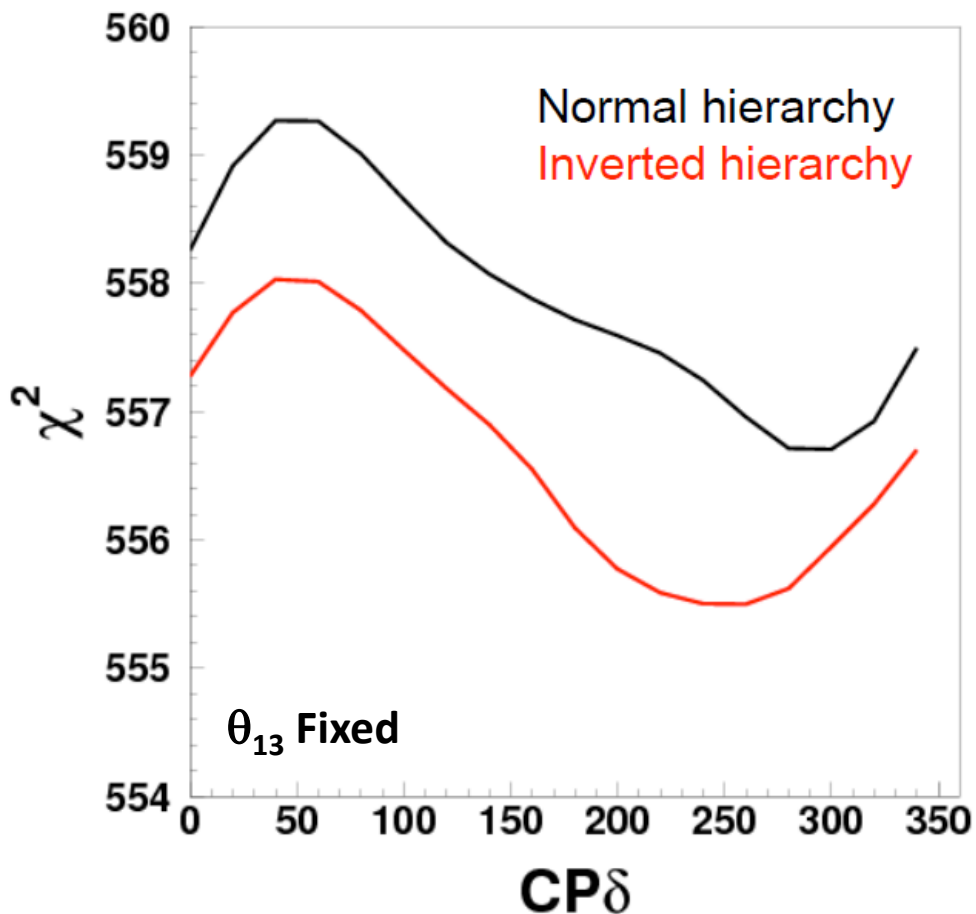
Inverted Hierarchy Results



IH	Best -Fixed
δm^2_{32}	2.66×10^{-3}
$\sin^2\theta_{23}$	0.575
δ_{CP}	260°

- Both fits find minima in the 2nd octant
- Weak constraint on δ_{cp} , all values allowed at 90% C.L. in both fits

Hierarchy Comparison



Hierarchy

Fit Result

Normal

556.7 / 477 dof

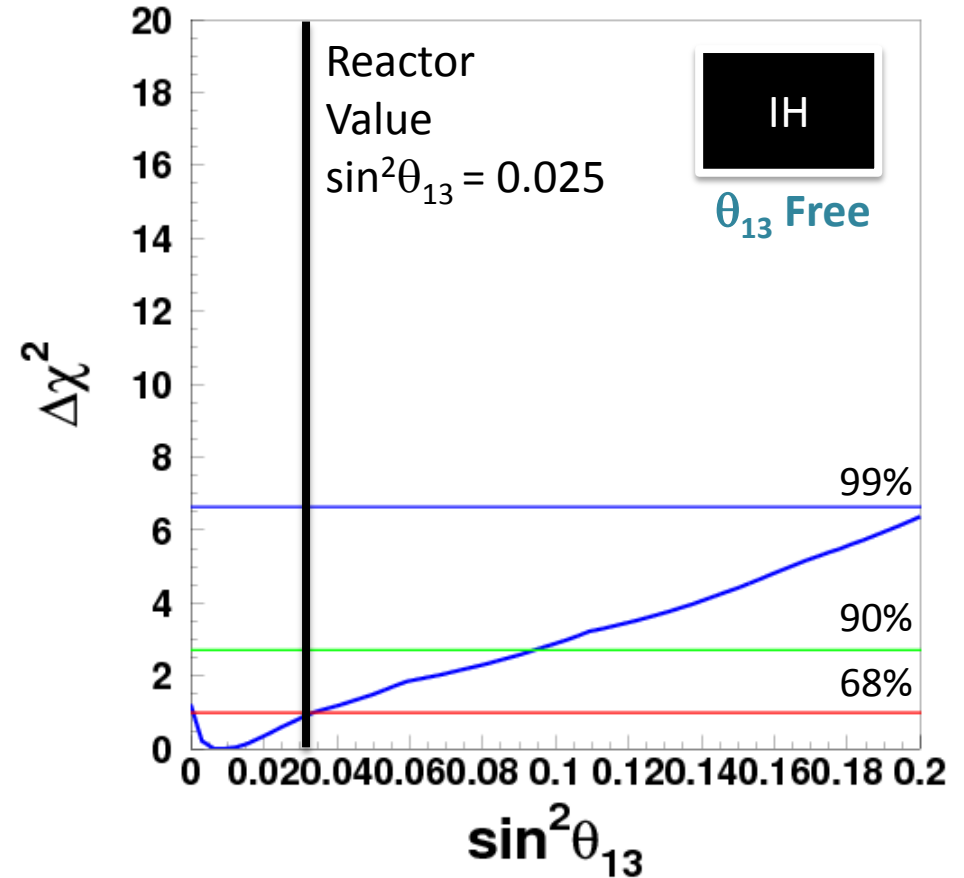
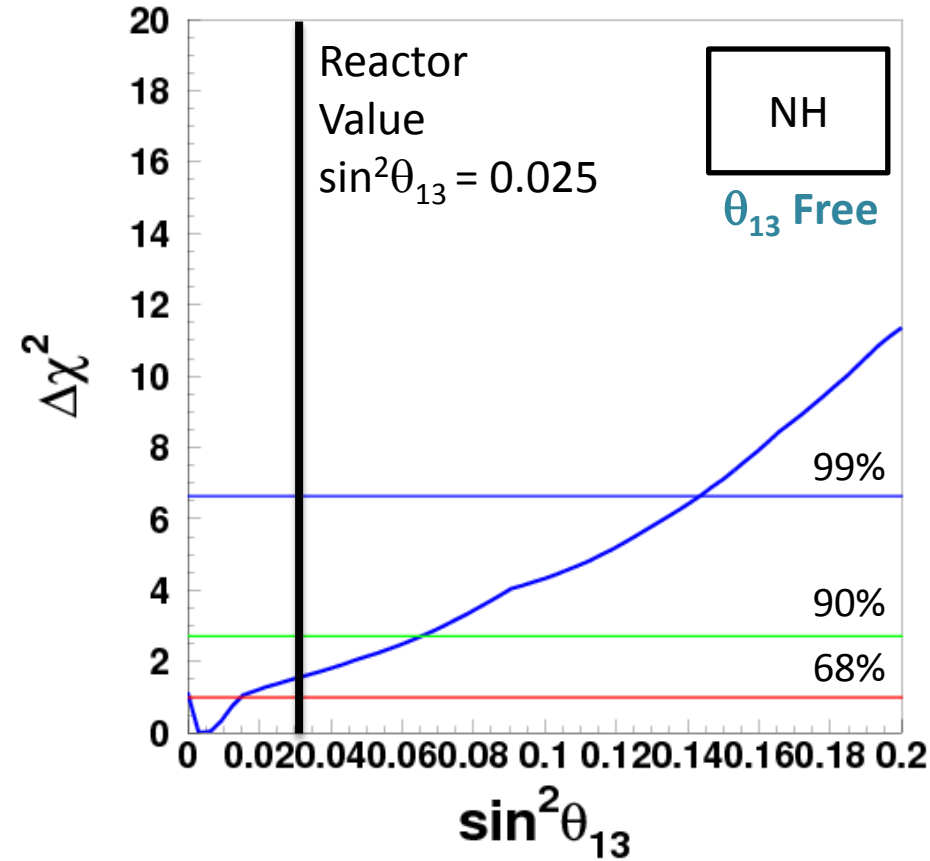
Inverted

555.5 / 477 dof

$$\chi^2_{\min} (\text{NH}) - \chi^2_{\min} (\text{IH}) = 1.2$$

- Expected sensitivity to the inverted hierarchy is $\Delta\chi^2 = 0.9$
- Data show a 1σ preference for the inverted hierarchy, not significant

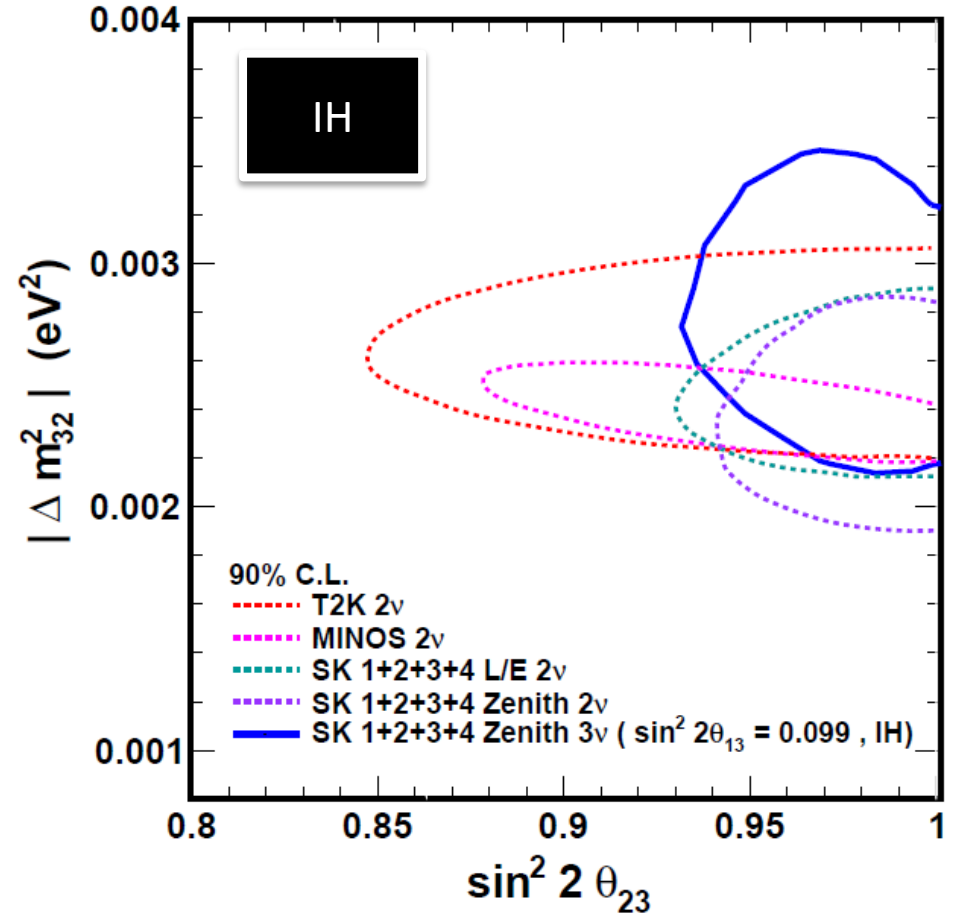
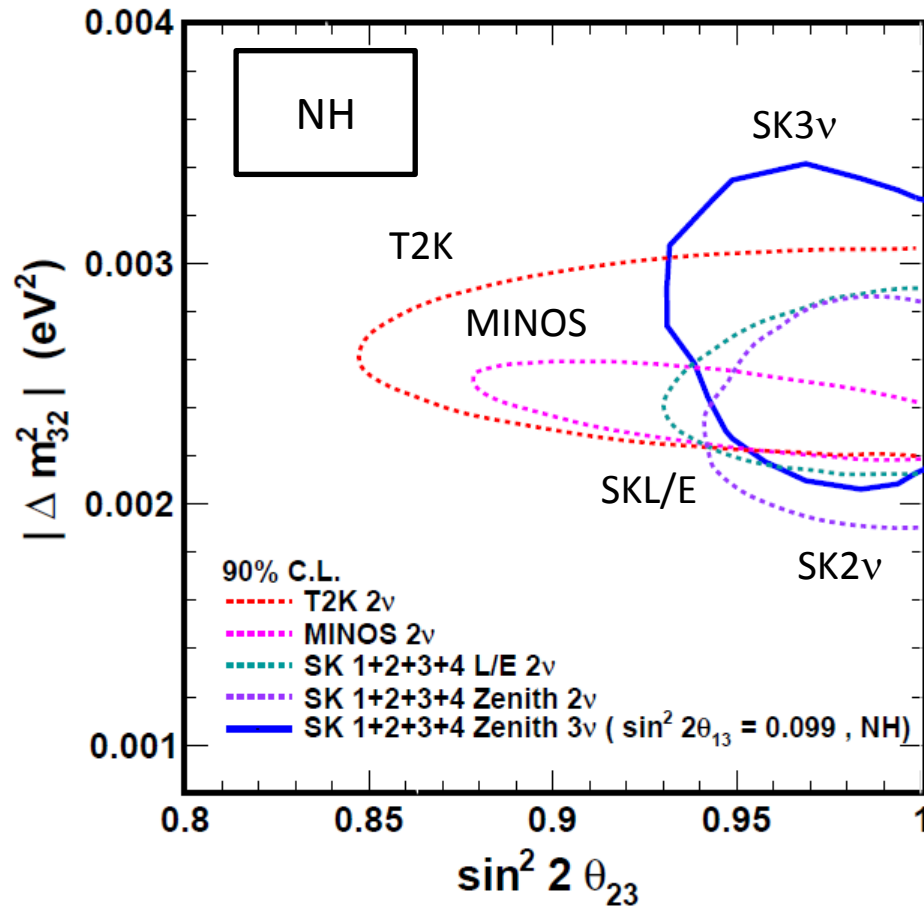
θ_{13} Measurement With Atmospheric Neutrinos



Hierarchy	Best $\sin^2\theta_{13}$	$\sin^2\theta_{13}$ - 90% CL	χ^2
Normal	0.003	0-0.066	555.3 / 476
Inverted	0.006	0-0.094	554.9 / 476

Still a slight preference for the inverted hierarchy

Disappearance Parameters Post Neutrino 2012



- Atmospheric neutrinos still provide the strongest constraint on θ_{23}
- Long-baseline experiments have measured Δm^2 more precisely
- Atmospheric neutrino measurement is consistent

Summary

- ❑ Super-K oscillation analysis has been updated to include constraints from global information on θ_{13}
- ❑ No decisive indication of the mass hierarchy, θ_{23} octant, or value of δ_{cp} yet ...

NH	Best fit	90% C.L.
Δm^2_{32}	2.66×10^{-3}	$[2.06 - 3.04] \times 10^{-3} \text{ eV}^2$
$\sin^2 \theta_{23}$	0.425	0.391 - 0.619
δCP	300°	All allowed at 90% C.L.

IH	Best fit	90% C.L.
Δm^2_{32}	2.66×10^{-3}	$[2.14 - 3.04] \times 10^{-3} \text{ eV}^2$
$\sin^2 \theta_{23}$	0.575	0.393 - 0.630
δCP	240°	All allowed at 90% C.L.

❑ The Future

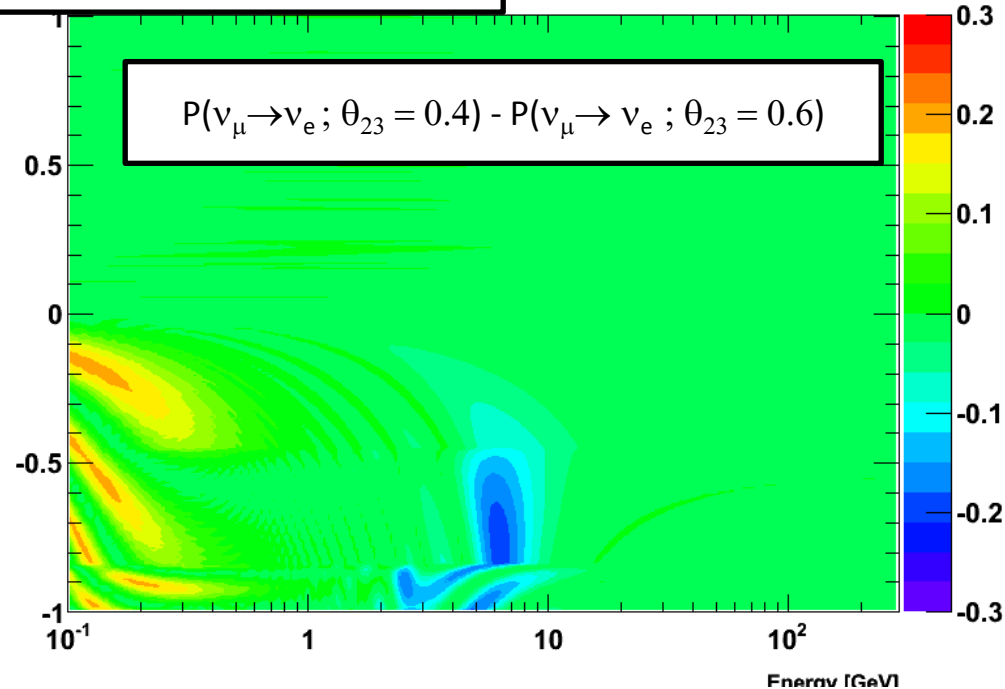
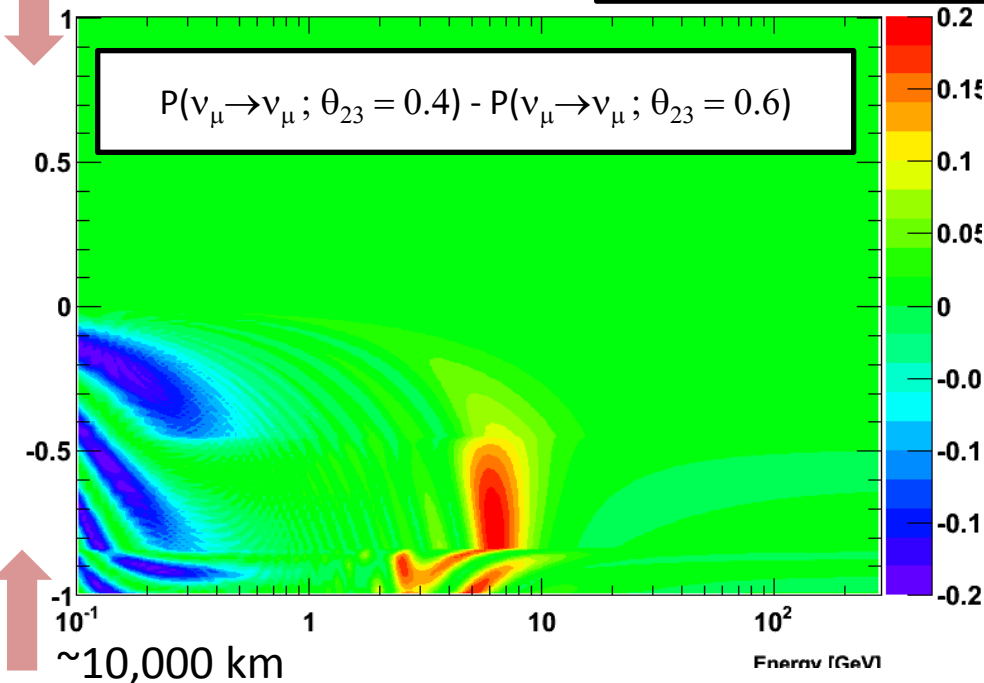
- ❑ Still statistics limited
- ❑ Constraints from external data may help focus the sensitivity of atmospheric neutrinos
 - ❑ Plans for an analysis incorporating global data underway

Supplements

Oscillation probability difference between the θ_{23} octants

$$P(1^{\text{st}} \text{ Octant}) - P(2^{\text{nd}} \text{ Octant})$$

~100 km



□ Matter effect gives improved sensitivity

□ Mass hierarchy

→ Asymmetry between neutrinos and antineutrinos

□ size of θ_{13} and δ_{cp}

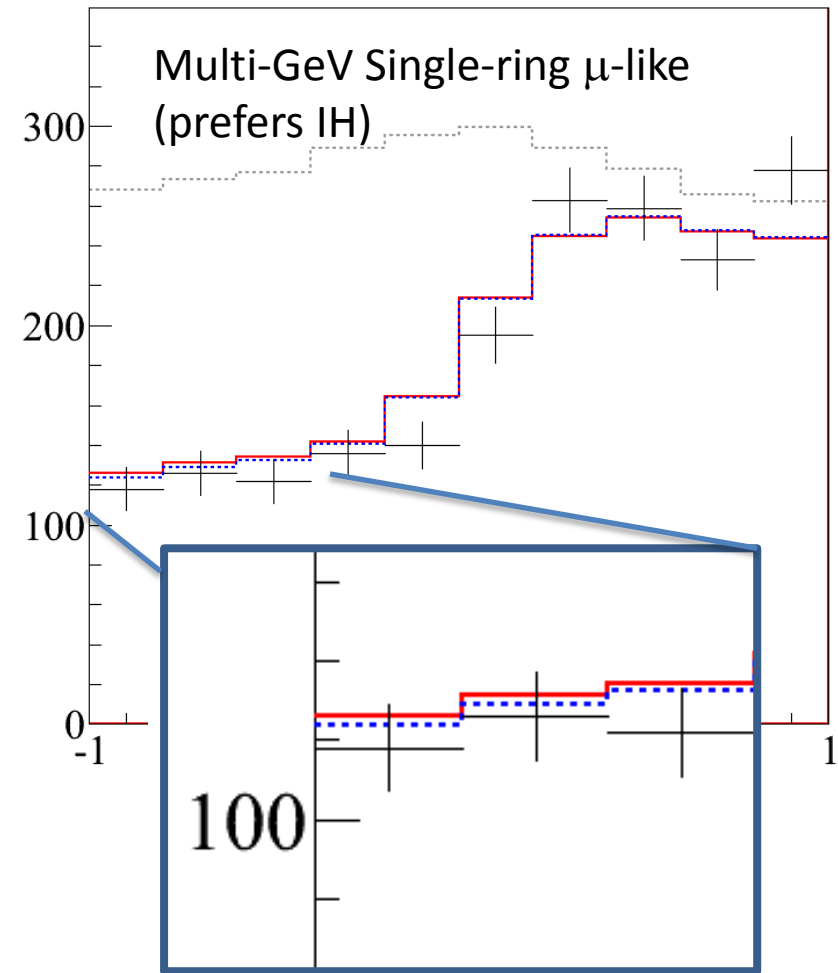
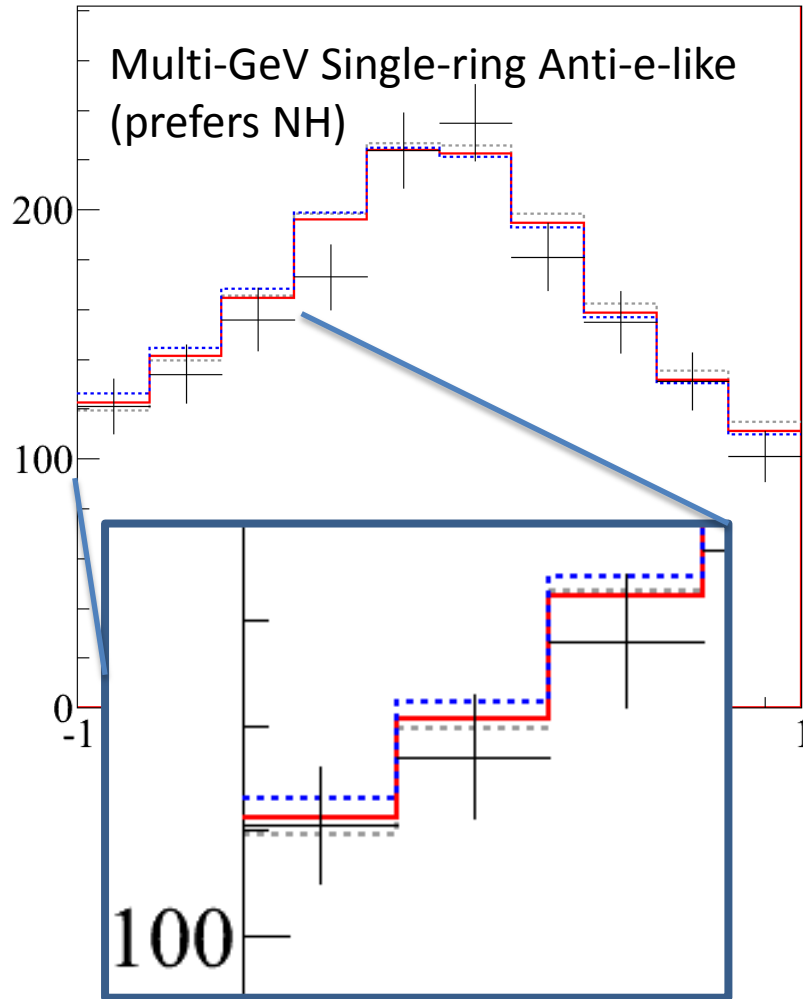
→ Magnitude of resonance effect

□ Octant of θ_{23}

→ Appearance and disappearance interplay

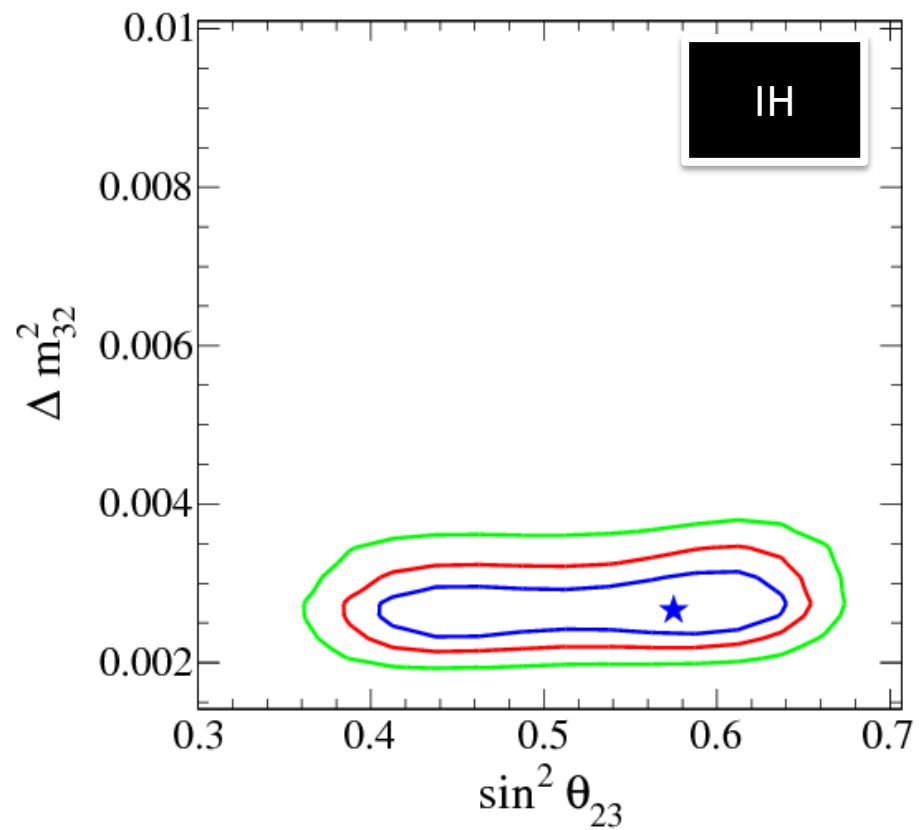
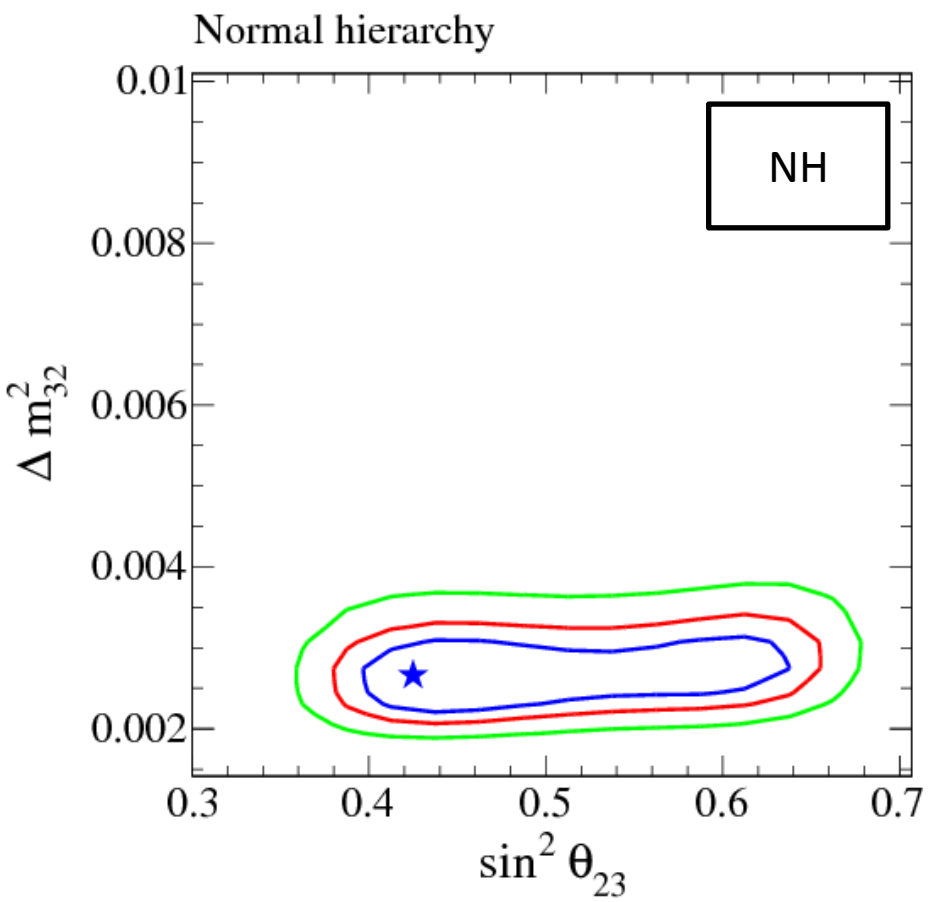
(Trends are Independent of Hierarchy)

Zenith angle distributions of the Data

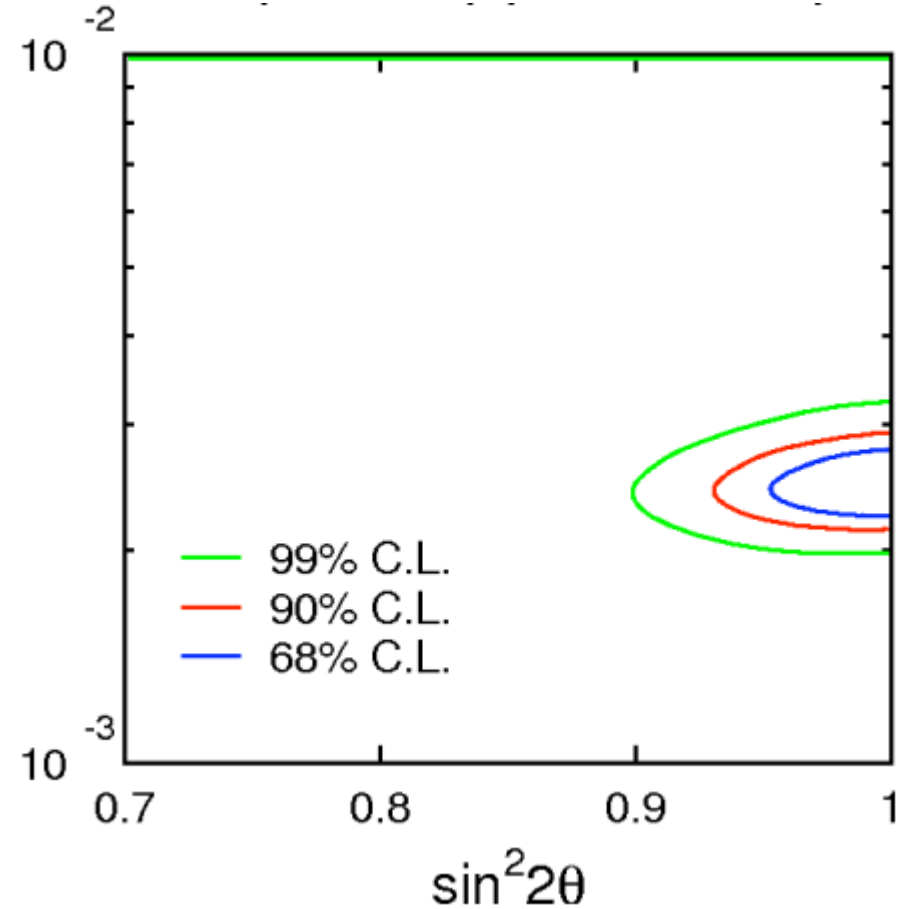
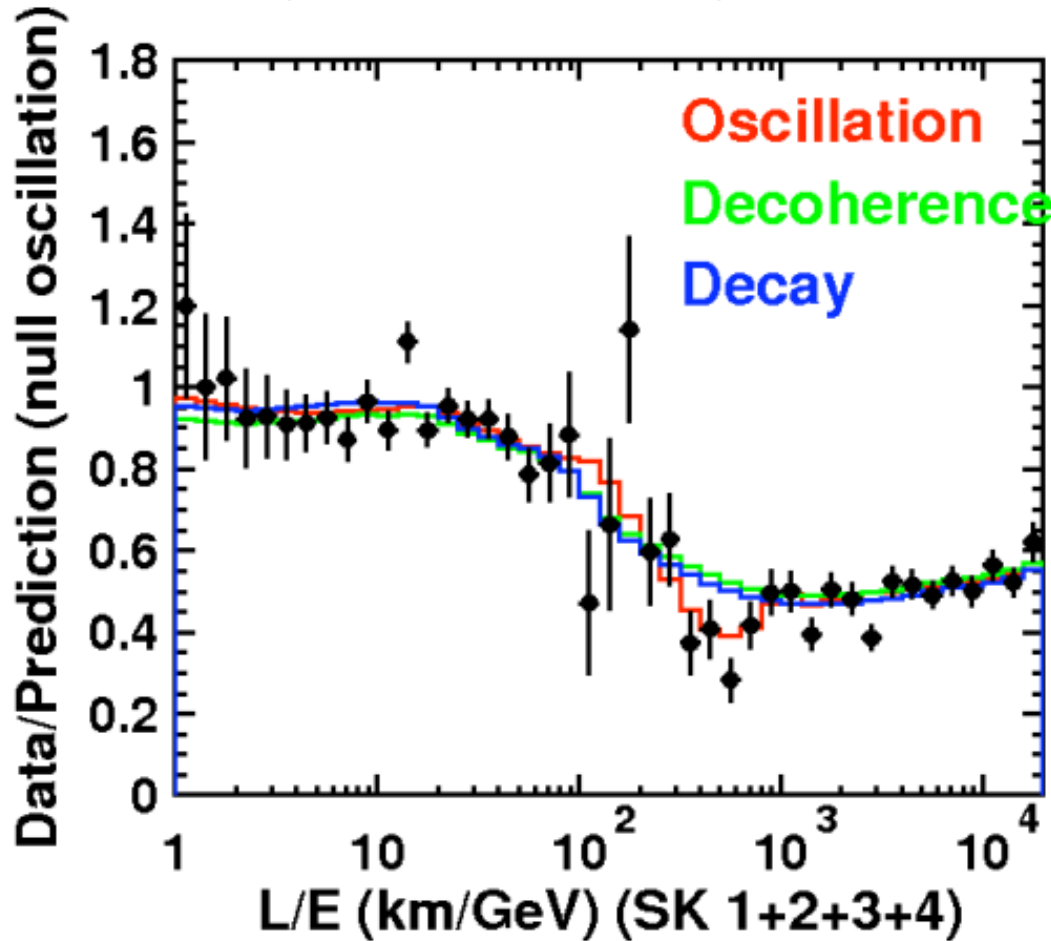


- ❑ So far no consistent story has emerged from the data
- ❑ Taking more data is a necessity !

— NH Fit
- - - IH Fit



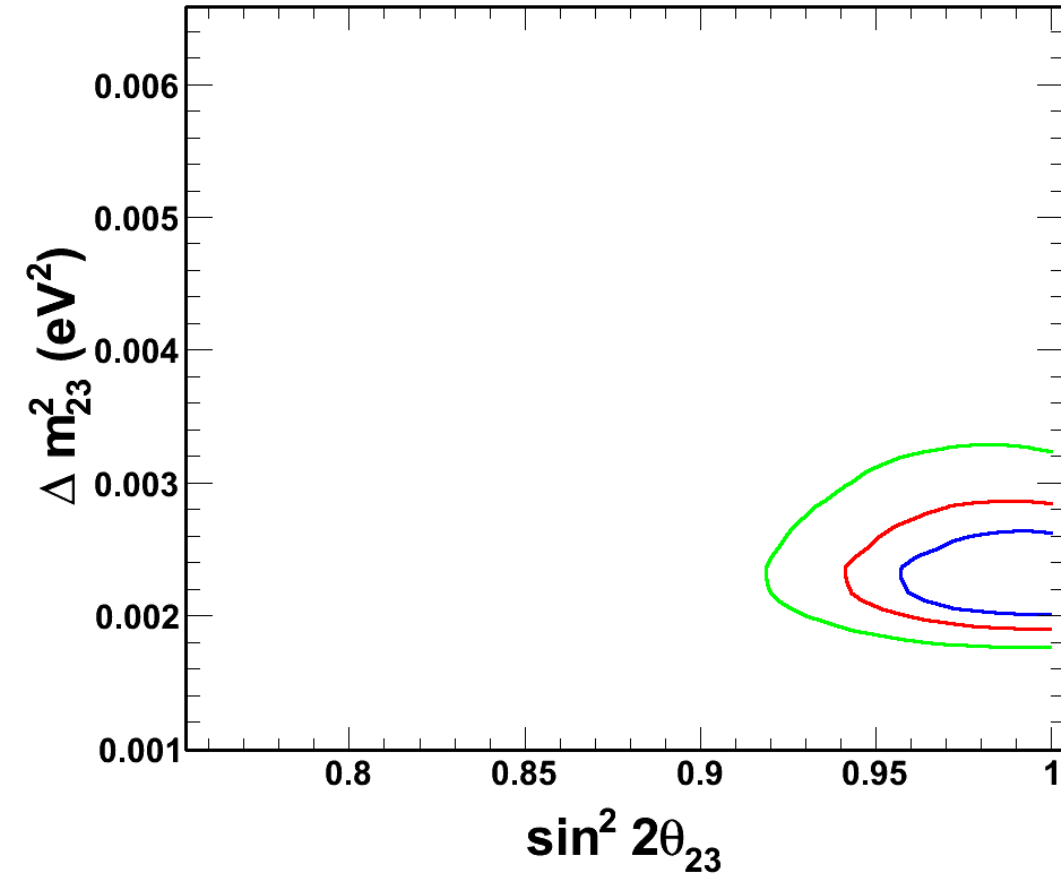
L/E Analysis – 3903 Days of Data



- Neutrino decay disfavored at 4.0σ
- Neutrino decoherence disfavored at 4.8σ

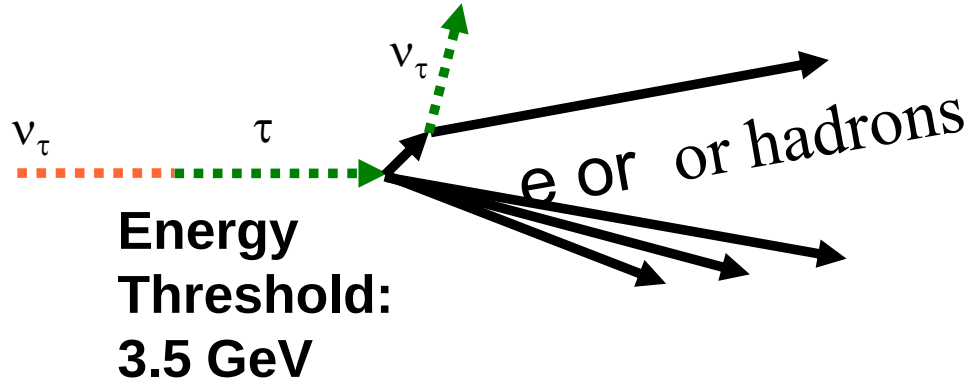
L/E	Best fit	90% C.L
Δm^2_{32}	2.5×10^{-3}	$[2.13 - 2.91] \times 10^{-3} \text{ eV}^2$
$\sin^2 2\theta_{23}$	1.0	> 0.93

Two Flavor Analysis



L/E	Best fit	1 σ CL
Δm^2_{32}	2.3×10^{-3}	$[2.1 - 2.46] \times 10^{-3} \text{ eV}^2$
$\sin^2 2\theta_{23}$	1.00	> 0.96 (90%)

ν_τ Events at Super-K



- » Many light producing particles
- » Most events are deep inelastic scattering interactions

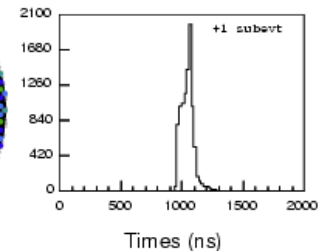
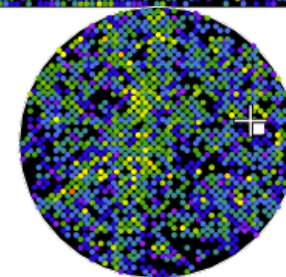
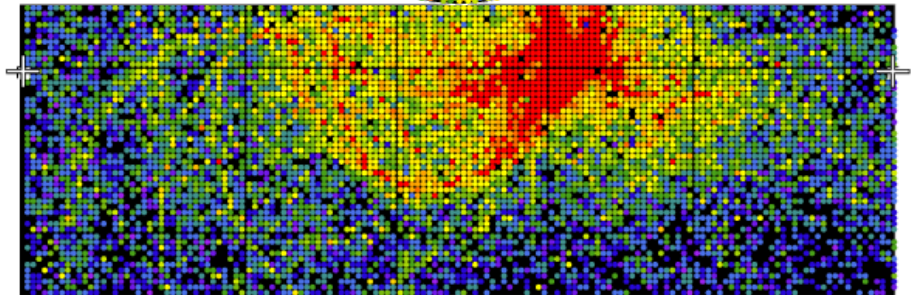
- » Complicated event topology complicate identification of the leading lepton
 - Use a Neural Network procedure
 - 80% efficient for signal
- » Negligible primary flux
 - Observed tau events would be oscillation induced

GOAL : Observe ν_τ events in the atmospheric data
How inconsistent is the “no appearance” hypothesis?

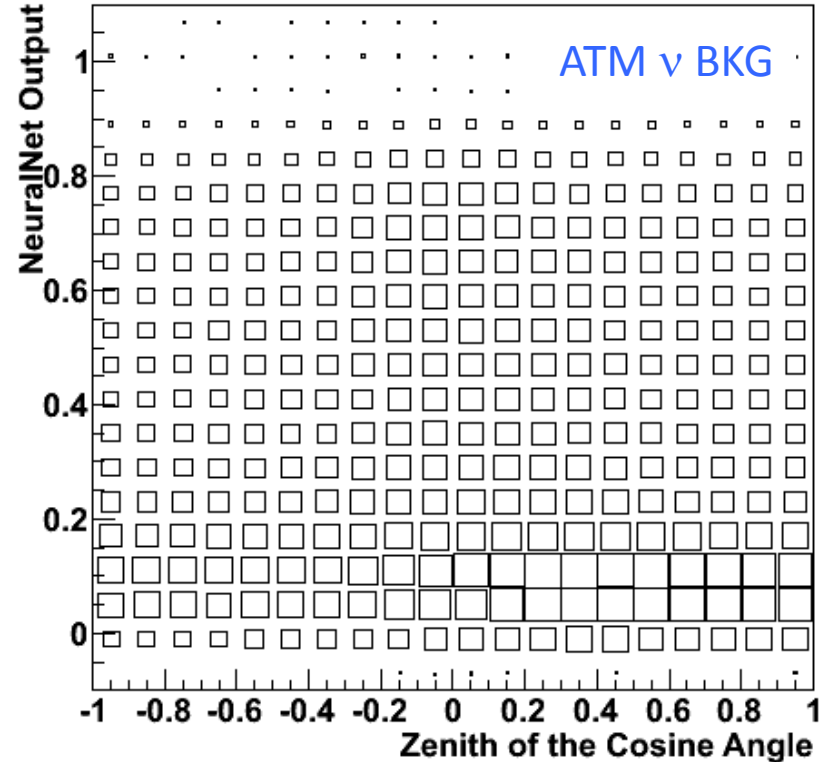
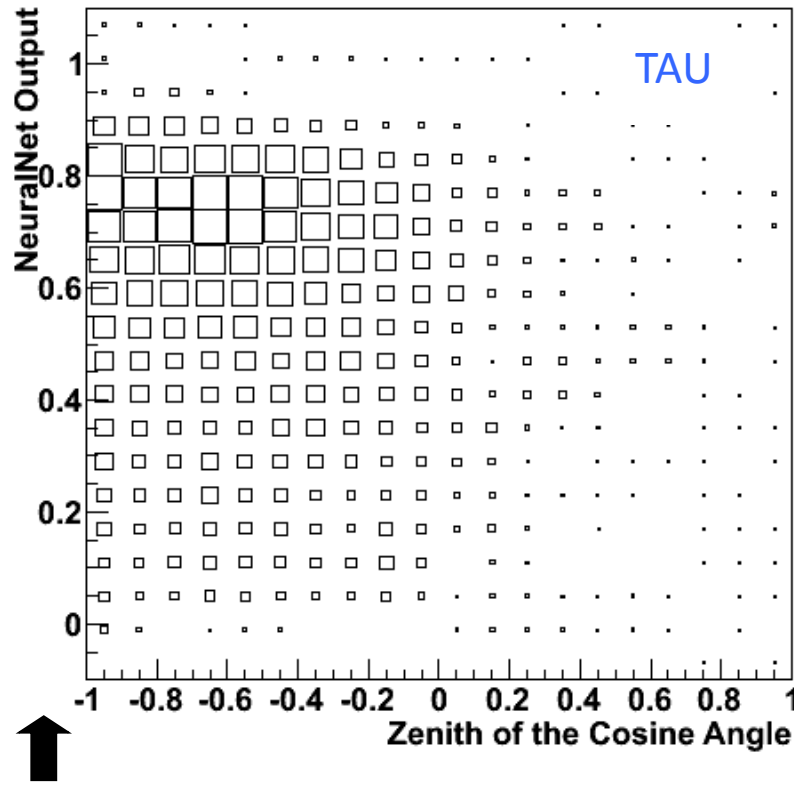
SK-I MC

Charge (pe)

- >26.7
- 23.3-26.7
- 20.2-23.3
- 17.3-20.2
- 14.7-17.3
- 12.2-14.7
- 10.0-12.2
- 8.0-10.0
- 6.2-8.0
- 4.7-6.2
- 3.3-4.7
- 2.2-3.3
- 1.3-2.2
- 0.7-1.3
- 0.2-0.7
- < 0.2



Neural Network and Unbinned likelihood fit

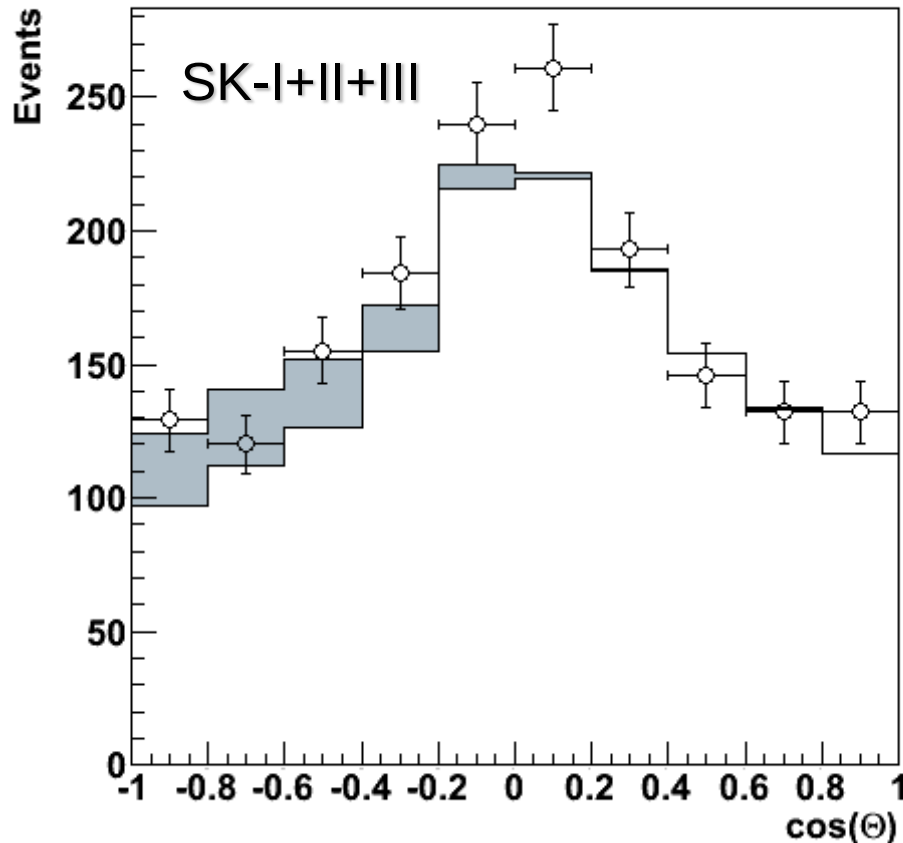


- Neural Network is built to separate tau events
 - Visible Energy, number of decay electrons, maximum distance to a decay electron, number of ring fragments, fraction of momentum in brightest right
- Fit normalization of signal and background PDFs as a function of NN output and Zenith Angle

Fit Results

If no ν_τ appearance, $\beta = 0$

$$Data = \alpha(\gamma) \times bkg + \beta(\gamma) \times signal$$



Fitted Excess

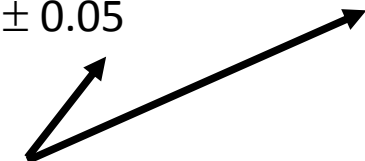


Atm ν BKG MC



Result	Background	Signal
SK-I	0.95	1.27
SK-II	0.96	1.47
SK-III	0.94	2.16
SK-I+II+III	0.94 ± 0.02	1.42 ± 0.35
DIS γ	1.10 ± 0.05	

Fitting Error only



□ Tau signal clearly appears in upward-going region

□ DIS fits to 1.10

□ Tau normalization fits to $1.42 \times$ expectation

Systematic Errors

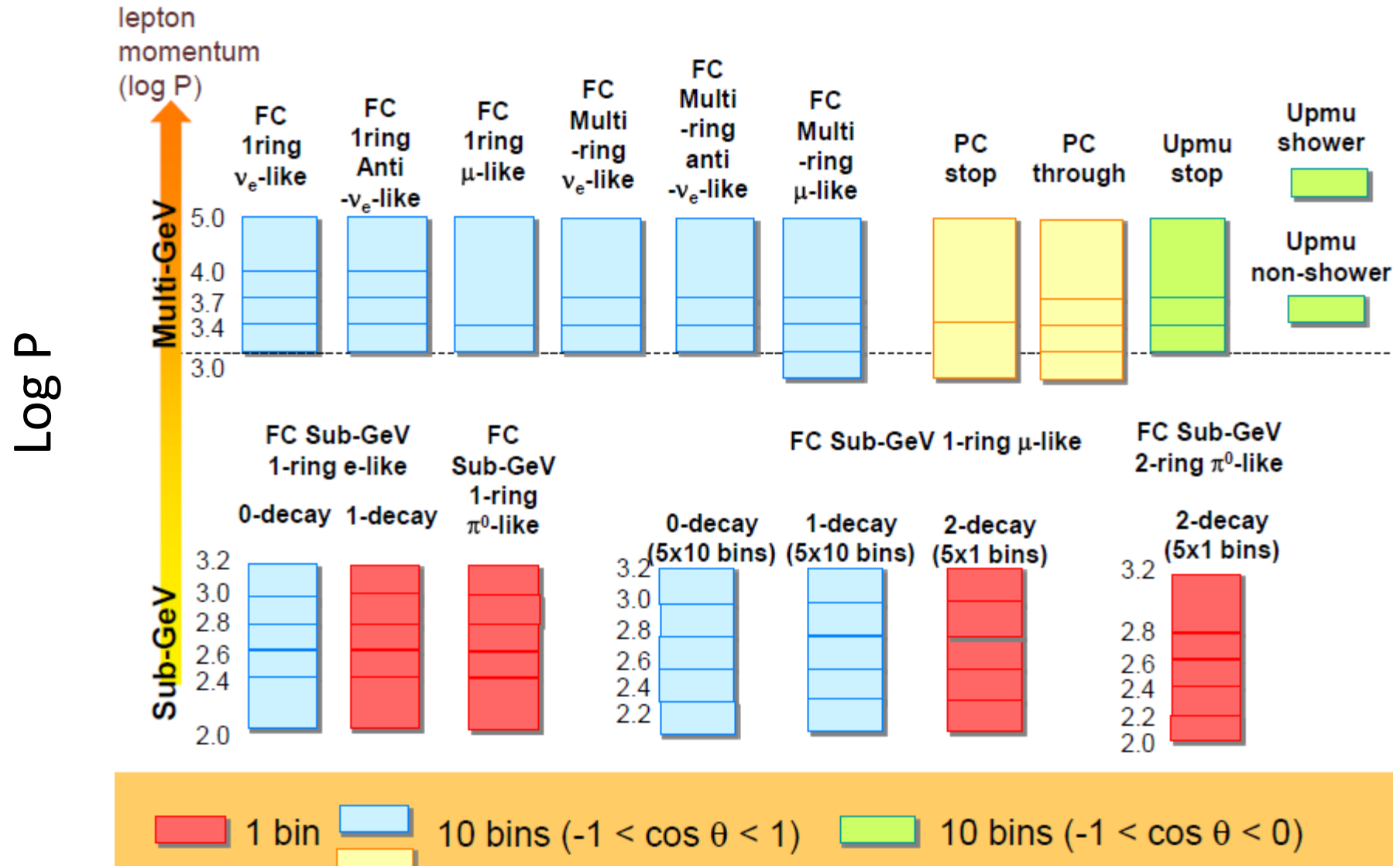
	+ %	- %
❑ Flux: Up/Down Ratio, Horizontal/ Vertical ratio , K/p ❑ X-sec: NC/CC ratio ❑ Dectector: Up/Down Energy cal. Asymmetry	7.9	8.5
❑ Oscillation Parameters: 1 - σ allowed atm.	5.4	1.3

$$\beta = 1.42 \pm 0.35 \text{ (stat)} \begin{matrix} +0.14 \\ -0.12 \end{matrix} \text{ (sys)}$$

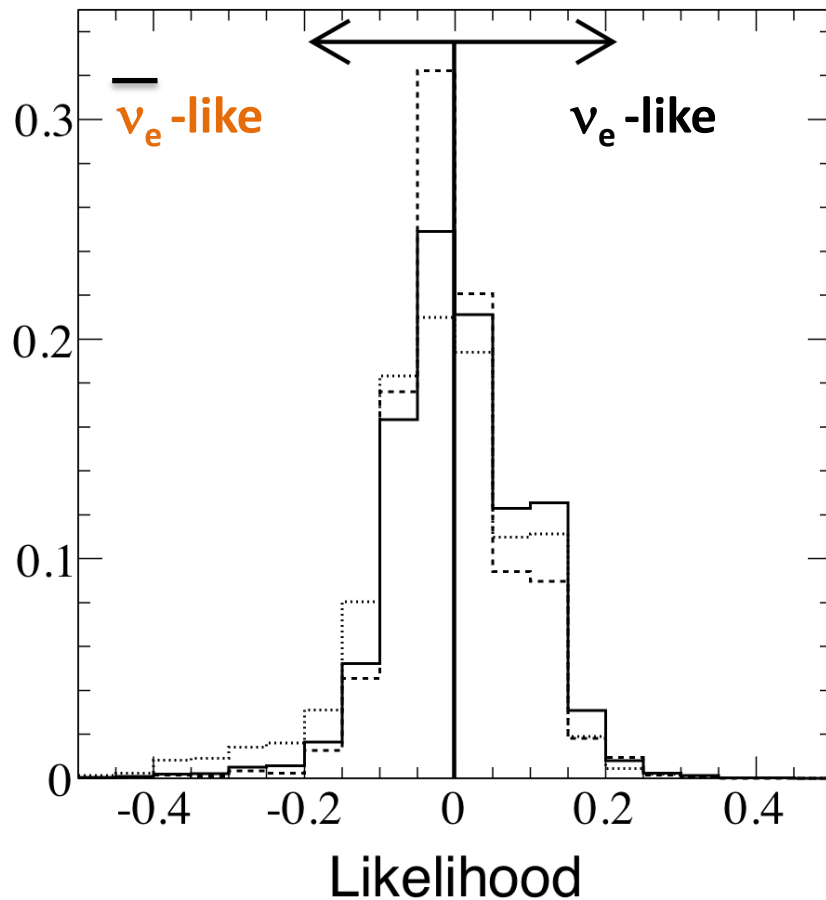
This corresponds to **180.1 $\pm$ 44.3 (stat) $^{+17.8}_{-15.2}$ (sys)**
events a **3.8 σ excess**
(Expected 2.7 σ significance)

SK Data disfavor 'no tau appearance' at 3.8 σ

Zenith Angle Analysis – 480 Bins in 18 Samples



Sample Selection : Multi-GeV Multi-Ring anti- ν_e and ν_e -like



— CC ν_e
- - CC anti- ν_e
..... CC $\nu_\mu + \text{anti-}\nu_\mu$

- Multi-ring events are in general more complicated
 - Many outgoing particles from DIS or multi- π interactions
- Likelihood function built on four variables
 - Number of decay electrons
 - Number of Cherenkov rings found
 - Transverse momentum

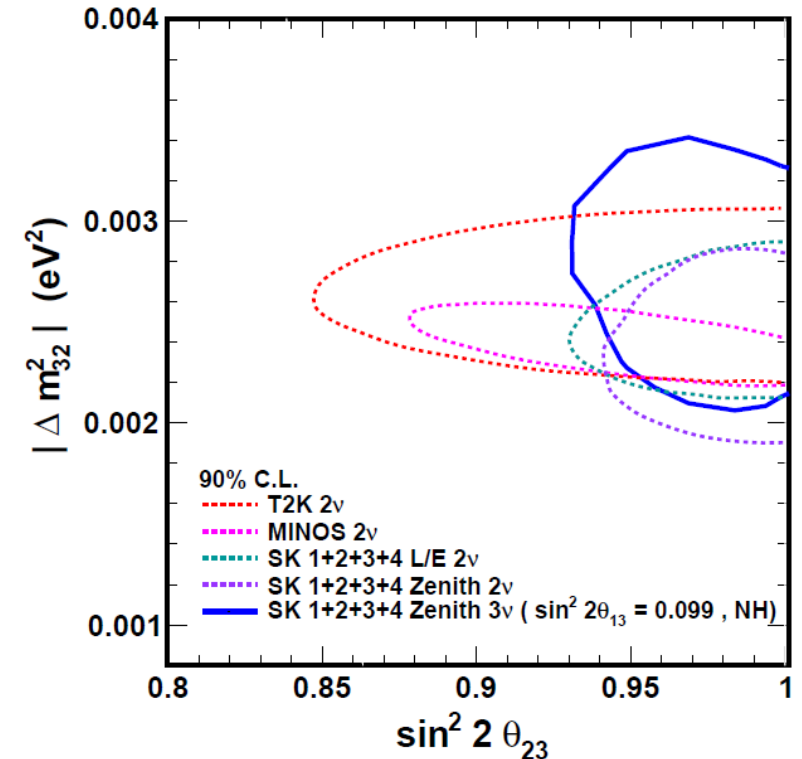
Variable	CC ne	CC anti-ne
N Rings	More	Fewer
N Decay e	More	Fewer
Transverse P	Larger	Smaller

Notes about Parameter Values

- In the studies below, unless specified otherwise the following inputs have been used to compute sensitivities

Parameter	Value	$\sin^2(2x)$	Comment
Δm^2_{32}	$2.4 \times 10^{-3} \text{ eV}^2$		Global Fit
$\sin^2\theta_{23}$	0.4-0.6	0.96 - 1.0	**
$\sin^2\theta_{13}$	0.025	0.10	Reactor Data
δ_{cp}	40 degrees		Min.P($\nu_\mu \rightarrow \nu_e$)
$\sin^2\theta_{21}$	0.31	0.85	Global Solar
Δm^2_{21}	$7.6 \times 10^{-5} \text{ eV}^2$		Global Solar
Hierarchy	Normal		Assumption

Neutrino 2012



** MINOS central value from Neutrino 2012: $\sin^2(2\theta_{23}) = 0.96$

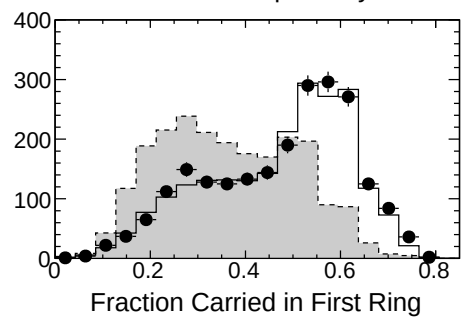
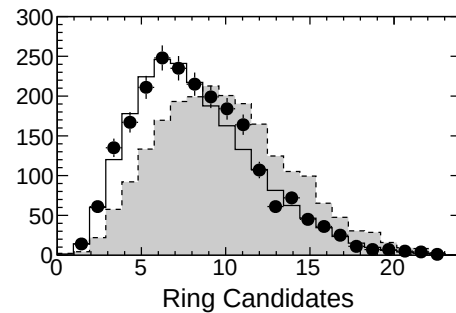
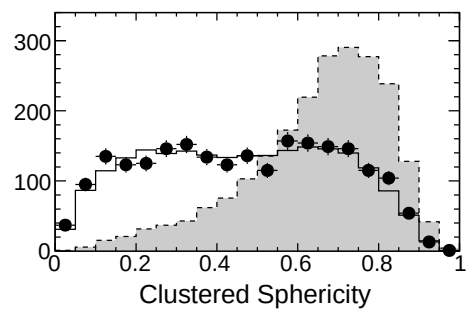
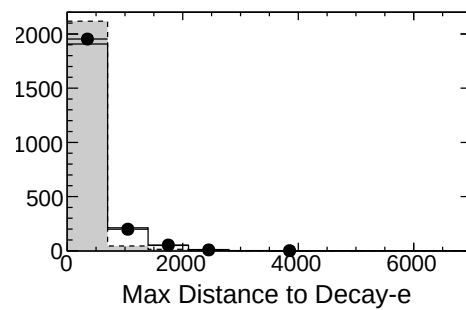
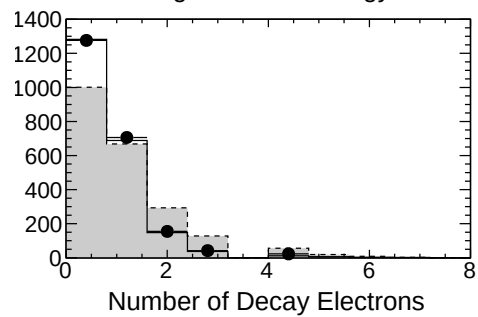
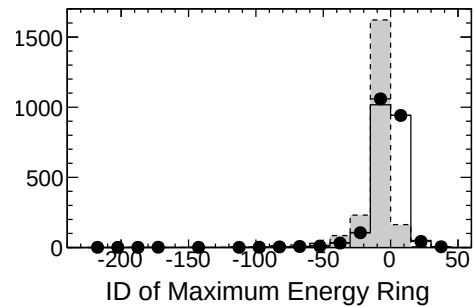
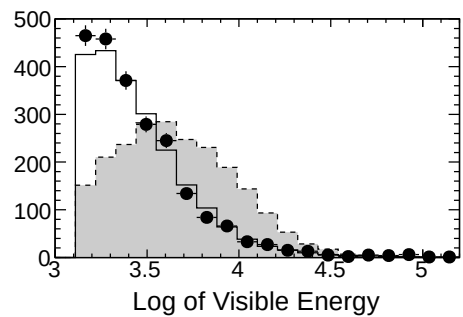
Systematic Error			fit value	σ
Flux normalization	$E_\nu < 1$ GeV		34.7	25 ^a
	$E_\nu > 1$ GeV		8.8	7 ^b
ν_μ/ν_e	$E_\nu < 1$ GeV		-1.9	2
	$1 < E_\nu < 10$ GeV		-2.5	3
	$E_\nu > 10$ GeV		-3.7	5 ^c
$\bar{\nu}_e/\nu_e$	$E_\nu < 1$ GeV		5.54	5
	$1 < E_\nu < 10$ GeV		1.13	5
	$E_\nu > 10$ GeV		-0.10	8 ^d
$\bar{\nu}_\mu/\nu_\mu$	$E_\nu < 1$ GeV		-0.48	2
	$1 < E_\nu < 10$ GeV		-1.35	6
	$E_\nu > 10$ GeV		-1.75	6 ^e
Up/down ratio	< 400 MeV	e -like	-0.07	0.1
		μ -like	-0.23	0.3
		0-decay μ -like	-0.84	1.1
	> 400 MeV	e -like	-0.61	0.8
		μ -like	-0.38	0.5
		0-decay μ -like	-1.29	1.7
	Multi-GeV	e -like	-0.53	0.7
		μ -like	-0.15	0.2
	Multi-ring Sub-GeV	μ -like	-0.15	0.2
	Multi-ring Multi-GeV	e -like	-0.23	0.3
		μ -like	-0.15	0.2
	PC		-0.15	0.2

Horizontal/Vertical ratio < 400 MeV	<i>e</i> -like	-0.01	0.1
	μ -like	-0.01	0.1
	0-decay μ -like	-0.03	0.3
> 400 MeV	<i>e</i> -like	-0.14	1.4
	μ -like	-0.19	1.9
	0-decay μ -like	-0.14	1.4
Multi-GeV	<i>e</i> -like	-0.33	3.2
	μ -like	-0.23	2.3
Multi-ring Sub-GeV	μ -like	-0.13	1.3
Multi-ring Multi-GeV	<i>e</i> -like	-0.29	2.8
	μ -like	-0.15	1.5
PC		-0.17	1.7
K/ π ratio in flux calculation		-12.9	10 ^f
Neutrino path length		-8.8	10
Sample-by-sample	FC Multi-GeV	-4.5	5
	PC + Up-stop μ	-7.1	5

About Systematics

- ❑ Super-K analysis considers 151 sources of systematic uncertainty from the usual cadre of errors
 - ❑ Flux, cross-section, detector performance
- ❑ This is a partial listing of things relevant to three-flavor issues

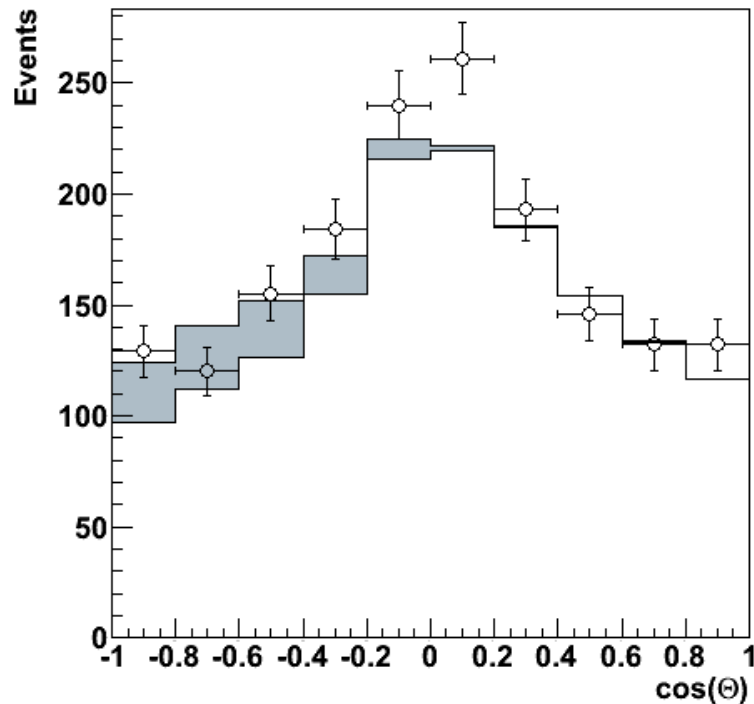
Error Source	Uncertainty
ν_e vs. anti- ν_e sample selection	7%
Charged-Neutral Pion Production	40%
Tau Production Cross section	25%
DIS Cross Section	5-10%
NC / CC Ratio	20%
Single-Pion Production	20%
Flux Normalization above 1 GeV	7%
Flux Ratio ν to $\bar{\nu}$ above 1 GeV	5-8%



Examine both the tau-like and non-tau like events in zenith

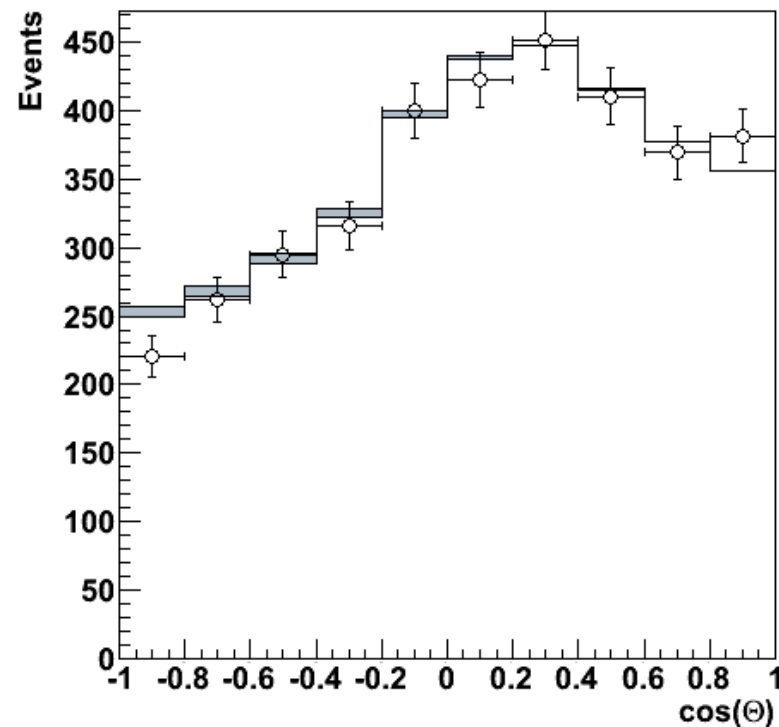
Tau like

Zenith Distribution



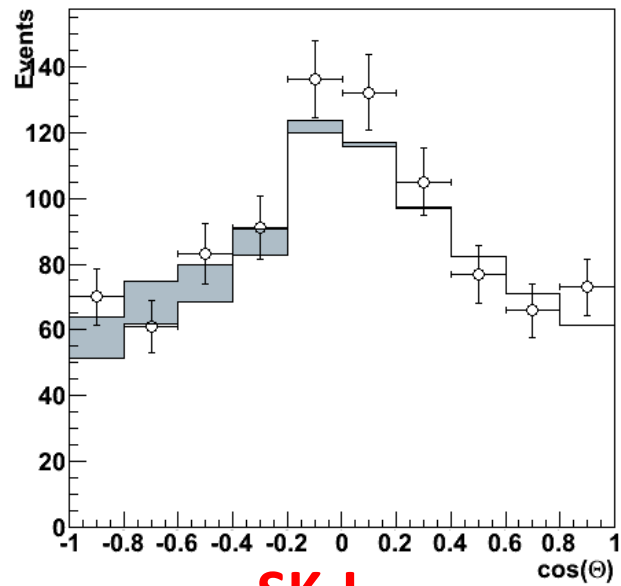
Non-Tau Like

Zenith non-tau like



Fit results separately for SK-I, SK-II, SK-III

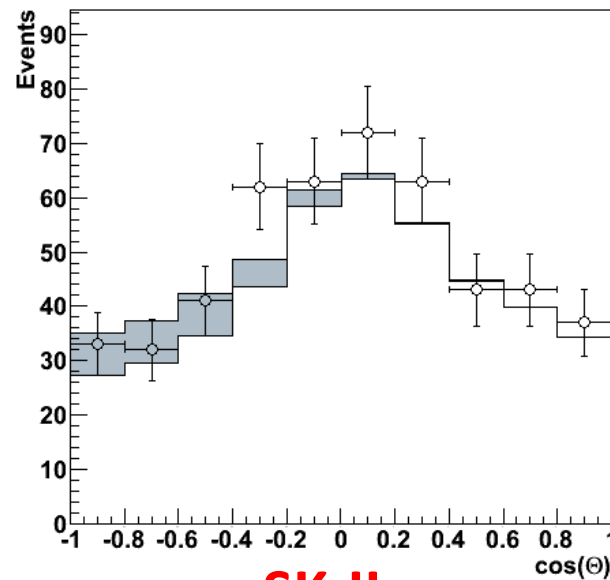
Zenith Distribution



SK-I

$$B = 1.27 \pm 0.49$$

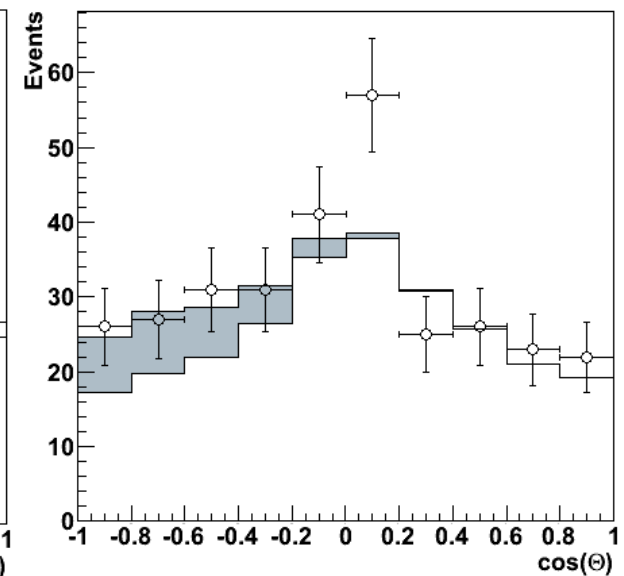
Zenith Distribution



SK-II

$$B = 1.47 \pm 0.62$$

Zenith Distribution

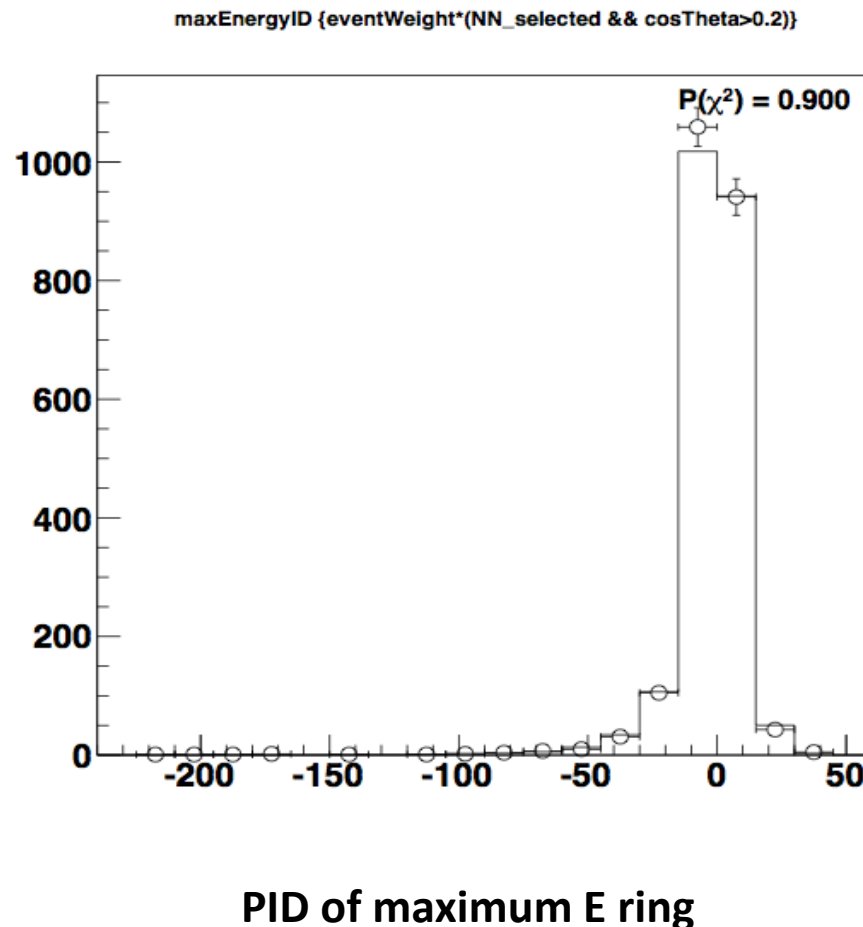
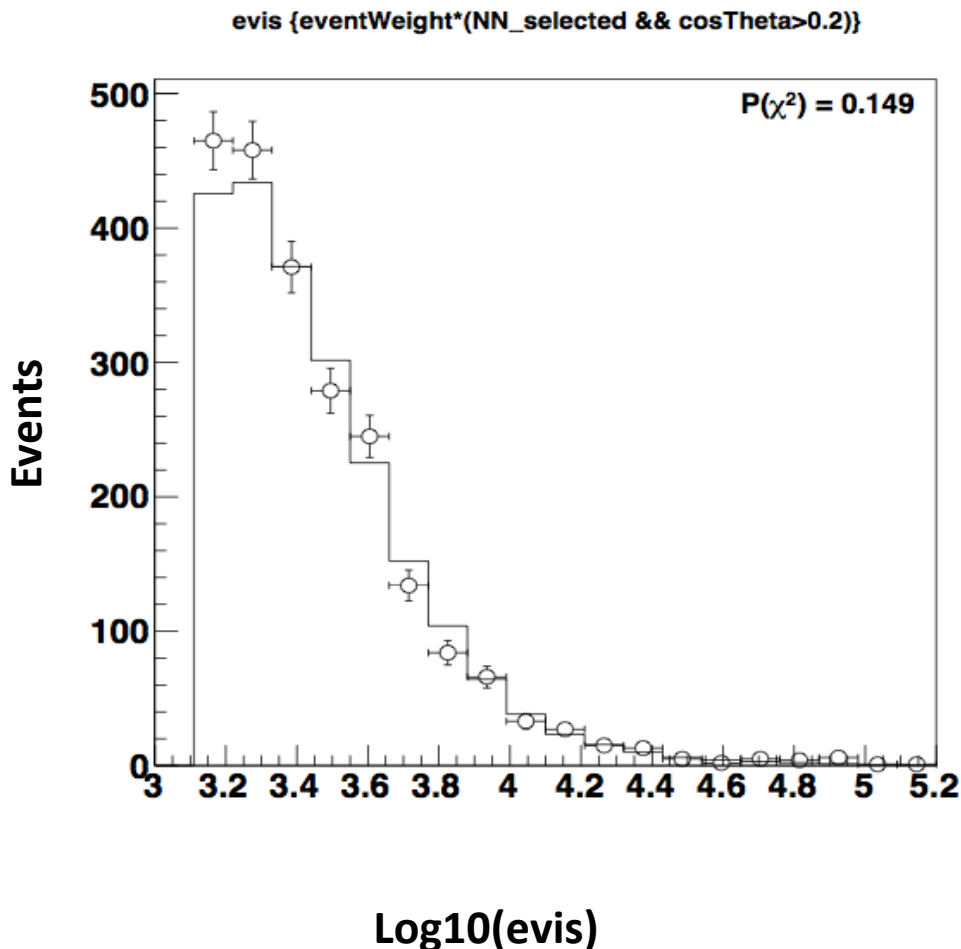


SK-III

$$B = 2.16 \pm 0.78$$

Now compare Downward going Data/MC

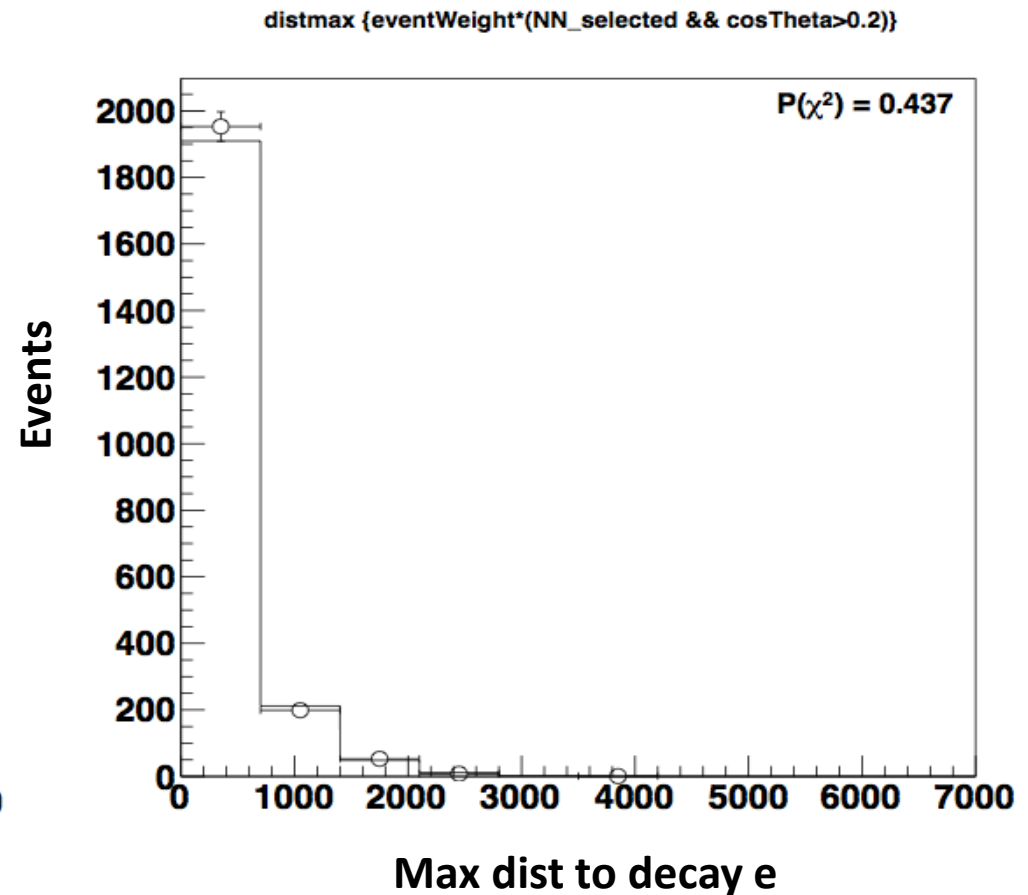
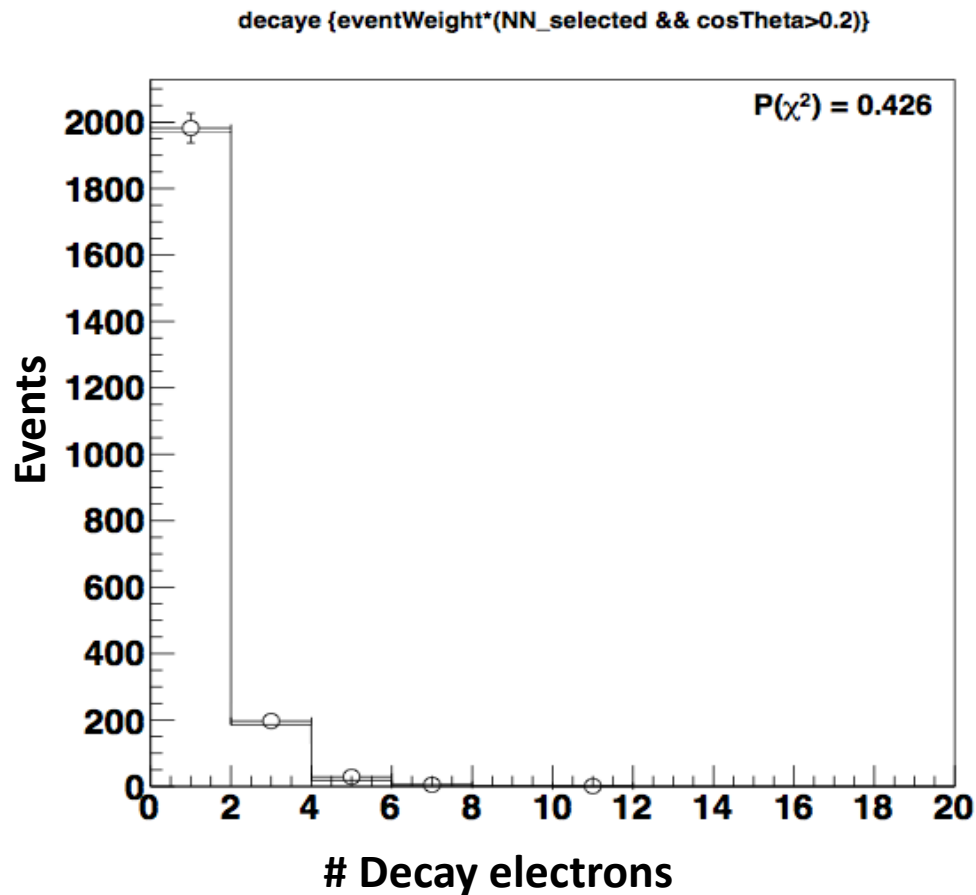
SKI + SKII + SKIII Combined



Note: Adjusted to fit values (DIS, bkg norm)

Data vs MC (downward)

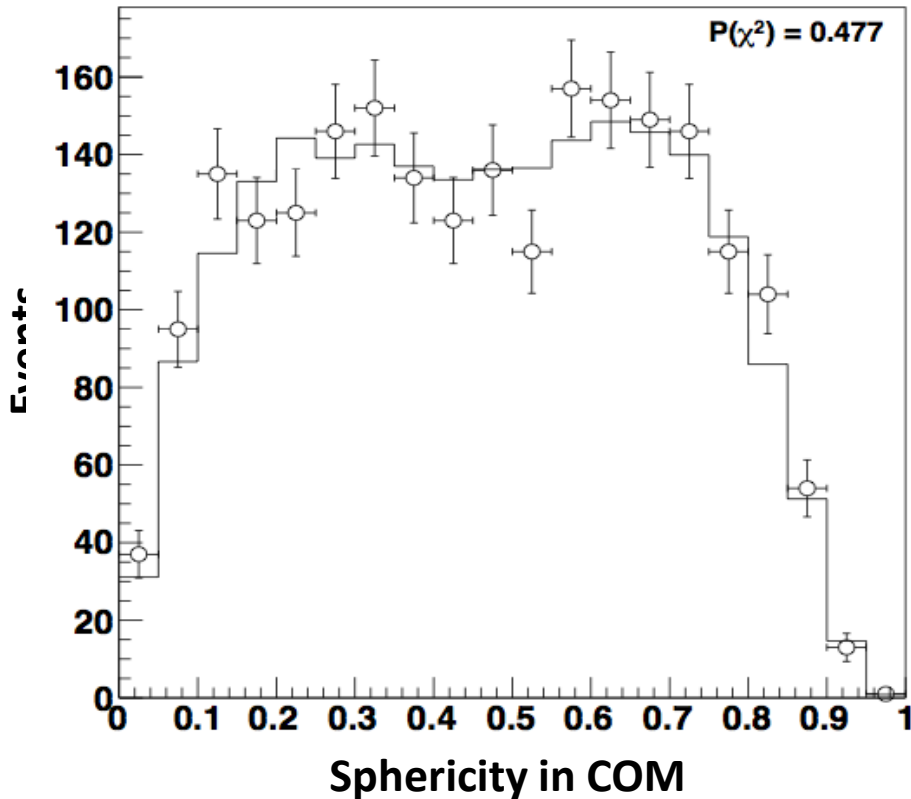
SKI + SKII + SKIII Combined



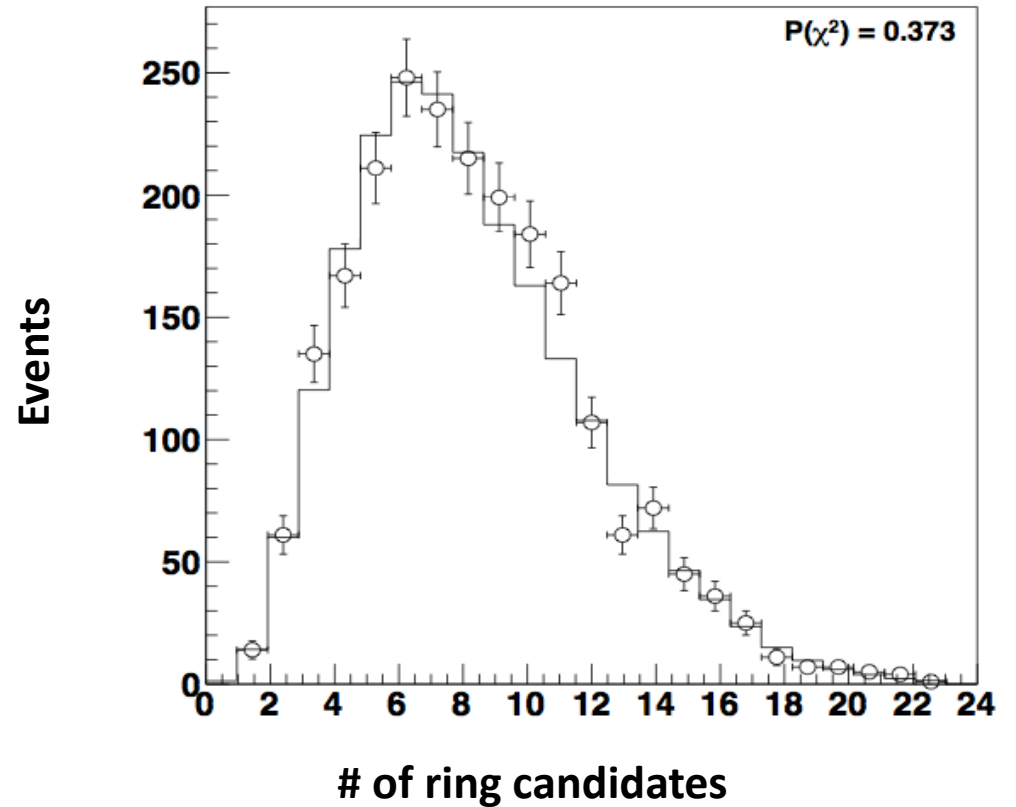
Data/MC (downward)

SKI + SKII + SKIII Combined

gmpsh {eventWeight*(NN_selected && cosTheta>0.2)}

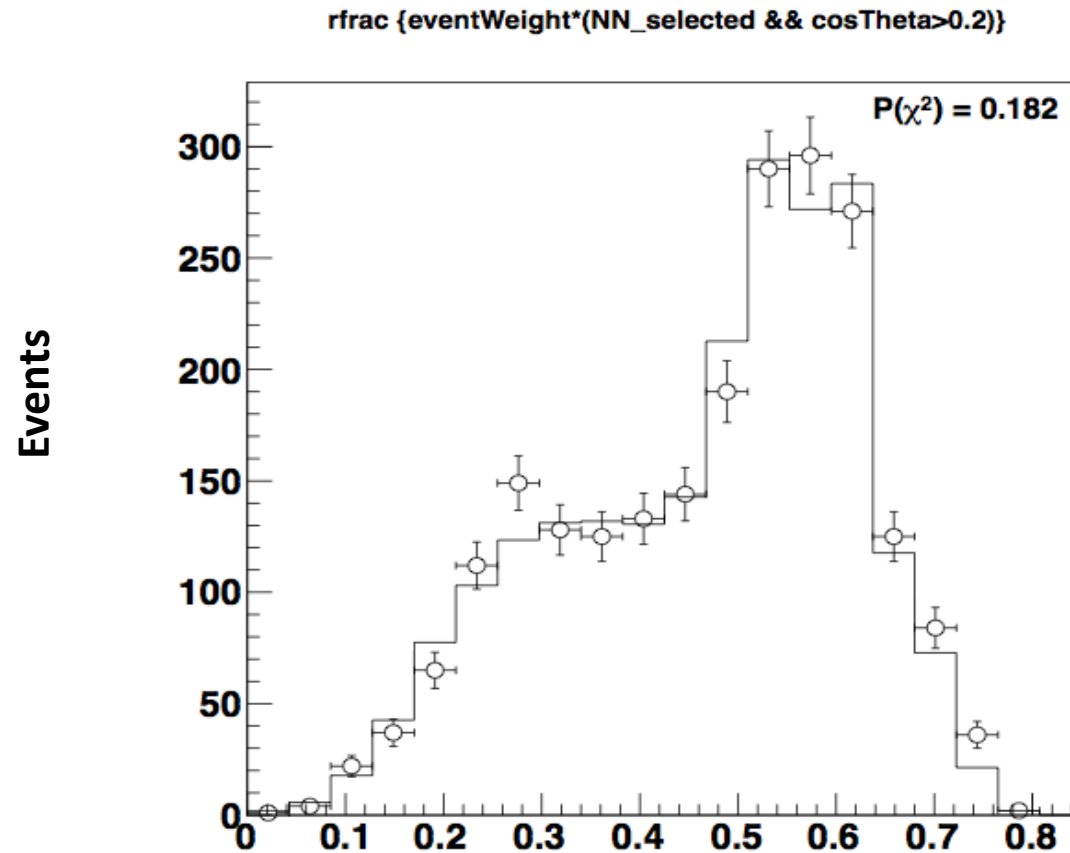


numseeds {eventWeight*(NN_selected && cosTheta>0.2)}



Data/MC (downward)

SKI + SKII + SKIII Combined

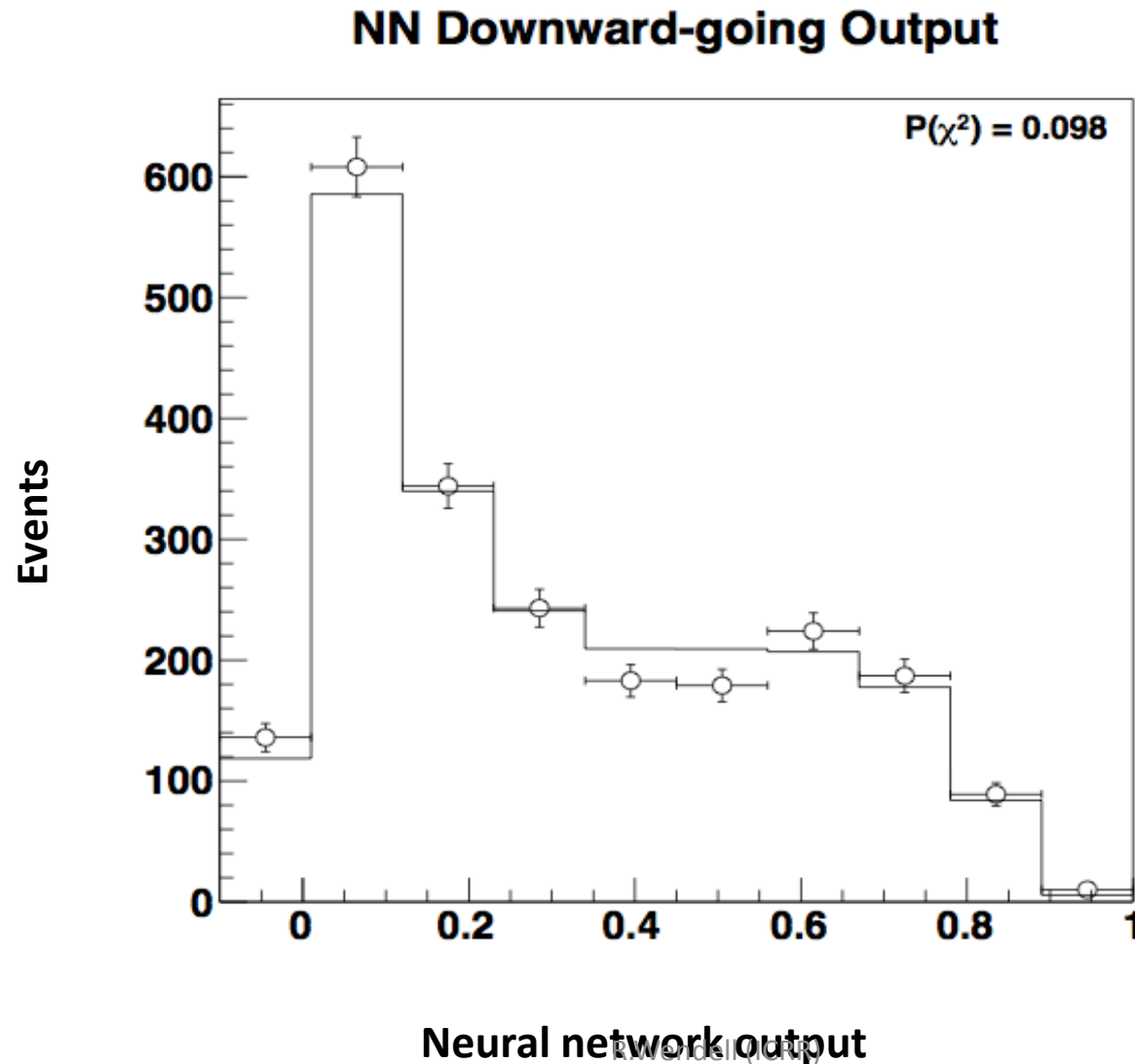


Fraction of photo electrons in leading ring

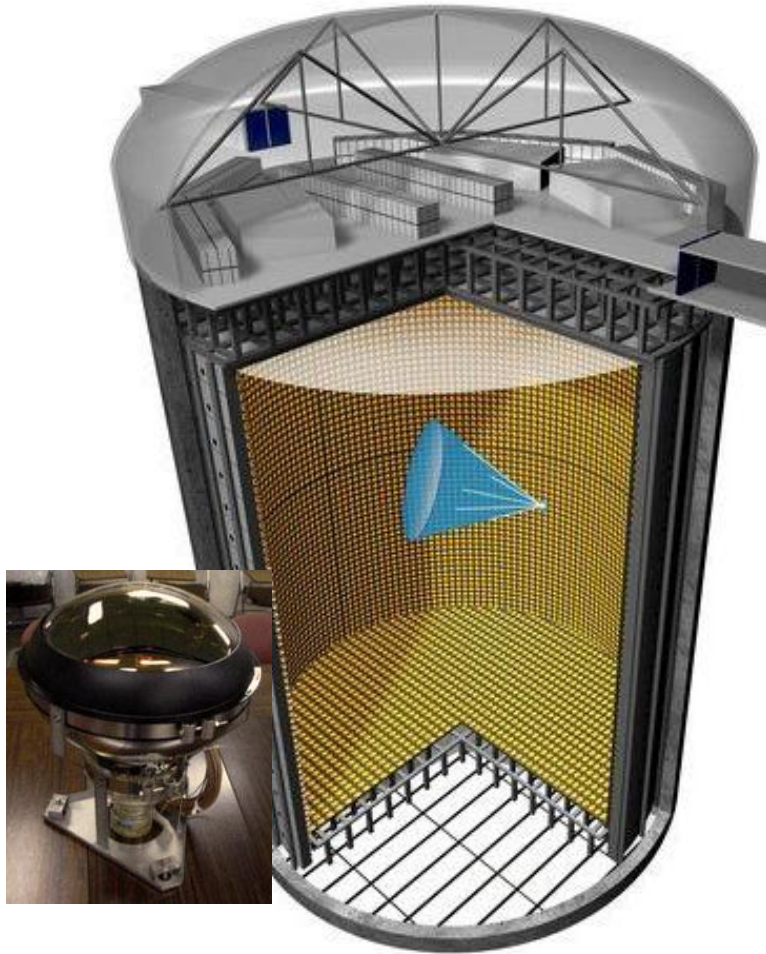
R.Wendell (ICRR)

NN Output Data/MC (downward)

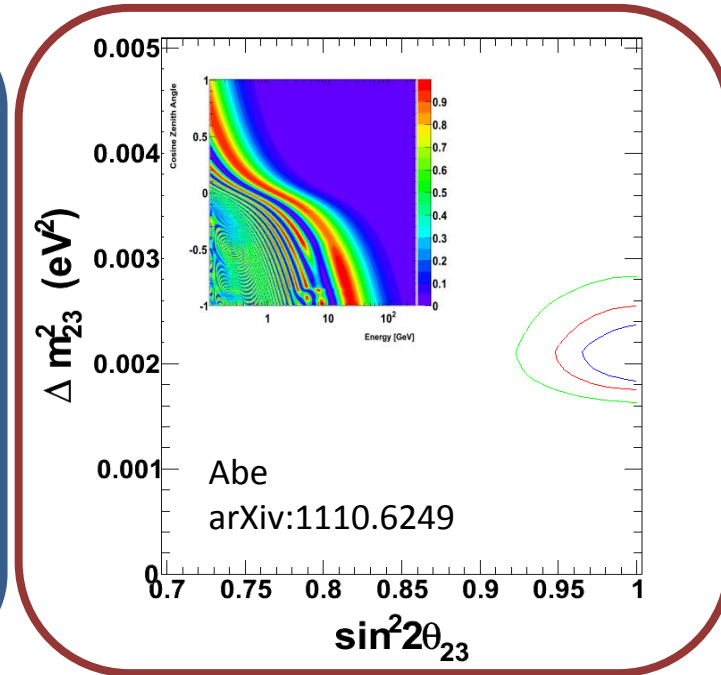
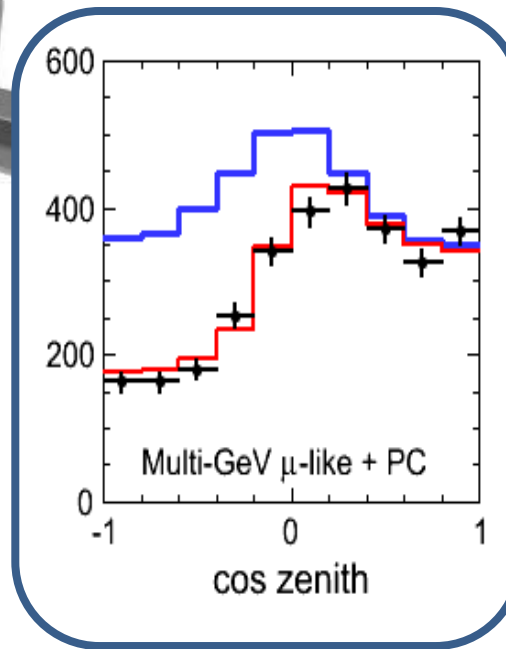
SKI + SKII + SKIII Combined



Super-Kamiokande : Introduction



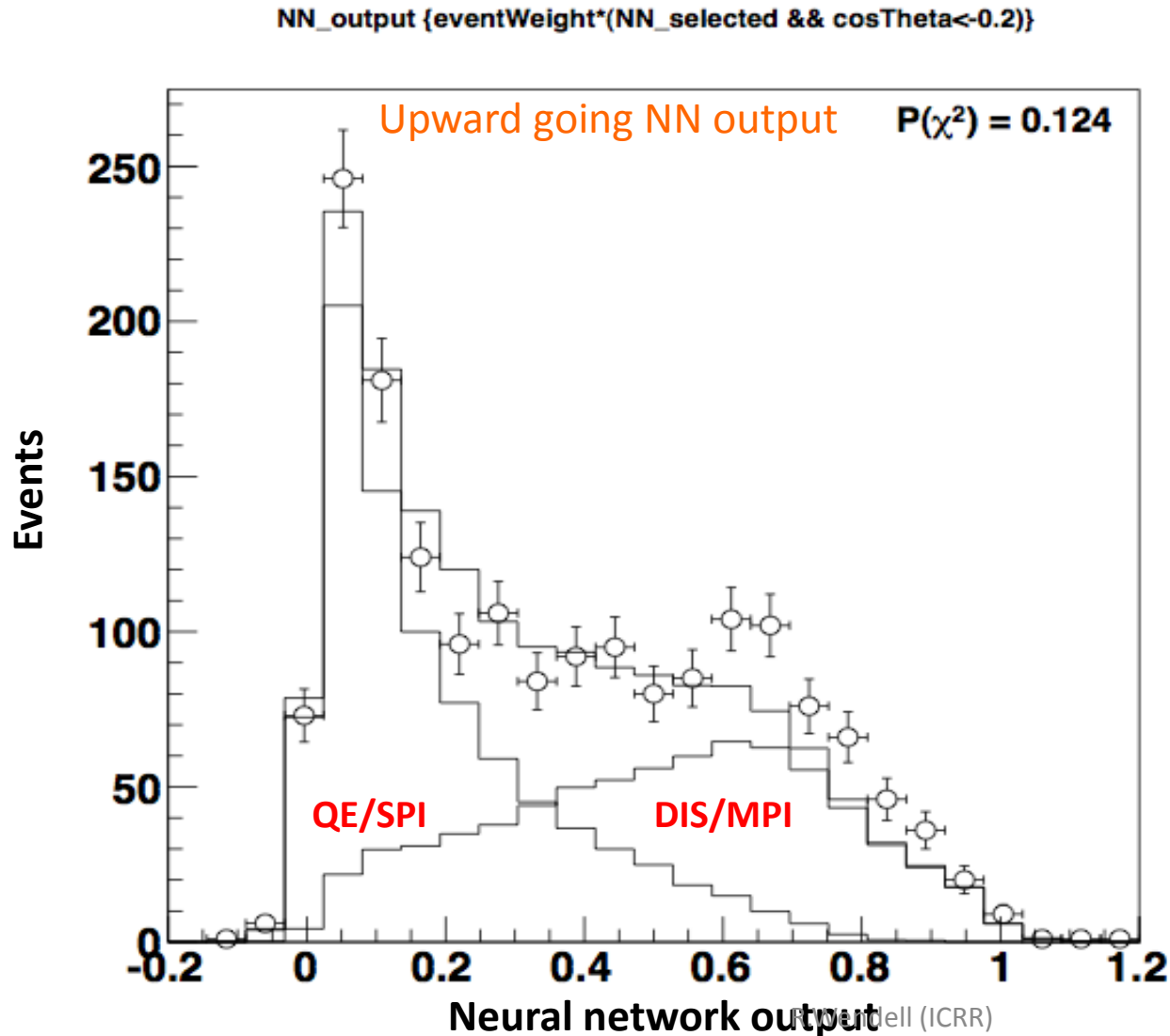
$$P(\nu_\mu \rightarrow \nu_\mu) = 1 - \sin^2 2\theta_{23} \sin^2(1.27 \Delta m_{23}^2 L/E)$$



- ❑ 22.5 kton FV Volume
- ❑ Ring Imaging Water Cherenkov detector
- ❑ 11,146 20" Phototubes
- ❑ Data taken over four periods since 1996

- ❑ SK discovered atmospheric neutrino oscillations exploiting the disappearance of upward-going muon events
- ❑ Currently **statistics limited**
 - ❑ Today: **3903 days of data**

Why we need to treat DIS separately



The neural net clearly separates
The bkg into DIS and non-DIS.
DIS events look more like taus.

The relative signal/bkg “slope”
changes by increasing the
amount of DIS events. So we
Need to adjust the central value in
the fit.