



# Final results from OPERA



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on behalf of the OPERA Collaboration





**26 institutions  
~150 physicists**



IRB Zagreb



METU Ankara



Technion Haifa



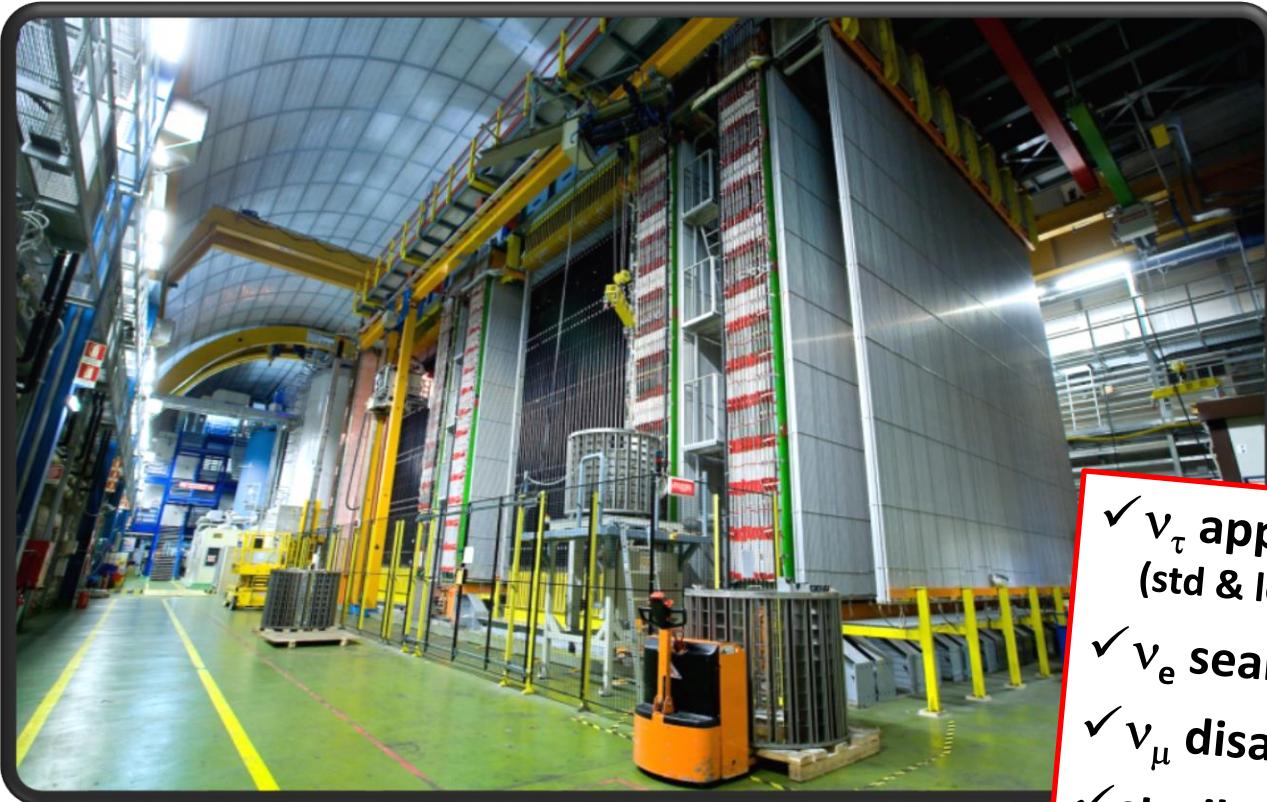
Bari  
Bologna  
LNF Frascati  
LNGS  
Napoli  
Padova  
Roma  
Salerno



LAPP Annecy  
IPHC Strasbourg



INR Moscow  
LPI Moscow  
SINP MSU Moscow  
JINR Dubna



- ✓  $\nu_\tau$  appearance (std & looser selection)
- ✓  $\nu_e$  search update
- ✓  $\nu_\mu$  disappearance
- ✓ sterile neutrinos
- ✓ non-oscillation physics

LHEP Bern

IHE Brussels

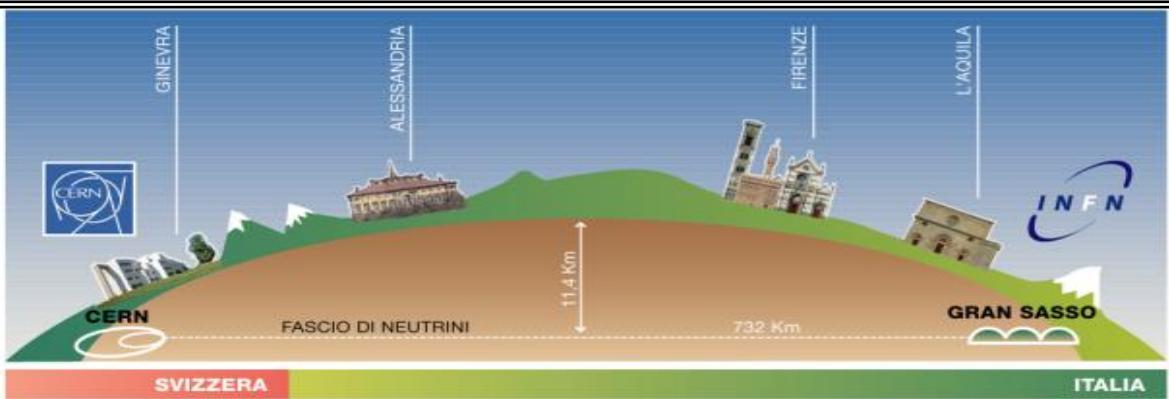
Hamburg

Jinjiu

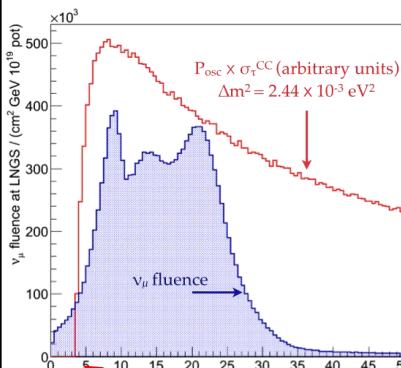
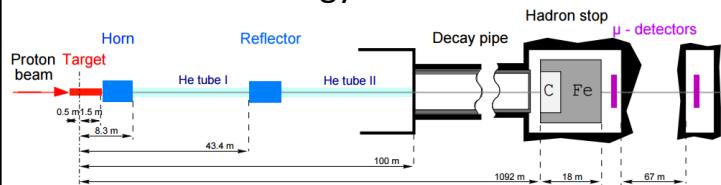
# Oscillations Project with Emulsion TRacking Apparatus

[CERN-SPSC-2000-028, 2000]

- Long baseline experiment: 735 km
- Aim: verify the  $\nu_\mu \rightarrow \nu_\tau$  oscillations at atmospheric  $\Delta m^2$  scale
- How:  $\nu_\tau$  appearance on event-by-event basis in a  $\nu_\mu$  beam



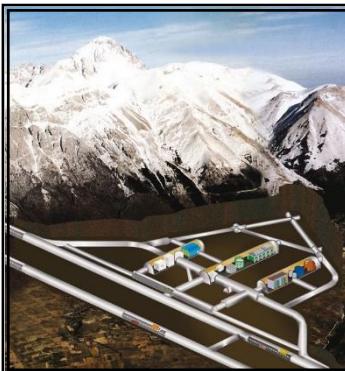
## Conventional muon neutrino beam neutrino mean energy: 17 GeV



- Optimized for  $\nu_\tau$  appearance at LNGS

- Maximize the number of  $\nu_\tau$  CC interactions

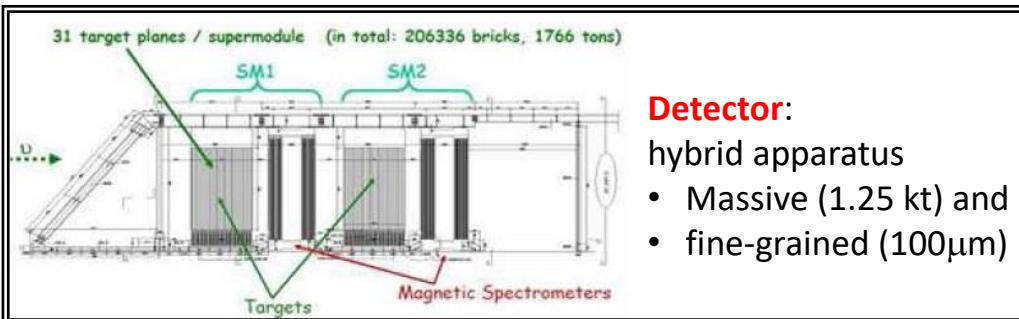
$(\nu_e + \bar{\nu}_e)/\nu_\mu$	0.9 %
$\bar{\nu}_\mu/\nu_\mu$	2.1%
$\nu_\tau$ prompt	negligible



## Low background environment

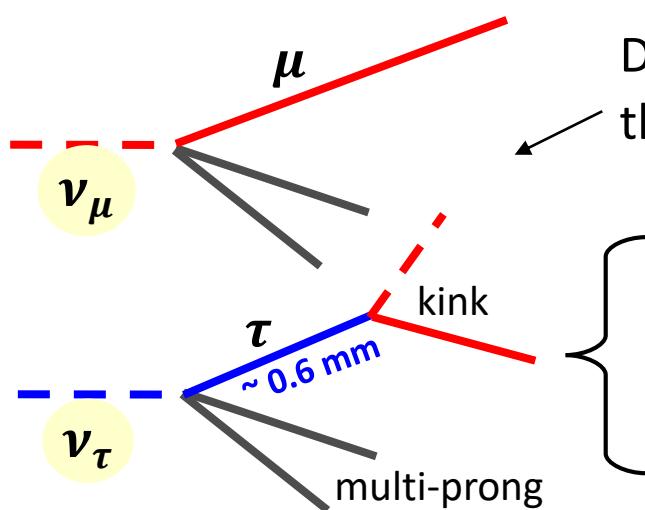
Laboratori Nazionali del Gran Sasso (Italy)

- 1400 m rock overburden
- atm.  $\mu$  reduction  $\sim 10^6$  [ $1\mu/(m^2 \cdot h)$ ]
- low radioactivity rock



- Detector:**  
hybrid apparatus
  - Massive (1.25 kt) and
  - fine-grained (100  $\mu$ m)

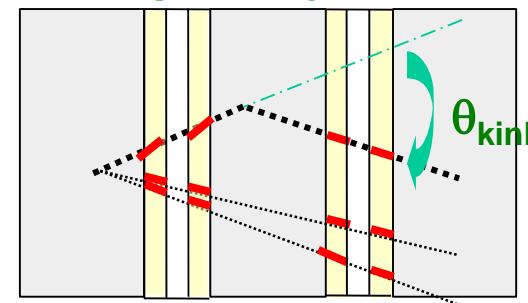
# The $\nu_\tau$ detection challenge



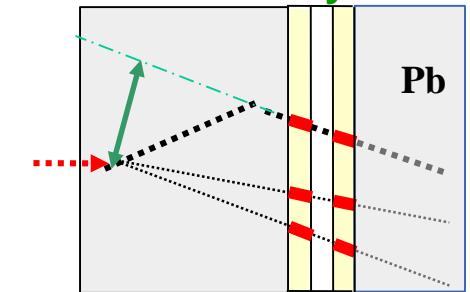
Detect a few  $\nu_\tau^{CC}$  from the bulk of  $\nu_\mu^{CC}$

$\tau^- \rightarrow \mu^- \nu_\tau \nu_\mu$	17 %
$\tau^- \rightarrow e^- \nu_\tau \nu_e$	18 %
$\tau^- \rightarrow h^- \nu_\tau n(\pi^0)$	50 %
$\tau^- \rightarrow \pi^- \pi^+ \pi^- \nu_\tau n(\pi^0)$	14 %

“long” decays: kink



“short” decays: I.P.



Modular detector of “Emulsion Cloud Chambers” (or bricks)

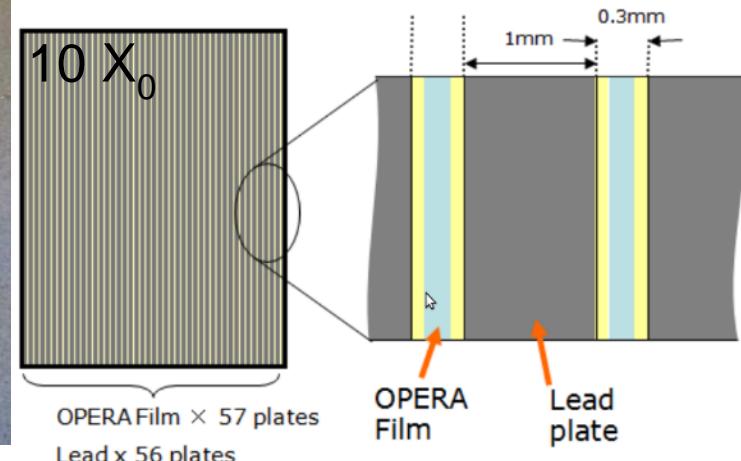
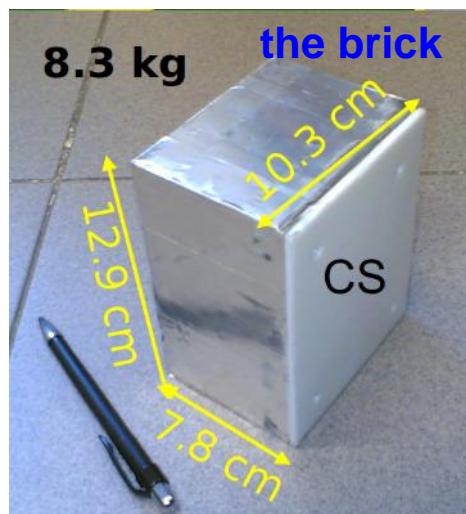
Match the needs for:

Large mass

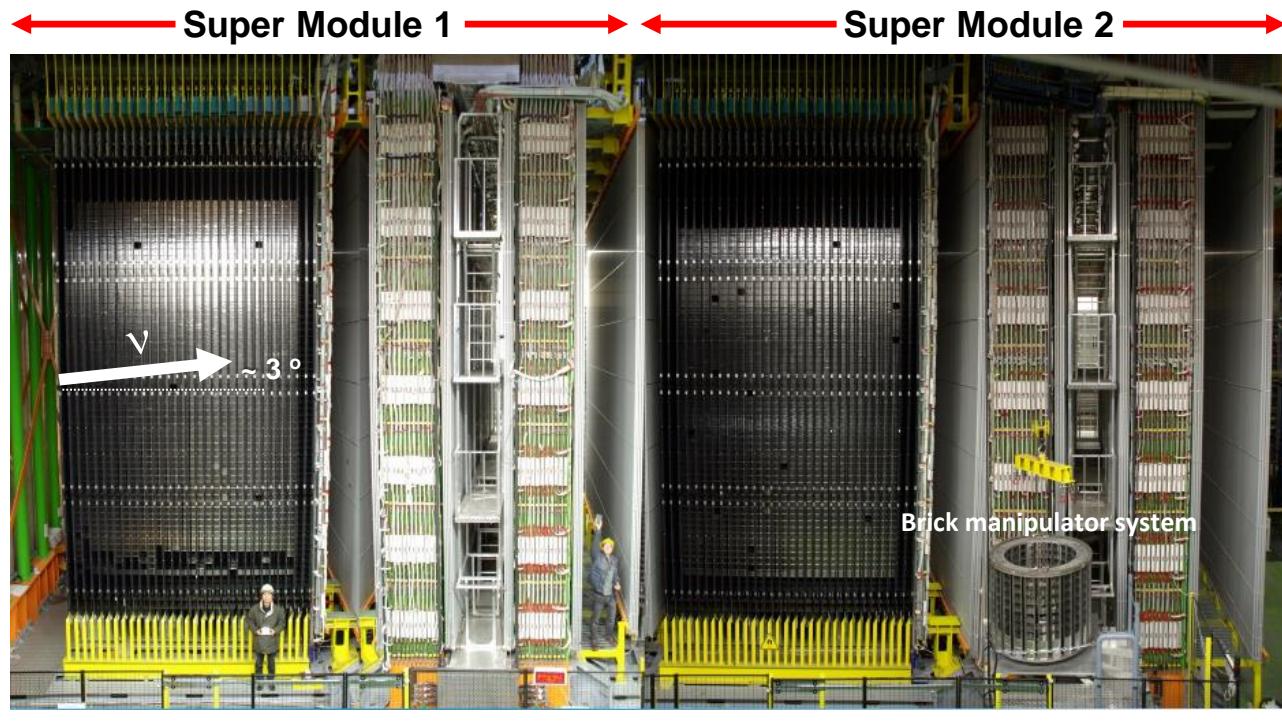
$$N_\tau \propto (\Delta m^2)^2 M_{\text{target}}$$

Extreme granularity

$\sim \mu\text{m}$  space resolution



# The OPERA detector



Target section ( $6.7 \times 6.7 \text{ m}^2$ ):

- Target
  - ~ 625 ton
  - ~ 75000 bricks in 27 walls
- Target Tracker
  - 31 XY doublets of 256 scintillator strips planes

Tracking of the target region

Brick selection

Calorimetry

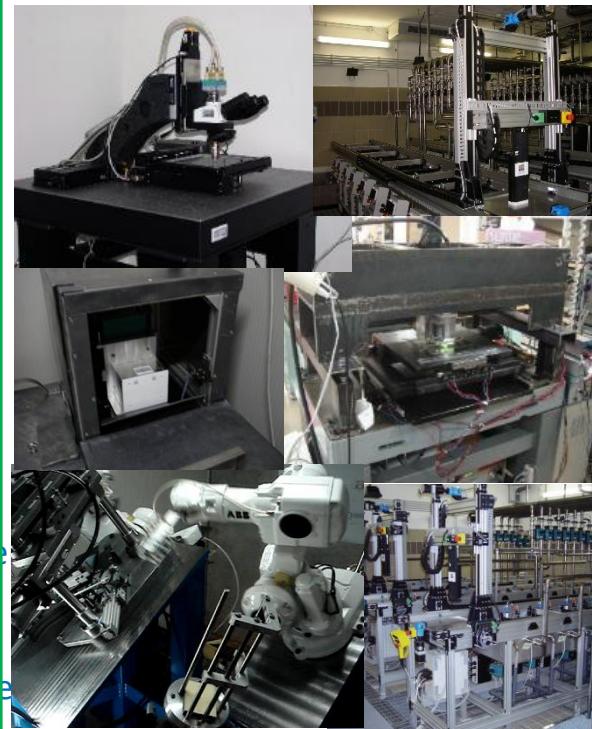
Muon spectrometer ( $8 \times 10 \text{ m}^2$ )

- 1.53 T magnet
- 22 XY RPC planes +  
2 RPC planes rotated by  $42.6^\circ$
- 6 stations of 4-fold drift tubes layer

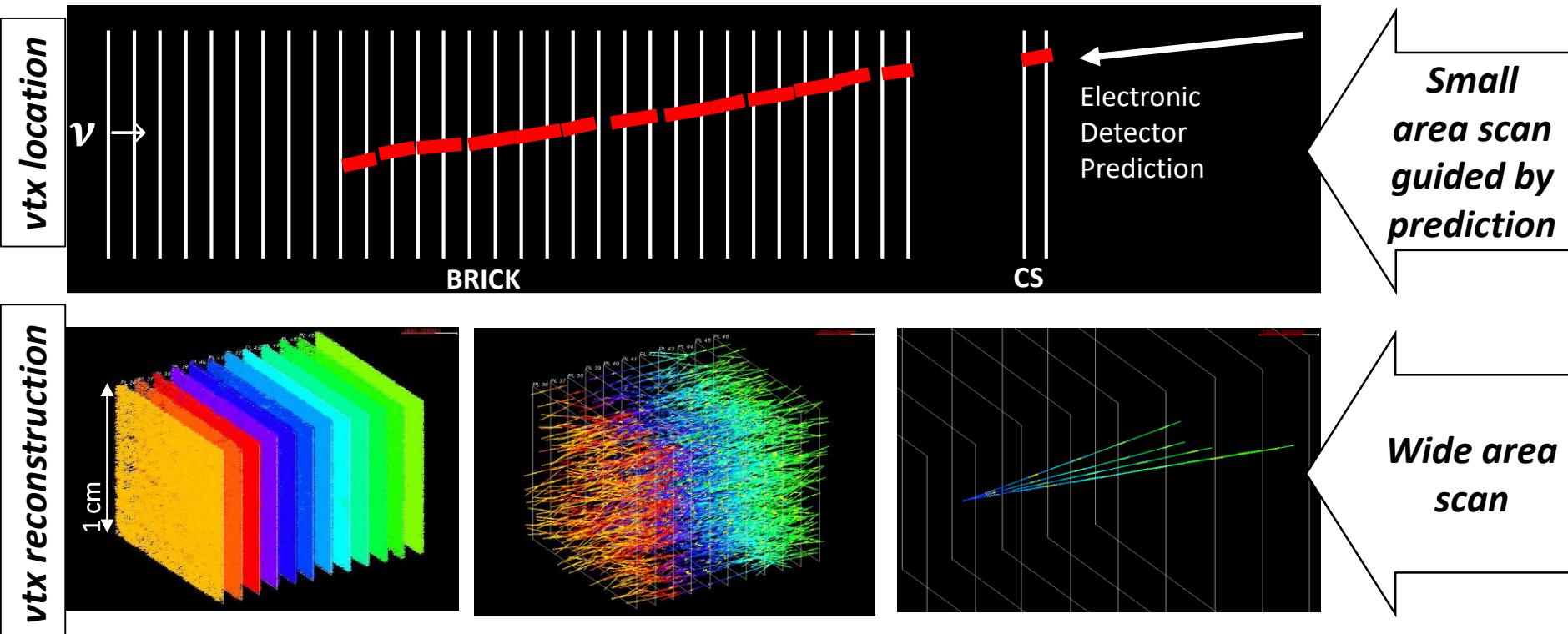
$\mu$  Identification +  
charge and momentum measurement

+ several ancillary facilities “off-site”:

- Assembly of bricks (LNGS)
- Brick Manipulator System (LNGS)
- Labelling and X ray marking (LNGS)
- Automatized development (LNGS)
- Scanning of CS doublets (LNGS+JP)
- Scanning bricks (European Labs + JP)



# Vertex hunting in the brick



- 0) tracks tagged in the CS films followed upstream to **stopping point**
- 1) 1 cm<sup>3</sup> **volume centered in the stopping point** scanned and tracks reconstructed
- 2) cosmic ray tracks (from a dedicated exposure) used for the fine **alignment** of films
- 3) passing through tracks discarded, the **vertexing algorithm** reconstructs the vertex
- 4) Short-lived particle decays identified (**decay search**)

# $\nu_\mu \rightarrow \nu_\tau$ background characterization

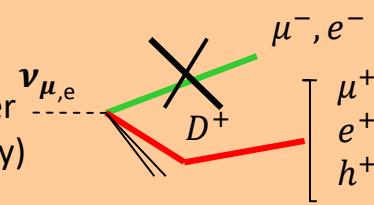
[JHEP 1311 (2013) 036]

## Monte Carlo simulation benchmarked on control samples

In order of decreasing relevance ↓

### CC with charm production (all channels)

If primary lepton is not identified and the daughter charge is not (or incorrectly) measured



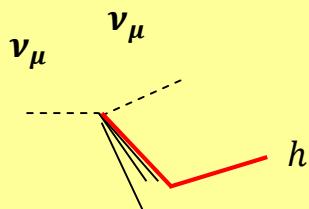
MC tuned on CHORUS data (cross section and fragmentation functions), validated with measured OPERA charm events.

Reduced by "track follow down", procedure and large angle scanning

[Eur.Phys.J. C74 (2014) 2986]

### Hadronic interactions

Background for  $\tau \rightarrow h$



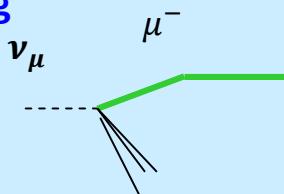
FLUKA + pion test beam data

Reduced by large angle scanning and nuclear fragment search

[PTEP9 (2014) 093C01]

### Large angle muon scattering

Background for  $\tau \rightarrow \mu$



Measurements in the literature (Lead form factor)

Improved MC simulations

[IEEE Trans.Nucl.Sci. 62 (2015) no.5, 2216-2225]

# $\nu_\tau$ appearance discovery (2015)

[Phys. Rev. Lett. 115, 121802 (2015)]

The 5 years long CNGS run

- $1.8 \times 10^{20}$  p.o.t. collected (80% of the design)
- 19505  $\nu$  interactions in the emulsion targets.
- 5  $\nu_\tau$  candidate events fulfill kinematical selection [S/B ratio ~10]

## Observed Data: 4 hadronic + 1 muonic candidates

Channel	Expected		Observed
	background	Expected signal	
$\tau \rightarrow 1h$	$0.04 \pm 0.01$	$0.52 \pm 0.10$	3
$\tau \rightarrow 3h$	$0.17 \pm 0.03$	$0.73 \pm 0.14$	1
$\tau \rightarrow \mu$	$0.004 \pm 0.001$	$0.61 \pm 0.12$	1
$\tau \rightarrow e$	$0.03 \pm 0.01$	$0.78 \pm 0.16$	0
Total	$0.25 \pm 0.05$	$2.64 \pm 0.53$	5

## Signal Background Modelization

- Multichannel (uncorrelated) **counting model** based on Poisson Statistics
- Gaussian for Background Uncertainties

$$\mathcal{L} = \prod \text{Pois}(n_i, \mu s_i + b_i) \text{Gaus}(b_{0i}, b_i, \sigma_{bi})$$

$\mu$  → strength of the signal (parameter of interest)  
 with  $\mu = 0$  : background-only hypothesis  
 and  $\mu = 1$  : nominal signal+background

test statistics:

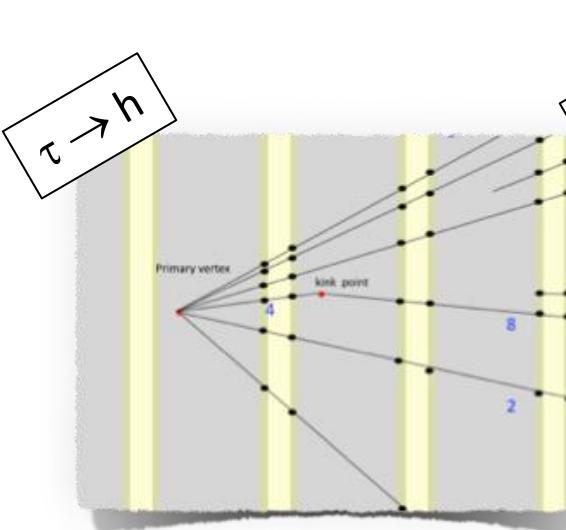
- i) Profile Likelihood Ratio;
- ii) Fisher's rule ( $\mu = 0$ ) .

Background-only hypothesis:

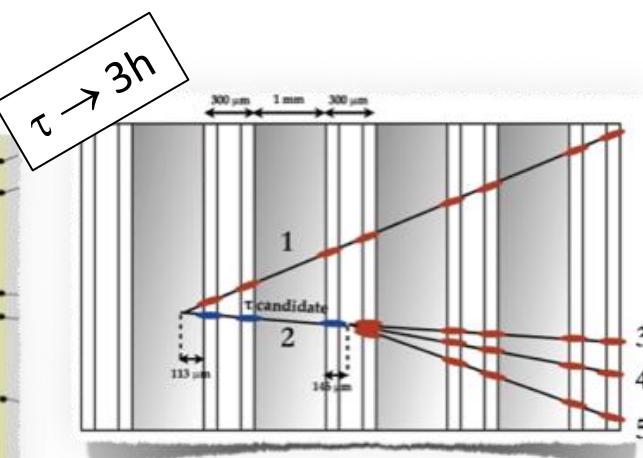
- p-value =  $1.1 \times 10^{-7}$
- excluded at  $5.1\sigma$  significance

Compatibility with 3 $\nu$  oscillation: Probability of less likely data:  
 $\hat{\mu} = 1.8^{+1.8}_{-1.1}$  at 90% C.L.      17% based on total number  
     6.4% if channels considered

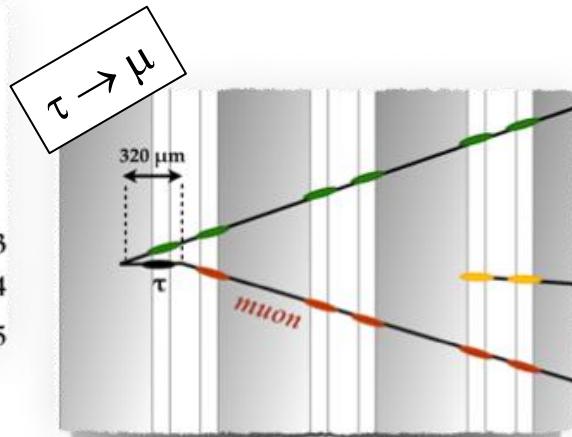
# The five $\nu_\tau$ candidates (2015)



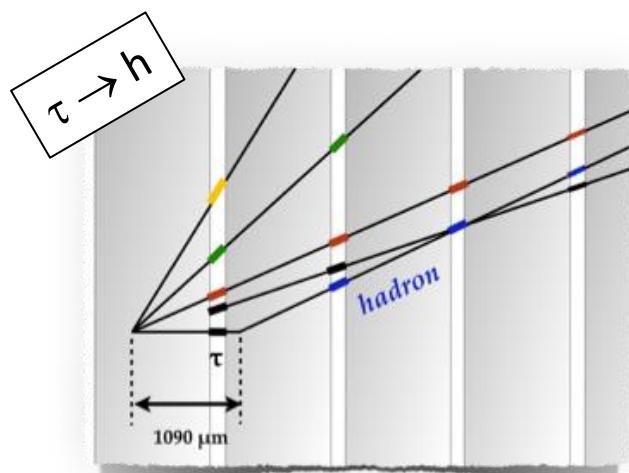
Phys. Lett. B 691 (2010) 138



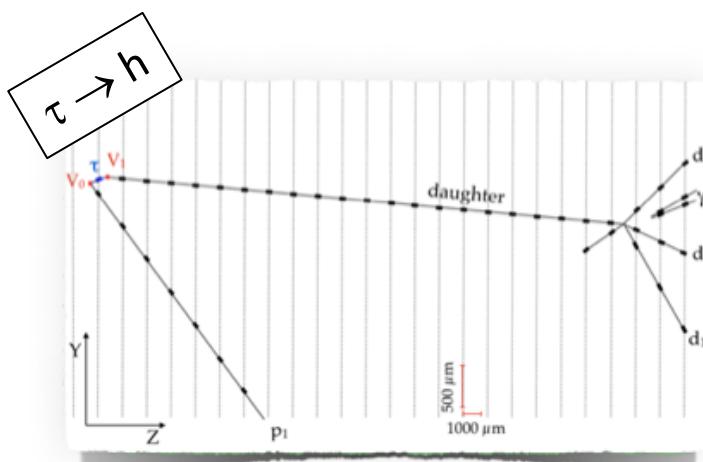
JHEP 11 (2013) 036



Phys. Rev. D 89 (2014) 051102



PTEP 2014 (2014) 10, 101C01



Phys. Rev. Lett. 115 (2015) no.12, 121802

# $\nu_\tau$ appearance: loose event selection (2018)

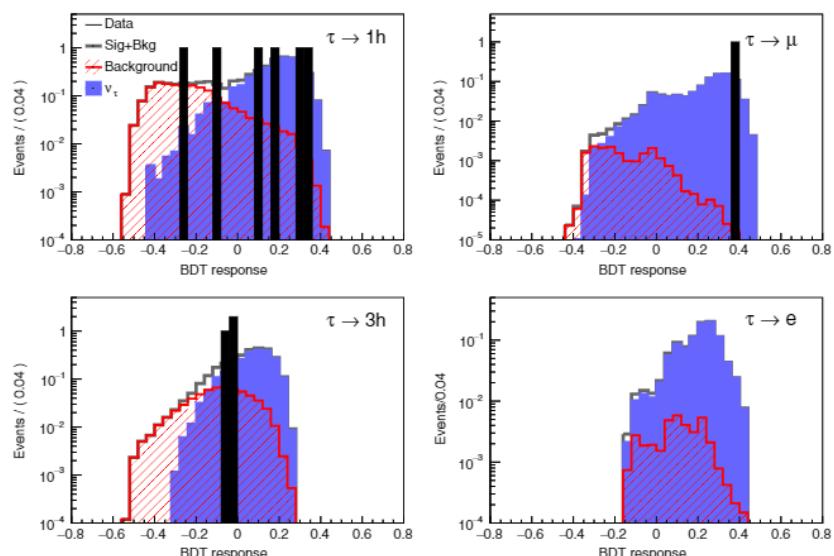
[Phys.Rev.Lett. 120 (2018) no.21, 211801]

- Loose kinematical cuts:
  - **Minimal requirements** to identify the topologies showing 2 vertices
  - **Negligible additional background** from K/ $\pi$  decays

Variable	$\tau \rightarrow 1h$	$\tau \rightarrow 3h$	$\tau \rightarrow \mu$	$\tau \rightarrow e$
$z_{\text{dec}}$ (mm)	<2.6	<2.6	<2.6	<2.6
$\theta_{\text{kink}}$ (rad)	>0.02	>0.02	>0.02	>0.02
$p_{2ry}$ (GeV/c)	>1	>1	[1, 15]	>1
$p_{2ry}^T$ (GeV/c)	>0.15		>0.1	>0.1
Charge <sub>2ry</sub>			Negative or unknown	

	Expected background	$\nu_\tau$ expected	Observed
Total	$2.0 \pm 0.4$	$6.8 \pm 1.4$	10

- **Increment of  $\nu_\tau$  sample: x2**
- **Reduction of S/B from  $\sim 10$  to  $\sim 3$**
- **Multivariate approach** (based on BDT)
  - exploit kinematical, topological information and their correlations  
→ higher discrimination power



⇒ Improvement in  $|\Delta m_{23}^2|$  or alternatively  $\langle \sigma \rangle$  estimation

# Statistical Analysis and Results (2018)

[Phys.Rev.Lett. 120 (2018) no.21, 211801]

- Likelihood:

$$\mathcal{L}(\mu, \beta_c) = \prod_{c=1}^4 \left( \mathcal{P}(n_c | \mu s_c + \beta_c) \prod_{i=1}^{n_c} f_c(x_{ci}) \right) \times \prod_{c=1}^4 \mathcal{G}(b_c | \beta_c, \sigma_{b_c})$$

- where

$$f_c(x_{ci}) = \frac{\mu s_c}{\mu s_c + \beta_c} \text{PDF}_c^{\text{sig}} + \frac{\beta_c}{\mu s_c + \beta_c} \text{PDF}_c^{\text{bkg}}$$

- Test statistic: **profile likelihood ratio**

- Using **asymptotic approximation** [Eur.Phys.J.C71:1554,2011], null hypothesis excluded with  **$6.1\sigma$**  significance

- Best-fit signal strength:

$$\mu = 1.1^{+0.5}_{-0.4}$$

$$\mu \propto |\Delta m_{32}^2|^2 \cdot \langle \sigma \rangle$$

$$|\Delta m_{32}^2| = (2.7^{+0.7}_{-0.6}) \times 10^{-3} \text{ eV}^2$$

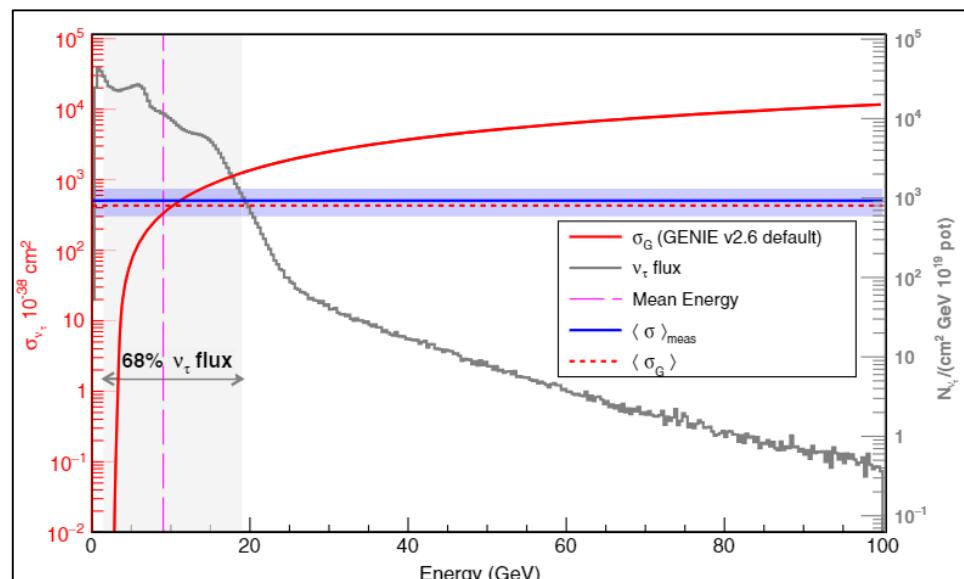
assuming maximal mixing  
first measure in appearance mode

effective tau neutrino cross section

$$\langle \sigma \rangle = (5.1^{+2.4}_{-2.0}) \times 10^{-36} \text{ cm}^2$$

assuming maximal mixing and  $|\Delta m_{32}^2| = 2.5 \times 10^{-3} \text{ eV}^2$

$$\langle \sigma_{\text{Genie}} \rangle = 4.29 \pm 0.04 \times 10^{-36} \text{ cm}^2$$

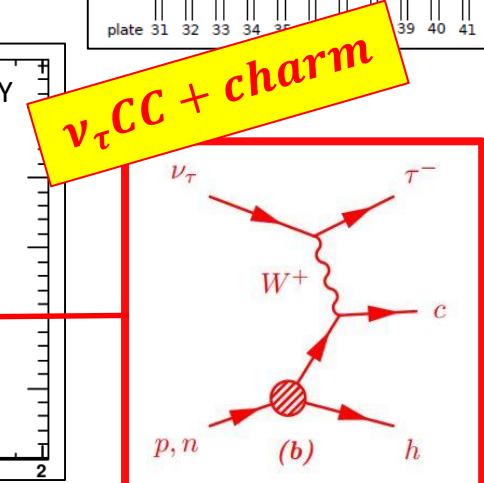
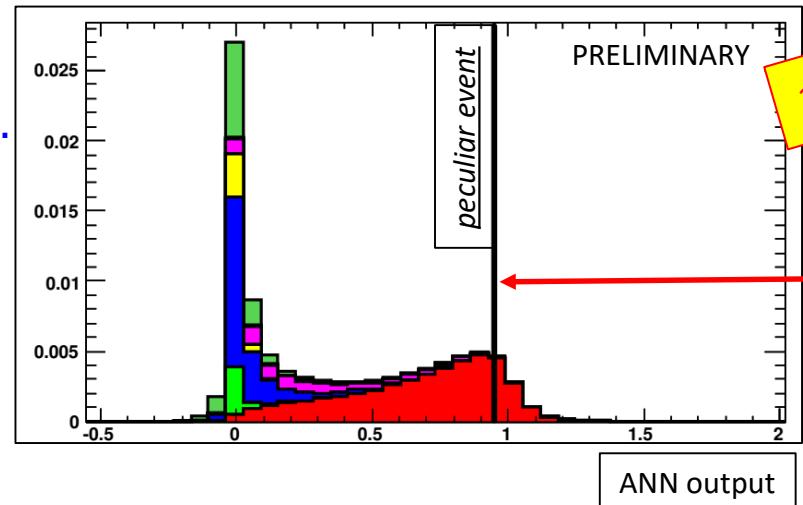
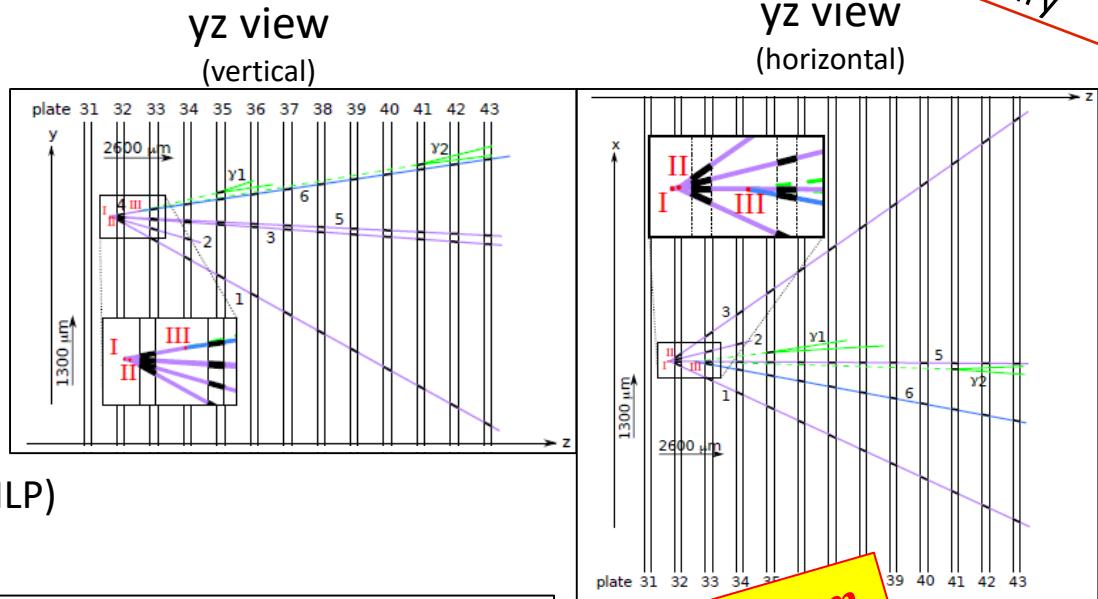


# Peculiar event

Preliminary

- Muon-less neutrino event
- Most probable topology:  
 **$\nu$  interaction vertex + 2 decay vertices**
- **Rare topology** not considered in the experiment proposal  
(0.1 events expected in full data sample)
- **Dedicated simulations + ANN** (2 Layers MLP) to disentangle possible interpretations:

- $\nu_\tau CC + c$
- $\nu_\mu CC + c + \text{had. int.}$
- $\nu_\mu NC + c\bar{c}$
- $\nu_\tau CC + \text{had. int.}$
- $\nu_\mu CC + 2 \text{ had. int.}$
- $\nu_\mu NC + 2 \text{ had. int.}$



The hypothesis the event not being  $\nu_\tau$  CC + charm is excluded:  
**p-value  $\sim 10^{-4} \rightarrow \text{Significance} = 3.4 \sigma$**

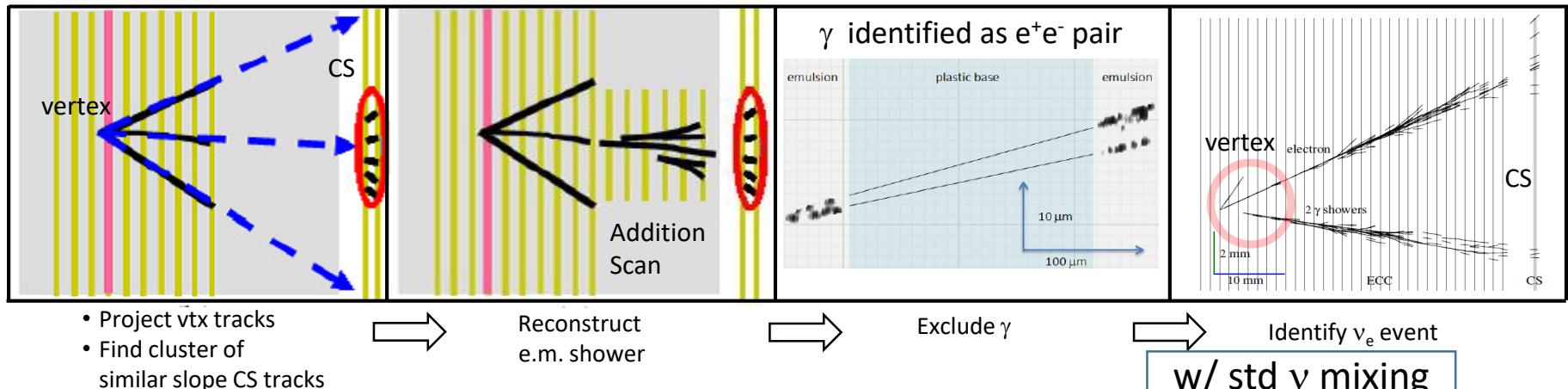
Preliminary

# $\nu_e$ search

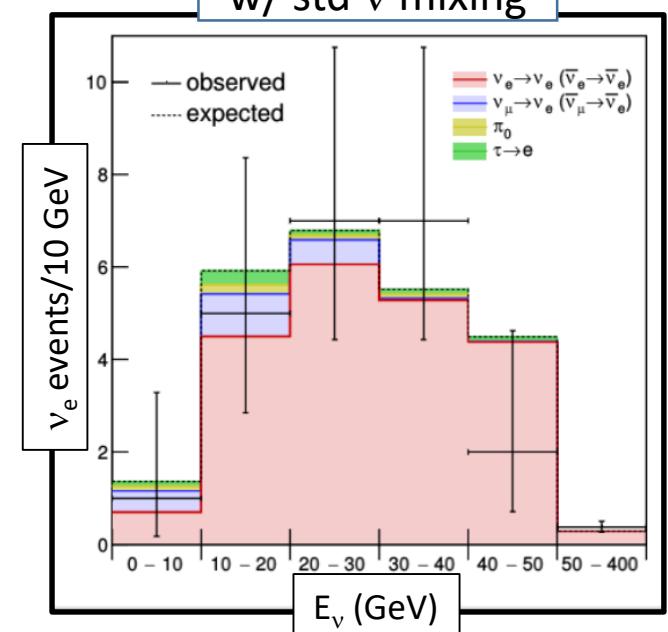
[JHEP 1806 (2018) 151]

[JHEP 1307 (2013) 004]

- OPERA detector granularity allows e.m. shower id →  $\nu_e$  search.
- A **dedicated procedure**, balancing time need vs efficiency.



Component	Expectations w/o $\nu$ mixing	Expectations w/ std $\nu$ mixing
$\nu_e \rightarrow \nu_e$ ( $\bar{\nu}_e \rightarrow \bar{\nu}_e$ )	30.7	31.1
$\tau$ (unidentified) → e	0.7	0.7
$\pi^0 \rightarrow \gamma$ (misidentified)	0.5	0.5
$\nu_\mu \rightarrow \nu_e$ ( $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$ )	0.0	2.0
Total	31.9	34.3
observed	35	



# $\nu_\mu$ disappearance

PRELIMINARY

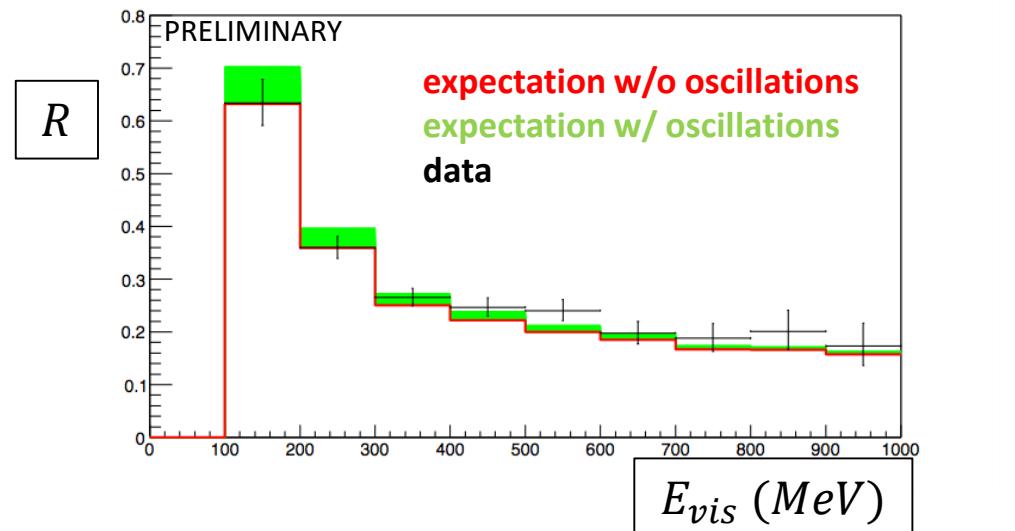
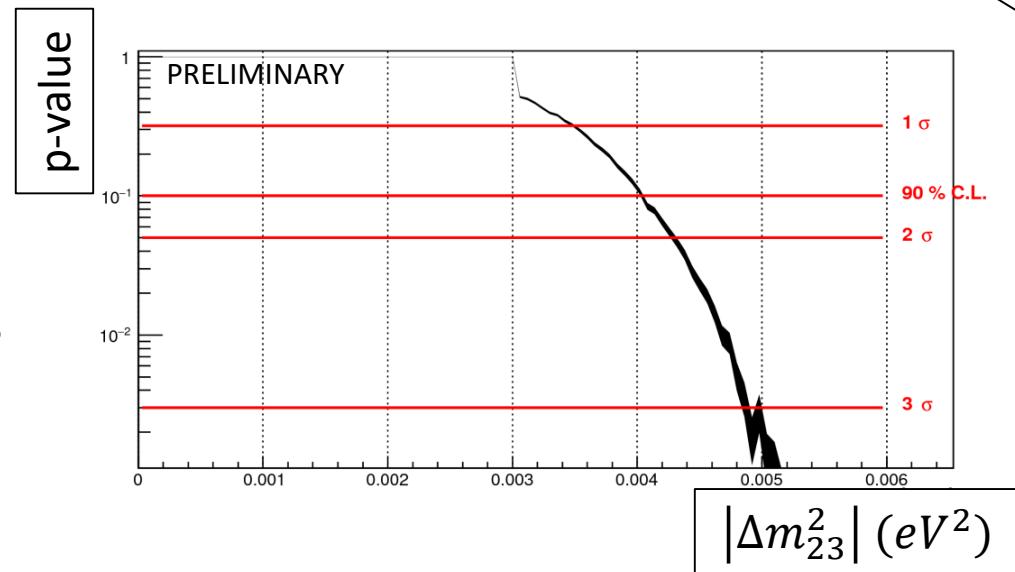
$\nu_\mu$  disappearance sensitivity  
limited by flux uncertainties  
→ no NEAR detector

**Ratio ( $R$ ) of NC-like over CC-like mitigates limitation due to flux uncertainties**

**Electronic detector data:**  
smaller uncertainties w.r.t. emulsion data

Test compatibility with expectation  
for given values of  $|\Delta m_{23}^2|$   
**(assuming maximal mixing)**

$$|\Delta m_{23}^2| < 4.1 \times 10^{-3} \text{ eV}^2 @ 90\% \text{ C.L.}$$



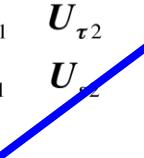
# Sterile neutrino search

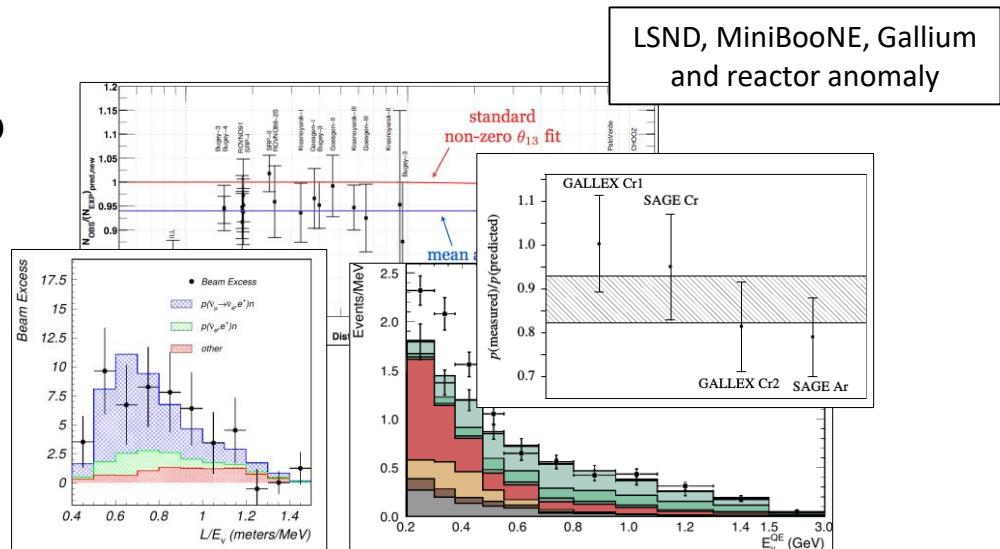
Some experimental results may hint to an additional massive ( $\sim 1 \text{ eV}^2$ ) **sterile** neutrino

Mixing described by  $4 \times 4$  matrix

$$\begin{bmatrix} U_{e1} & U_{e2} & U_{e3} & U_{e4} \\ U_{\mu 1} & U_{\mu 2} & U_{\mu 3} & U_{\mu 4} \\ U_{\tau 1} & U_{\tau 2} & U_{\tau 3} & U_{\tau 4} \\ U_{s1} & U_{s2} & U_{s3} & U_{s4} \end{bmatrix}$$

**$\nu_e$  appearance**  
 **$\nu_\mu$  disappearance**  
 **$\nu_\tau$  appearance**  
**NC disappearance**





**OPERA** can test the sterile neutrino hypothesis looking for deviations from predictions in the electron neutrino appearance or tau neutrino appearance channels.

Predictions of the **3+1 model** evaluated with **GLOBES**

- $\Delta m^2_{21}$  fixed to *PDG* value
- Gaussian constraint on  $\Delta m^2_{31}$  (*PDG* mean and sigma)
- **Matter effects:** constant Earth crust density (PREM onion shell model)  
[Phys. Earth Planet. Interiors 25 (1981) 297]
- $\Delta m^2_{41} > 0$  favored by  $\sum m_\nu$  result from cosmological surveys [A&A 594, A13 (2016)]
- **Profiled likelihood ratio  $\lambda$**  (nuisance parameter profiled out)
- Representation:  $U = R_{34}R_{24}\hat{R}_{23}R_{14}\hat{R}_{13}\hat{R}_{12}$

$\nu_\mu \rightarrow \nu_\tau$  oscillation probability in presence of a sterile neutrino:

$$P(Energy) = C^2 \sin^2 \frac{\Delta_{31}}{2} + \boxed{\sin^2 2\theta_{\mu\tau} \sin^2 \frac{\Delta_{41}}{2}}$$

*pure exotic oscillation*

Effective mixing parameter  
(leading mixing term at SBL)

$+ \frac{1}{2} C \sin 2\theta_{\mu\tau} \cos \phi_{\mu\tau} \boxed{\sin \Delta_{31}} \sin \Delta_{41}$

$- C \sin 2\theta_{\mu\tau} \boxed{\sin \phi_{\mu\tau}} \sin^2 \frac{\Delta_{31}}{2} \sin \Delta_{41}$

$+ 2 C \sin 2\theta_{\mu\tau} \cos \phi_{\mu\tau} \sin^2 \frac{\Delta_{31}}{2} \sin^2 \frac{\Delta_{41}}{2}$

$+ C \sin 2\theta_{\mu\tau} \boxed{\sin \phi_{\mu\tau}} \boxed{\sin \Delta_{31}} \sin^2 \frac{\Delta_{41}}{2}$

$+ \Delta m_{21}^2 \text{ terms ...}$

*Mass Hierarchy dependence*

*interference terms*

*CP-violating terms*

*Effective parameters*

$$C = 2|U_{\mu 3}| |U_{\tau 3}|$$

$$\phi_{\mu\tau} = \text{Arg}(U_{\mu 3} U_{\tau 3}^* U_{\mu 4}^* U_{\tau 4})$$

$$\sin^2 2\theta_{\mu\tau} = 2|U_{\mu 4}| |U_{\tau 4}|$$

... with  $\nu_\tau$

PRELIMINARY

Update of  
[JHEP 1506 (2015) 069]

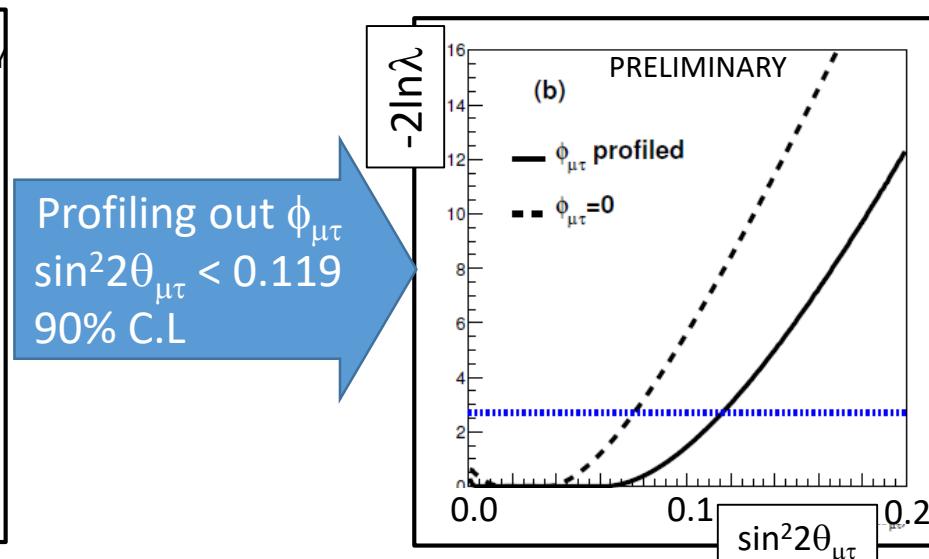
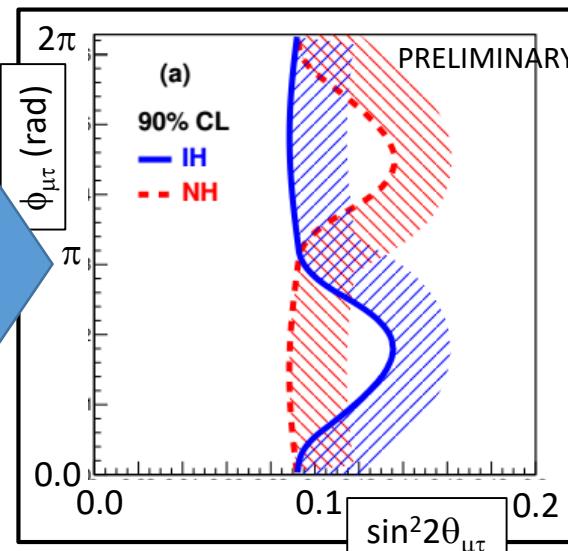
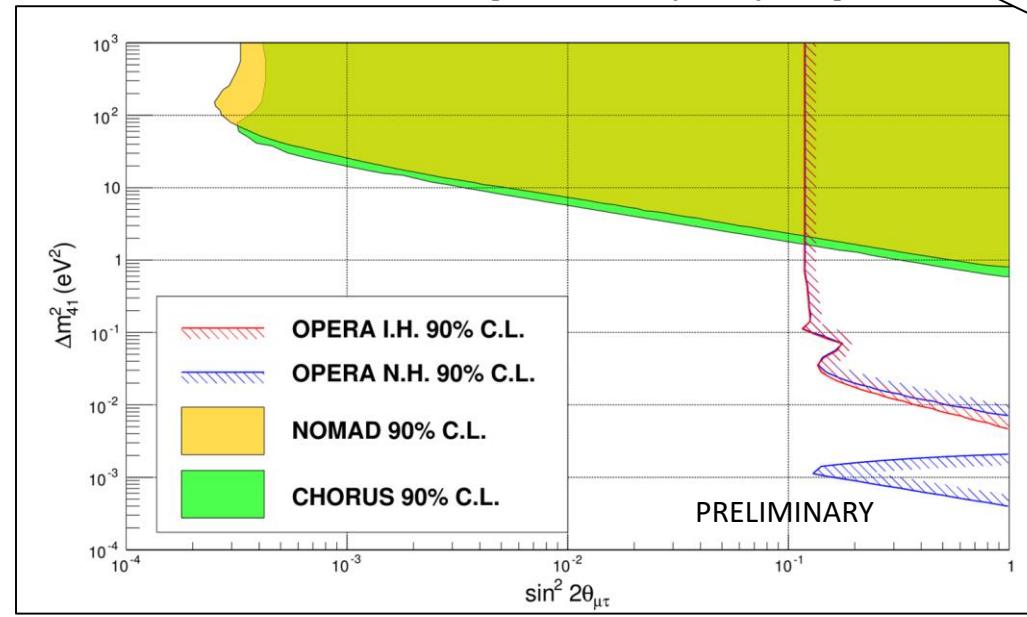
- Counting analysis

$$L = \text{Pois}(n; \mu) \times \text{Gaus}(\widehat{\Delta m_{23}^2}; \Delta m_{23}^2, \widehat{\sigma_{\Delta m}})$$

$\mu$ : expectation (GLoBES)  
 $n$ : observation (data)

$\widehat{\Delta m_{23}^2}, \widehat{\sigma_{\Delta m}}$  PDG values

- Both **normal** and **inverted** neutrino mass hierarchies considered
- Exclusion region on  $\Delta m_{41}^2$  vs  $\sin^2 2\theta_{\mu\tau}$  plane
- Energy selection** ( $E_\nu < 30$  GeV) maximizes sensitivity



- $\nu_e$  energy distribution  
to evaluate exclusion region on:

$$\Delta m_{41}^2 \text{ vs } \sin^2 2\theta_{\mu e}$$

$$\text{where } \sin^2 2\theta_{\mu e} = 4 |U_{\mu 4}|^2 |U_{e 4}|^2$$

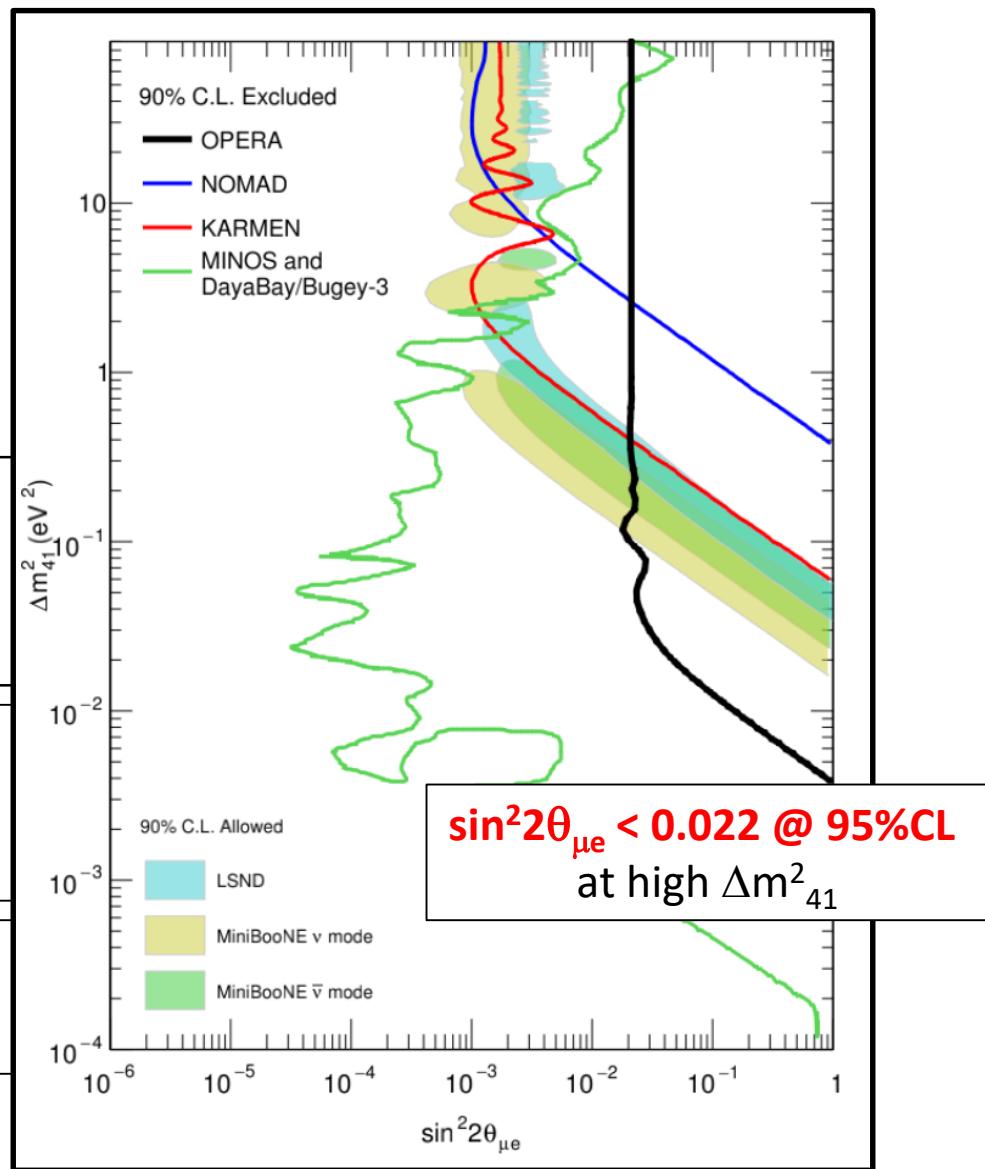
Systematics errors  $\sigma_i$  due to:

- Beam and efficiencies uncertainties
- 20%  $E_\nu < 10$  GeV & 10%  $E_\nu > 10$  GeV
- Bin-to-bin uncorrelated (conservative approach)

$$L = \left( \prod_i \text{Pois}(n_i; \mu_i(1 + k_i)) \times \text{Gaus}(0; k_i, \sigma_i) \right) \times \text{Gaus}(\widehat{\Delta m_{23}^2}; \Delta m_{23}^2, \widehat{\sigma_{\Delta m}})$$

Likelihood

Constraints on  $\Delta m_{23}^2$   
 $\widehat{\Delta m_{23}^2}, \widehat{\sigma_{\Delta m}}$  from PDG



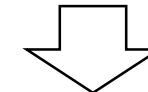
# Combining $\nu_\tau$ and $\nu_e$

PRELIMINARY

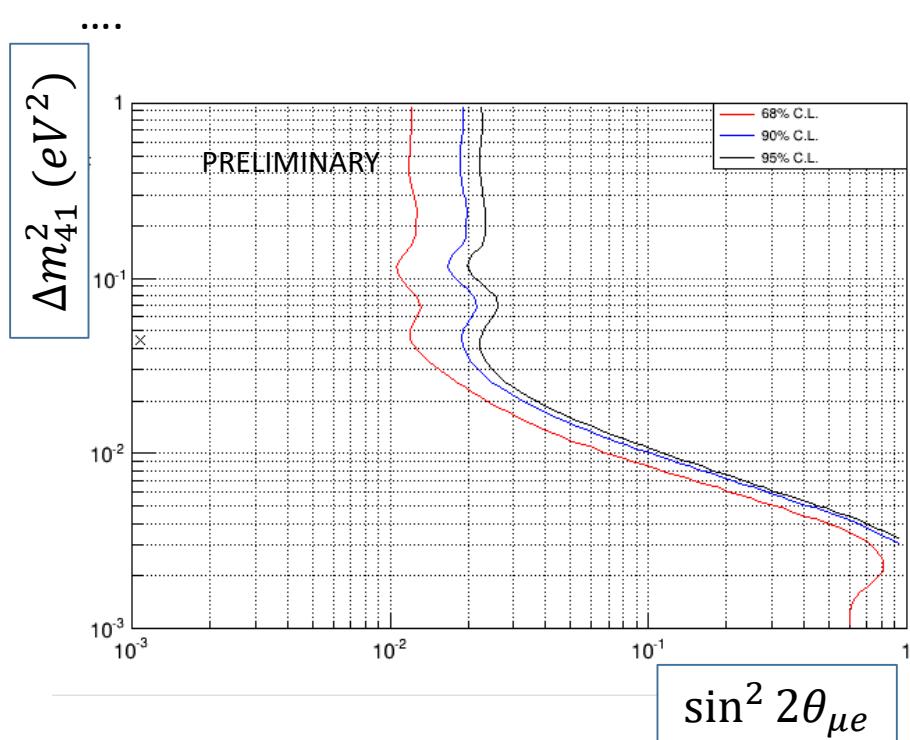
Exploiting simultaneously results of

- $\nu_\tau$  search: 10 candidates
- $\nu_e$  search: 35 candidates

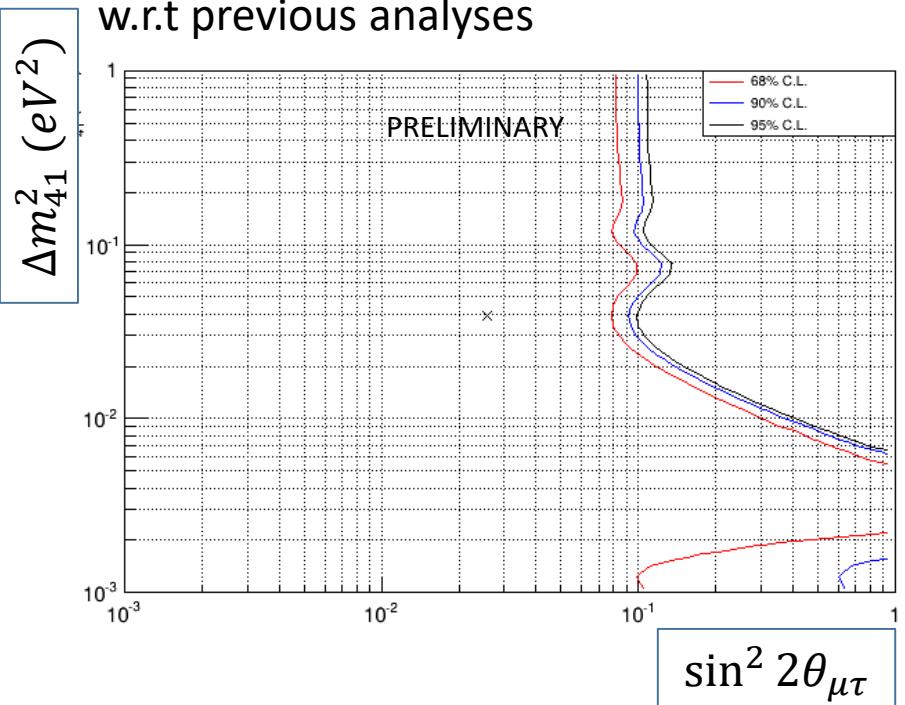
... to extract limits on the parameters of  
the 3 + 1 neutrino model



(Small) exclusion power enhancement  
w.r.t previous analyses



$\sin^2 2\theta_{\mu e} < 0.019$  [90% C.L.]  
@  $\Delta m_{41}^2 \sim 1$  eV<sup>2</sup>



$\sin^2 2\theta_{\mu \tau} < 0.099$  [90% C.L.]  
@  $\Delta m_{41}^2 \sim 1$  eV<sup>2</sup>

# Annual $\mu$ rate modulation

**PRELIMINARY**

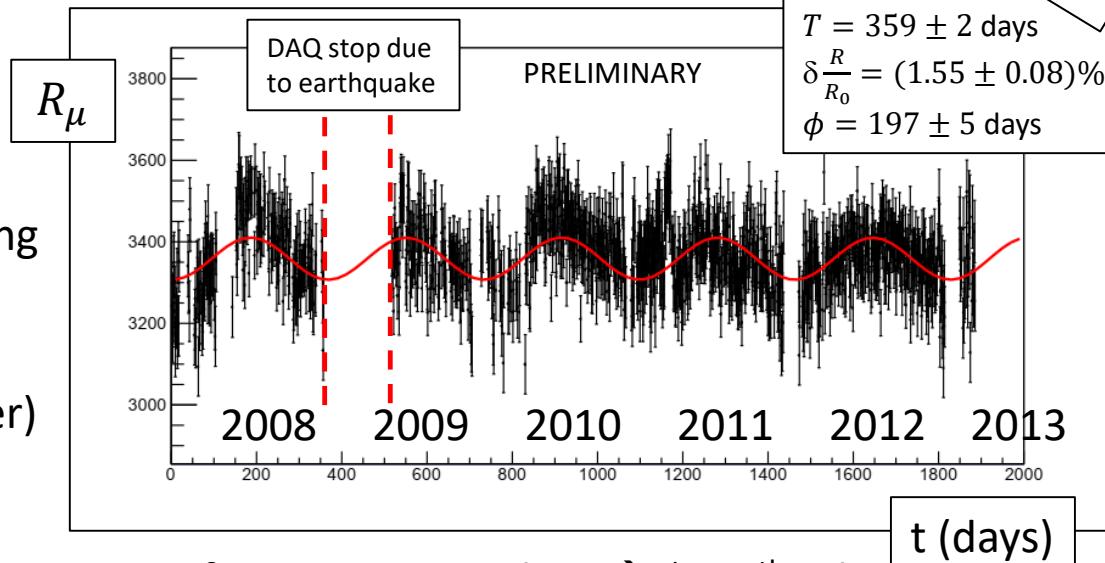
$\Delta T$  in the upper atmosphere

- ⇒ variation in atm. density
- ⇒ variation in  $\pi$  interaction length
- ⇒ variation in the fraction of mesons decaying before interacting

**Annual modulation of  $\mu$  rate ( $R_\mu$ )**

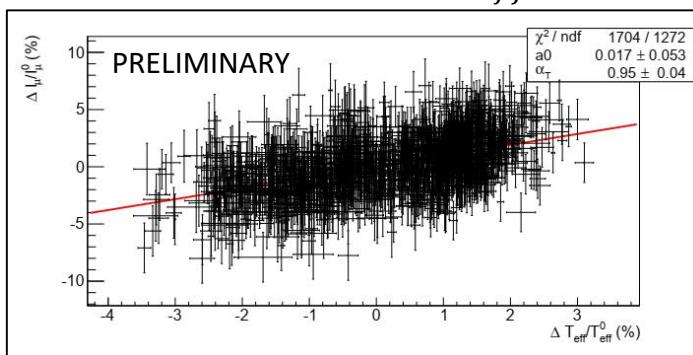
(More muons in summer than in winter)

$$\text{Fit with: } R_\mu = R_0 + \delta R \cos \frac{2\pi}{T} (t - \phi)$$



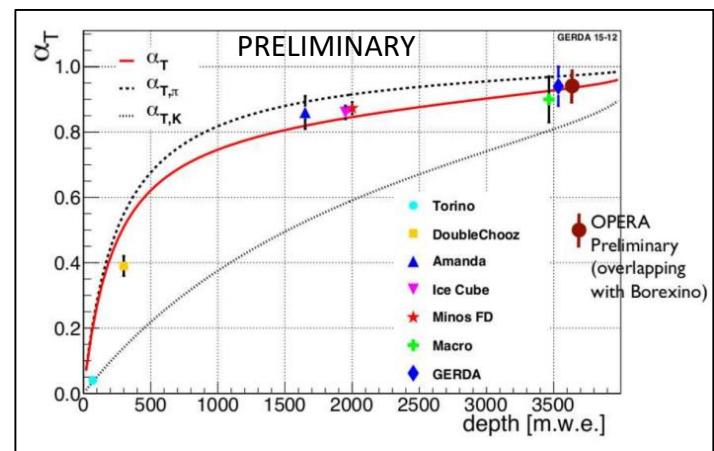
If  $T$  is set to 365 days  $\rightarrow \phi = 5^{\text{th}} \text{ July}$

**Correlation of  $R_\mu$  and the effective temperature ( $T_{eff}$ )**



$$\alpha_T = \frac{\Delta R_\mu}{\Delta T_{eff}} = 0.95 \pm 0.04$$

**$\alpha_T$  VS depth**

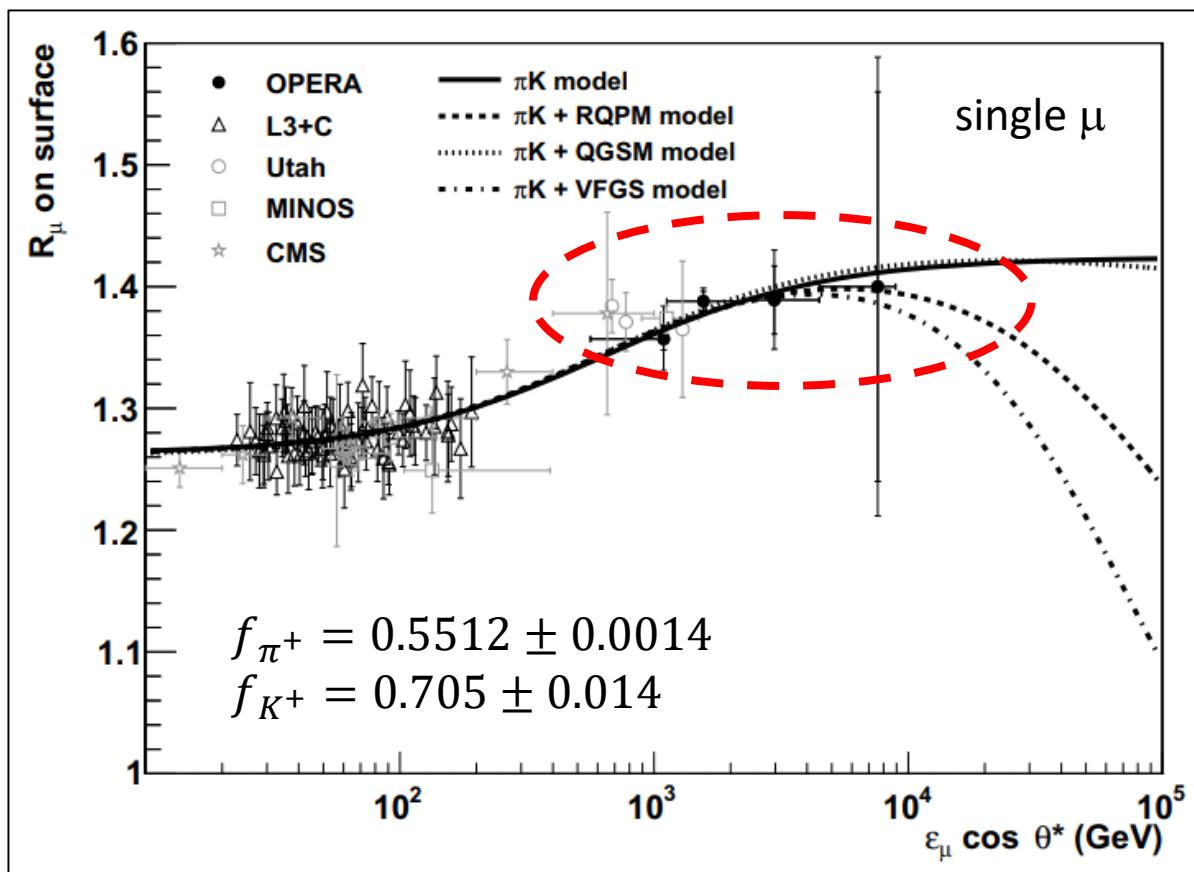


# Atmospheric muon charge ratio

[Eur. Phys. J. C (2014) 74]

- **Highest-E region** reached
- **Opposite magnet polarities** runs  
→ lower **systematics**
- Strong reduction of the charge ratio for multiple muon events
  - single- $\mu$   $1.377 \pm 0.006$
  - multi- $\mu$   $1.098 \pm 0.023$
- Results compatible with a simple  **$\pi$ -K model**
- **No** significant contribution of the **prompt component** up to  $E_\mu \cos \theta^* \sim 10$  TeV
- Validity of **Feynman scaling** in the fragmentation region up to  $E_\mu \sim 20$  TeV ( $E_N \sim 200$  TeV)

$$\phi_{\mu^\pm} \propto \frac{a_\pi f_{\pi^\pm}}{1 + b_\pi \mathcal{E}_\mu \cos \theta / \epsilon_\pi} + R_{K\pi} \frac{a_K f_{K^\pm}}{1 + b_K \mathcal{E}_\mu \cos \theta / \epsilon_K}$$

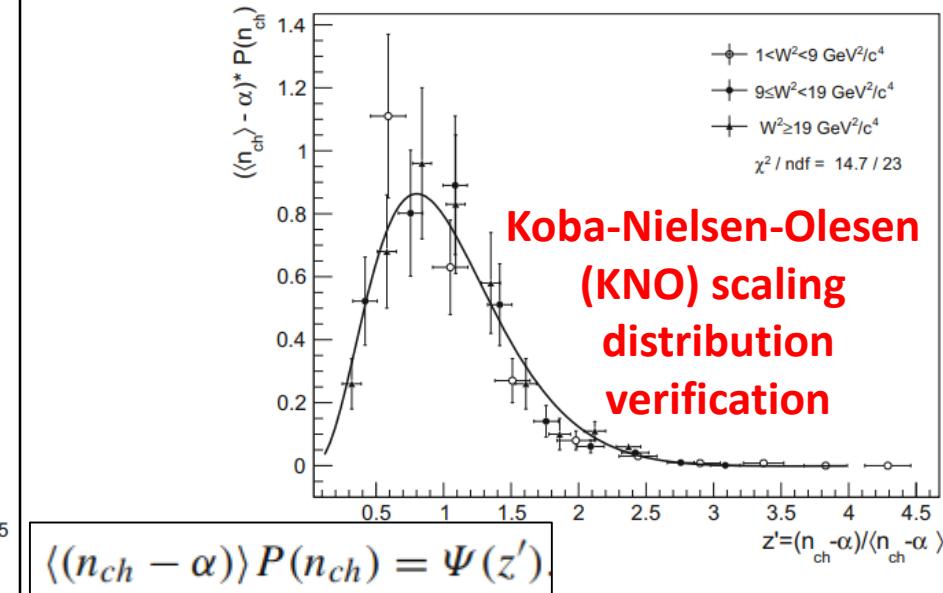
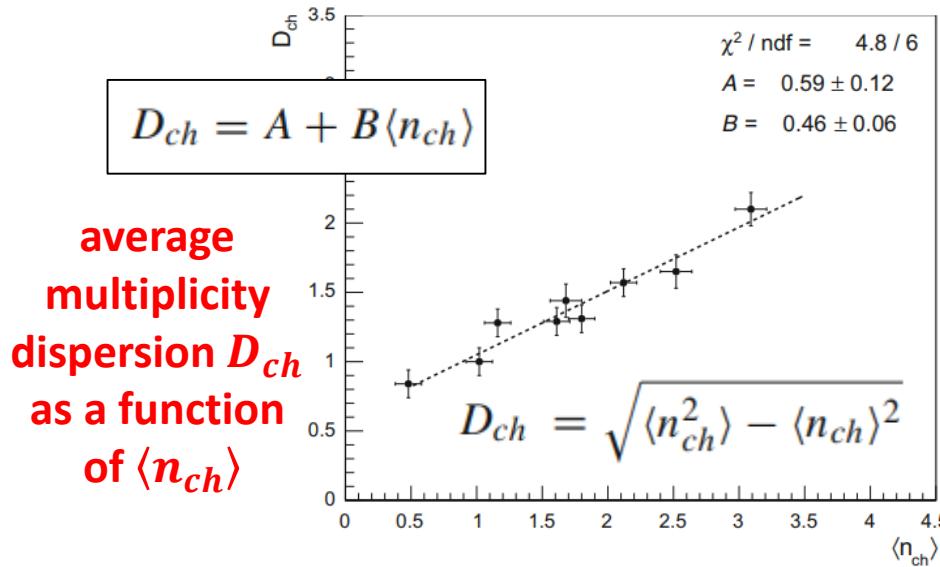
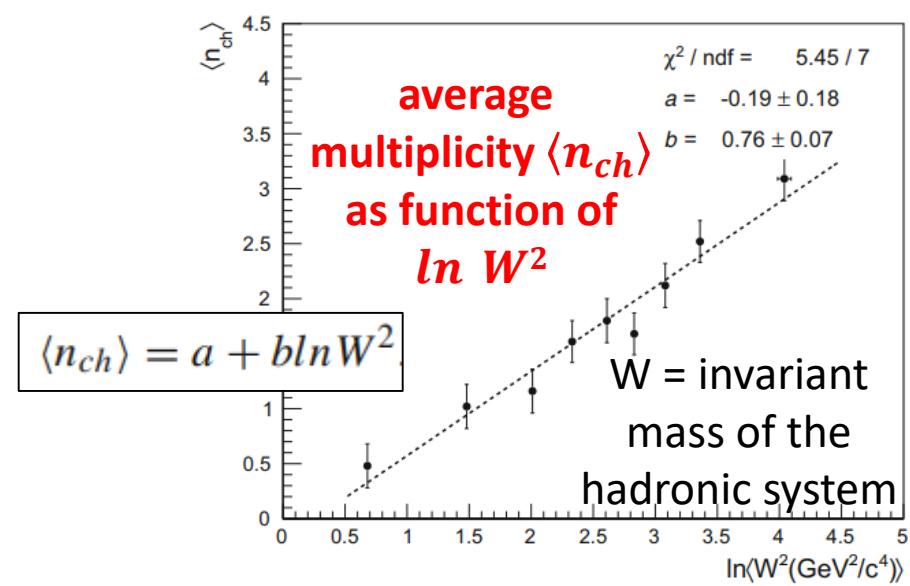
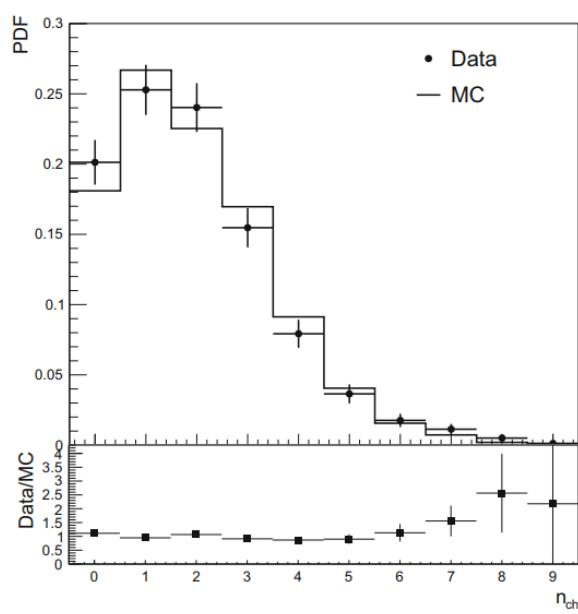


# Neutrino interactions multiplicity

unbiased sample of  
 $\nu_\mu CC$  interactions

[Eur.Phys.J. C78 (2018) no.1, 62]

**charged hadron  
multiplicity  
distribution**



# Summary

- Discovery of  $\nu_\mu \rightarrow \nu_\tau$  appearance in the CNGS neutrino beam:  $5.1\sigma$
- Loose selection analysis increase discovery significance  $6.1\sigma$ 
  - Measurement of  $\Delta m^2_{23}$  (first measurement in appearance mode)
  - Measurement of effective  $\nu_\tau$  cross-section
- Muon-less double decay event has been reported.  
Favored interpretation  $\nu_\tau$  CC interaction with charm production
- Final results from  $\nu_\mu \rightarrow \nu_e$  oscillation search
- Search for  $\nu_\mu$  disappearance
  - Upper limit on  $\Delta m^2_{23}$
- Constraints on sterile neutrinos  
from  $\nu_\mu \rightarrow \nu_e$ ,  $\nu_\mu \rightarrow \nu_\tau$  and their combination  
in the 3+1 flavor model
- Non-oscillation Physics:
  - atmospheric muons charge ratio
  - annual modulation of atmospheric muons rate
  - Neutrino interactions charged multiplicity study

# OPERA taking a "selfie"... Thank you!



Image taken using **OPERA nuclear emulsion film**  
with a pinhole hand made camera  
courtesy by Donato Di Ferdinando